

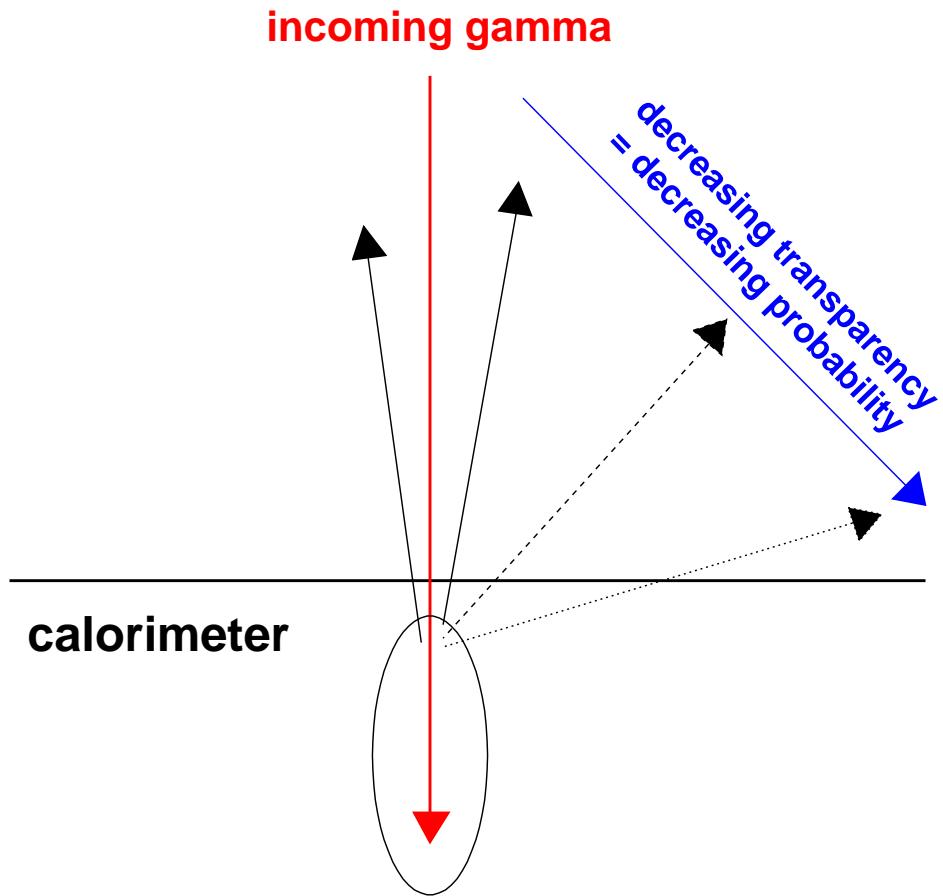
Backsplash study with the ACD

(thanks to Eric Charles and Alex Moiseev)

- backsplash and other splashes
- ACD tiles signal
- backsplash measurement
- effect of cracks
- effect of incoming angle
- how can we test backsplash at SPS ?

Pure backsplash

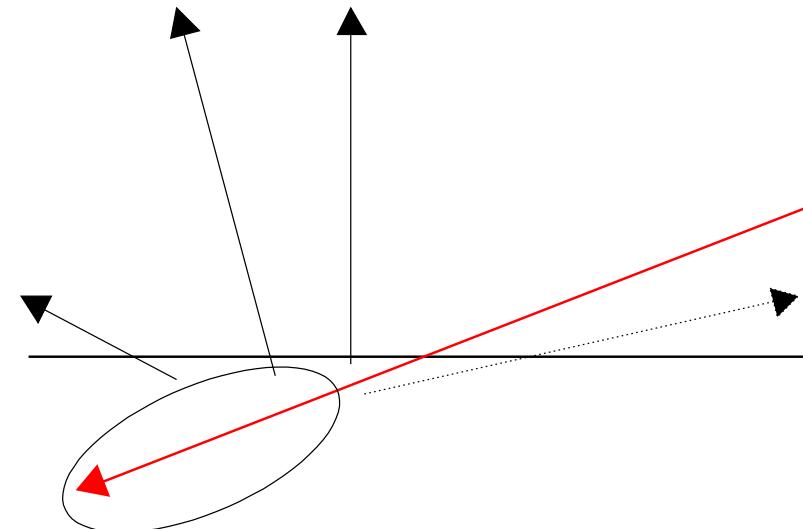
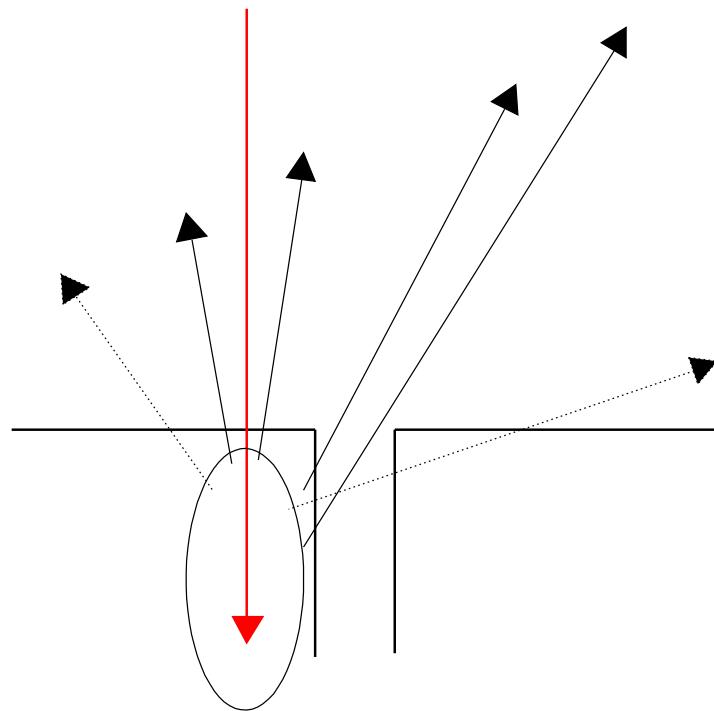
- high energy incoming gamma (> 10 GeV)
- some low energy outgoing photons (hundreds of keV)



- backsplash should be maximal at 180 deg.
(\Rightarrow back)
- and should decrease to zero at 90 degrees

Crack-splash and side-splash

- particles of the shower escaping through cracks
- when no more on-axis, the maximum of backsplash is no more at 180 degrees
- no more only photons ...



Simulation

Simulation of high energy gammas

- in the entire LAT
- using v7r3
- pencil beam with radius = 1cm
- 10, 50, 100 and 200 GeV
- various impact positions and incoming angles

Selection

- at least one track
- conversion point inside the tracker (= not inside ACD)

Backsplash in ACD

- on the top : 5×5 tiles (face 0)
- on each side (face 1,2,3,4) :
 - 3 rows of 5 tiles
 - + one large bottom tile

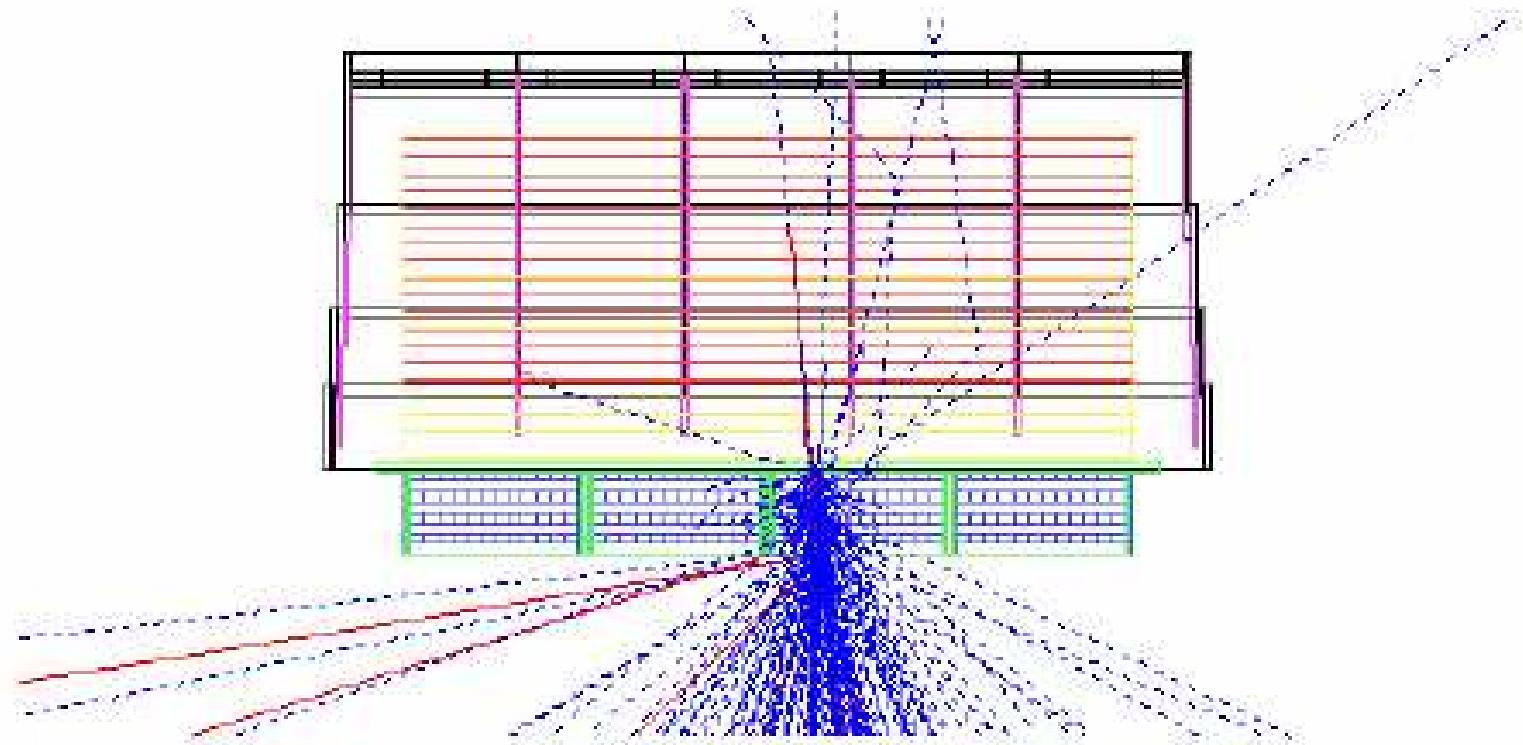
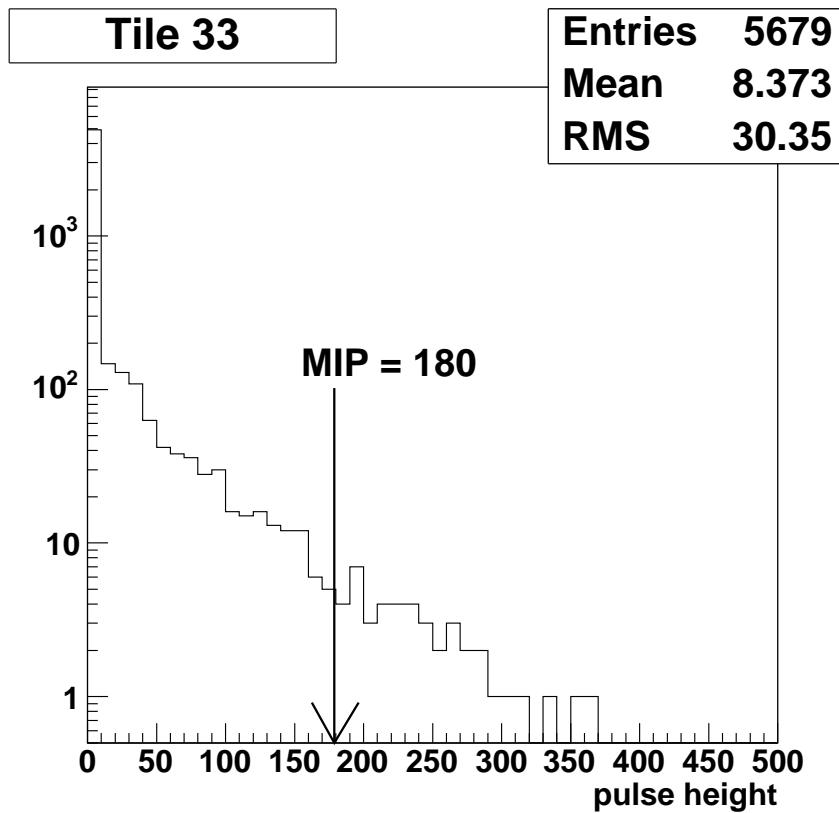


figure from Alex et al talk (IA workshop, July 14, 2005)

ACD tiles signal

usual backsplash pulse height distribution
in tile above impact point when on-axis

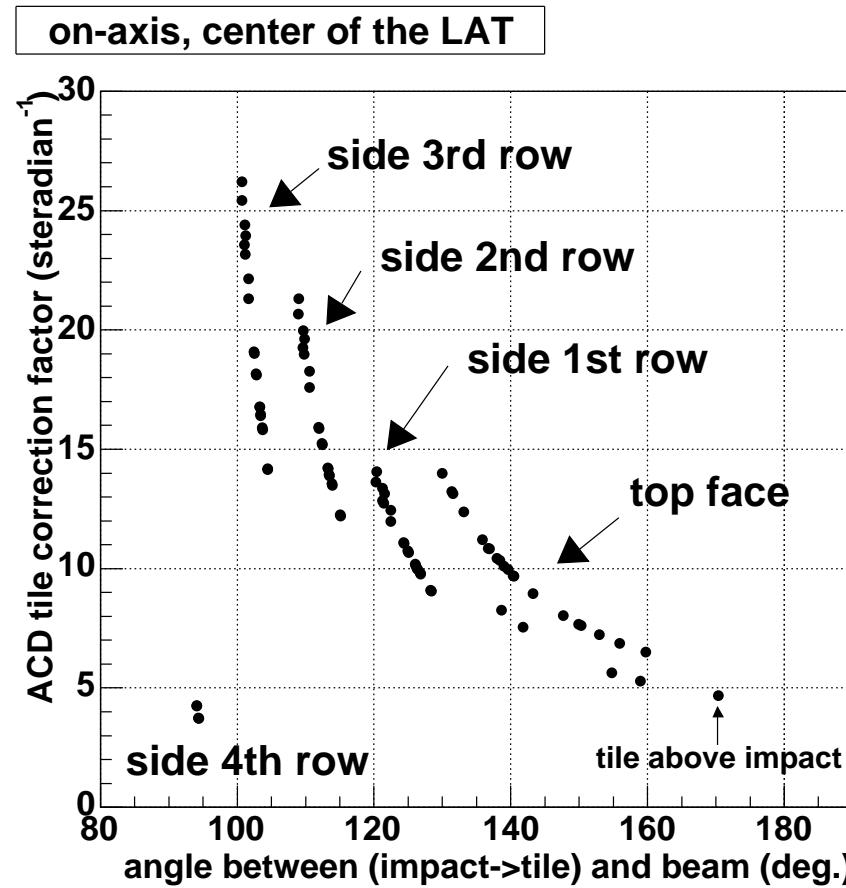


- two estimators :
- pulse height average
 - fraction of events above a certain fraction of MIP

Geometrical correction

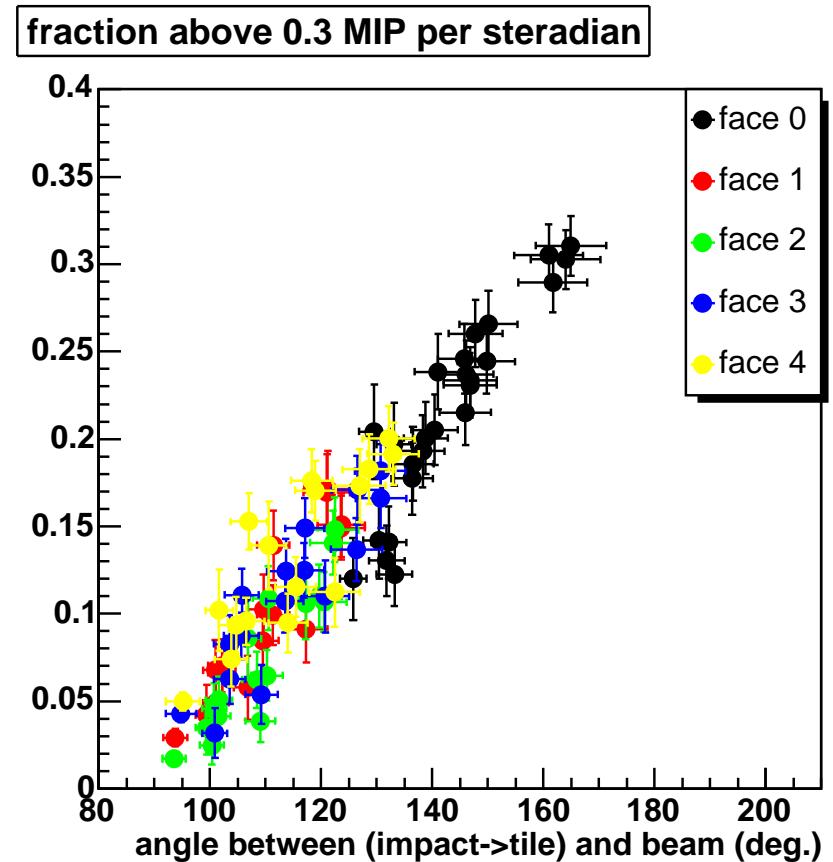
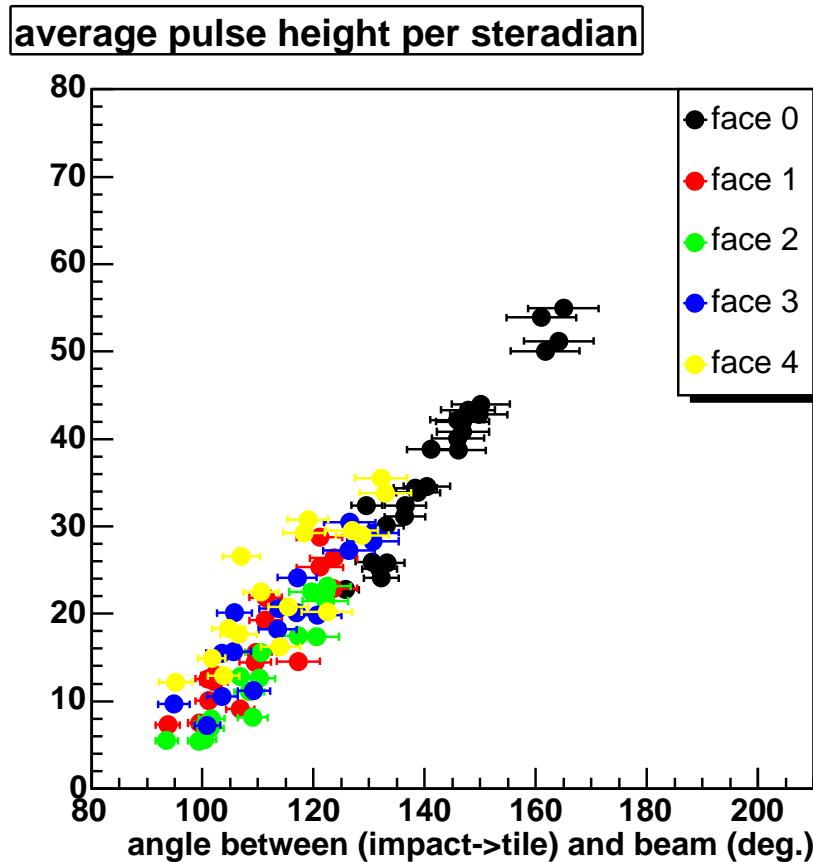
In order to compare signals of different tiles :

⇒ signal $\times 1/(\text{solid angle}) \times 10\text{mm}/(\text{mean path length})$



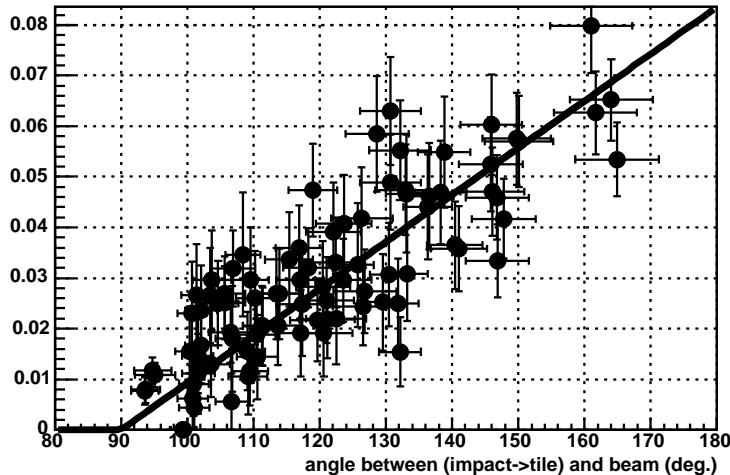
Backsplash angular distribution

- the correction seems to work well
- from 0 at 90 degrees to maximum at 180 degrees

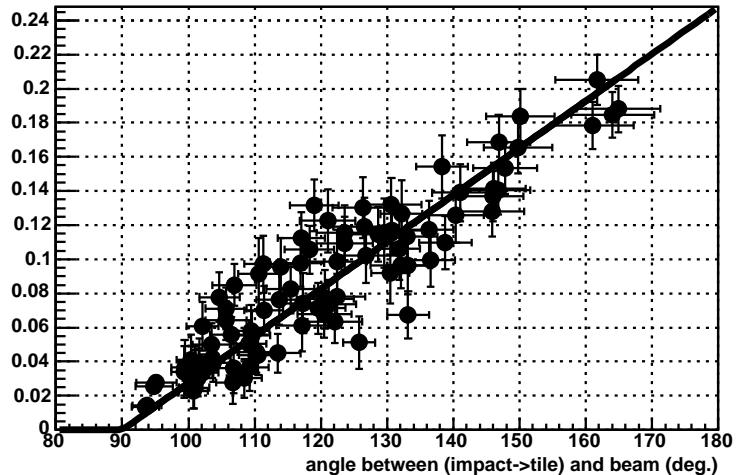


Energy dependence

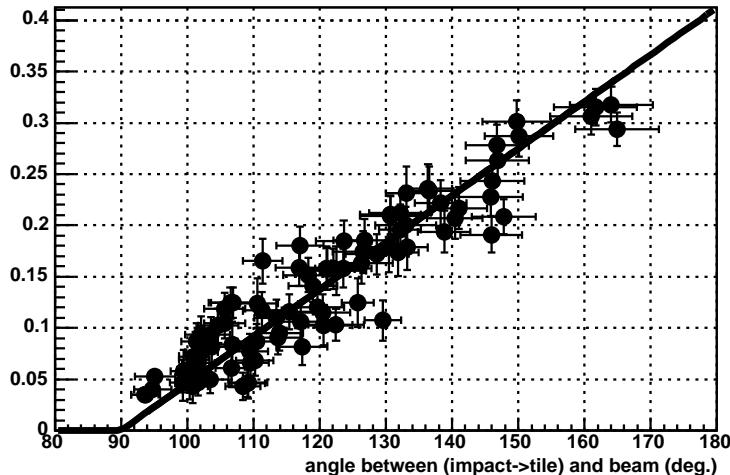
10 GeV : fraction above 0.3 MIP per steradian



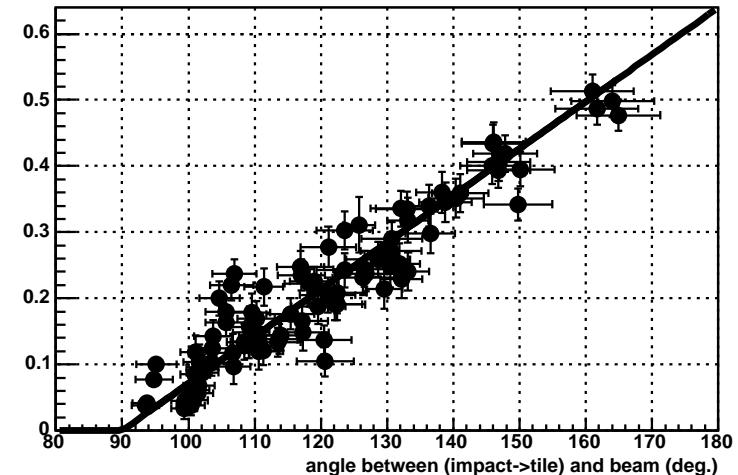
50 GeV : fraction above 0.3 MIP per steradian



100 GeV : fraction above 0.3 MIP per steradian



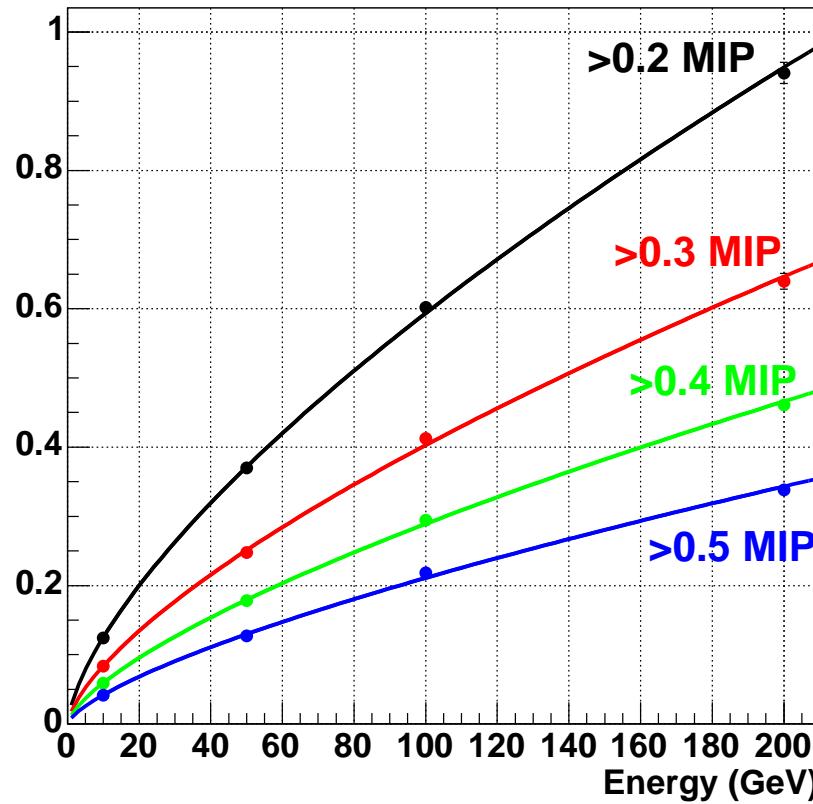
200 GeV : fraction above 0.3 MIP per steradian



Energy parameterization

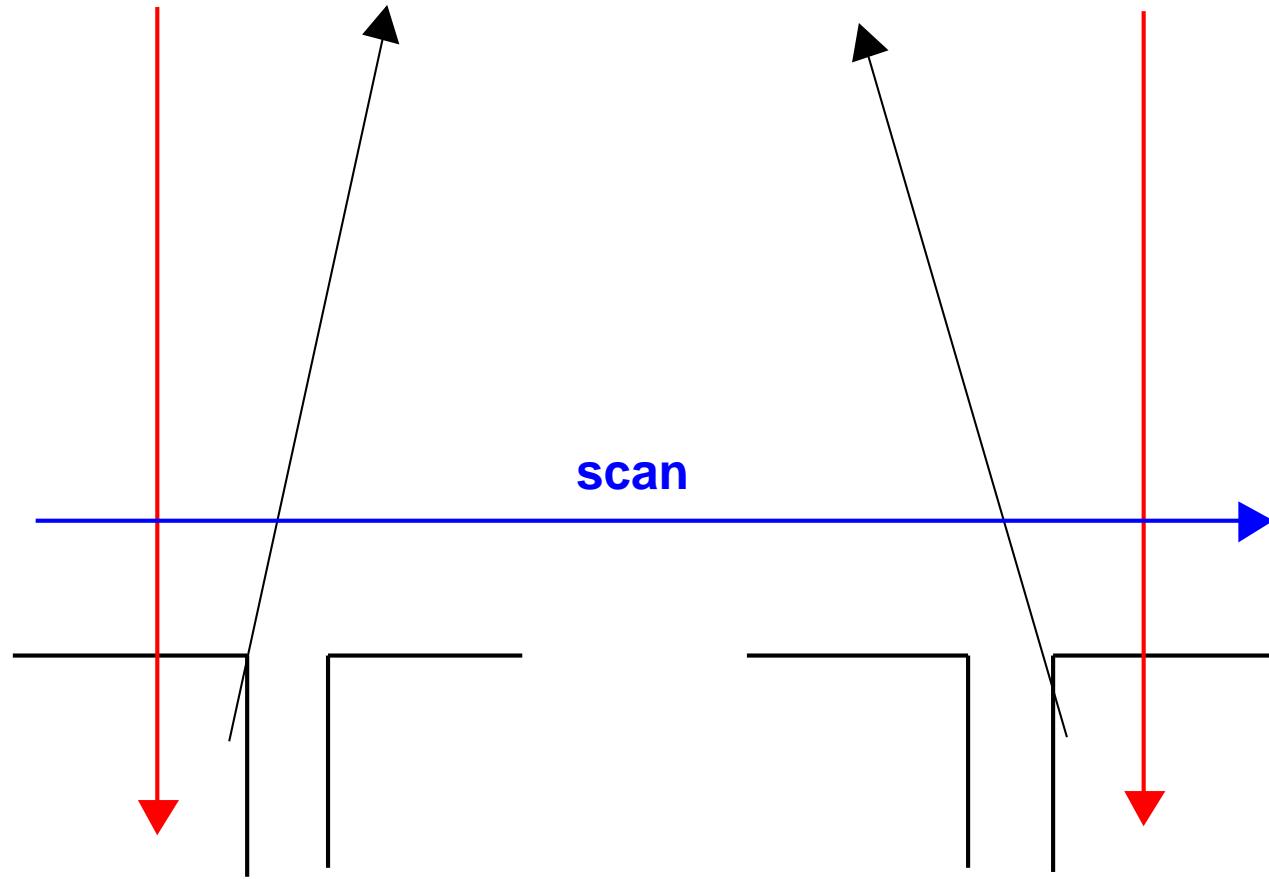
fraction per steradian at 180 deg. $\propto E^{0.7}$

fraction per steradian at 180 deg.



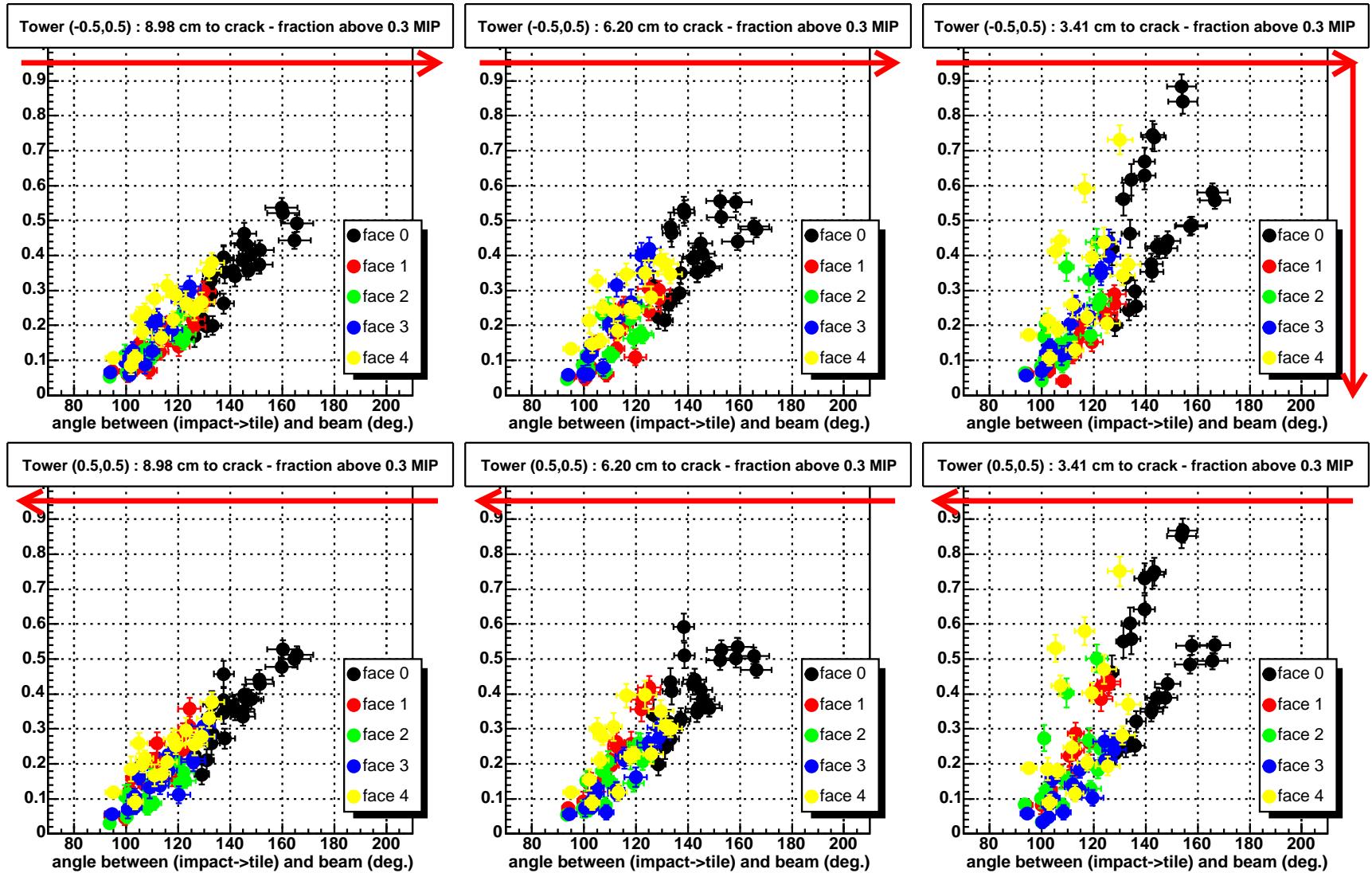
Cracksplash

the signal should be enhanced in the tiles
on the opposite side of the crack



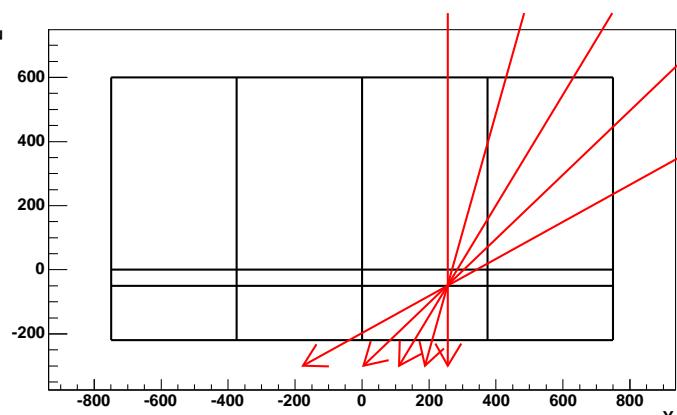
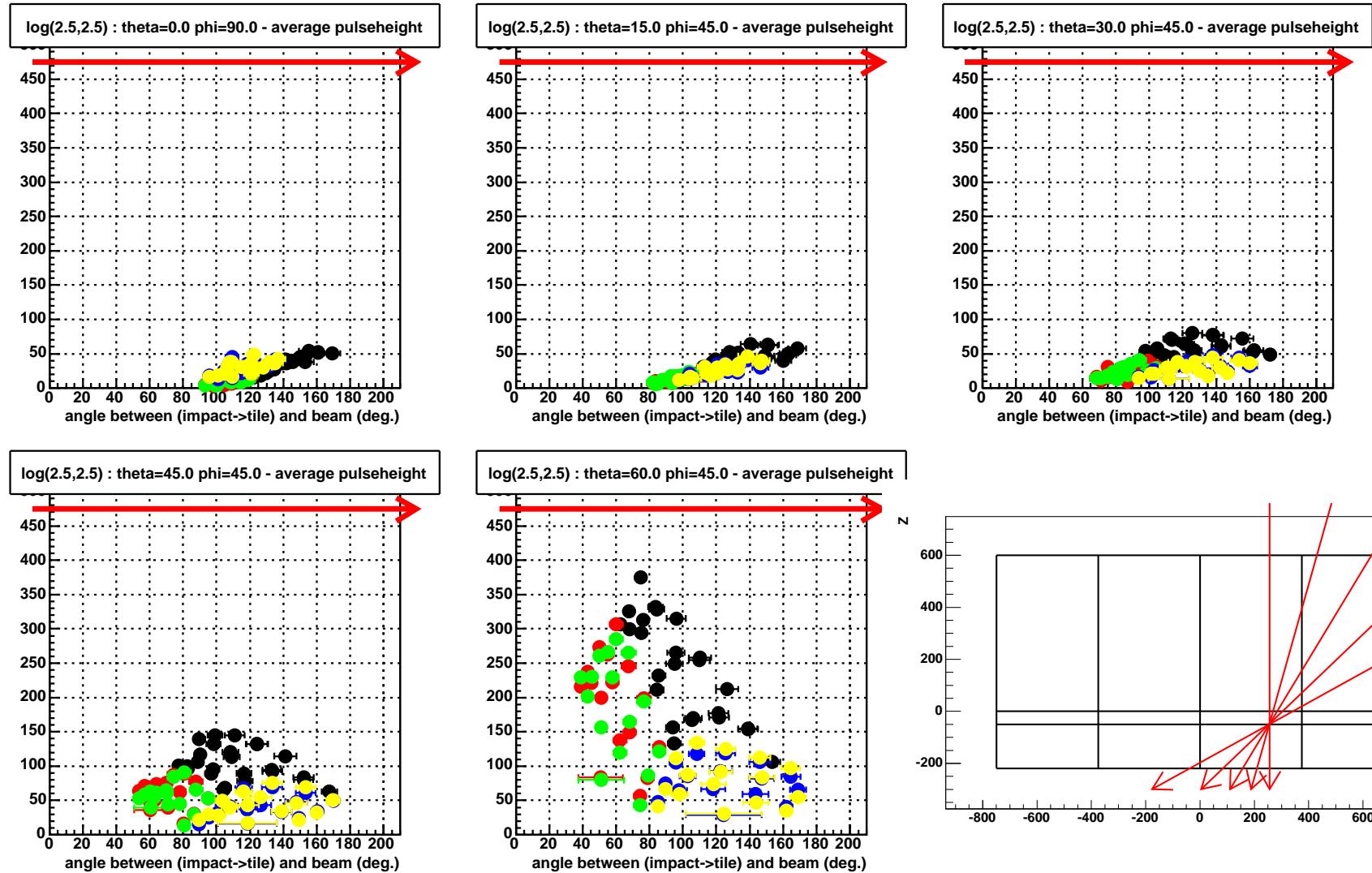
Scan close to crack (200 GeV on-axis)

distance to crack : -9cm → -6 → -3.5 → +3.5 → +6 → +9cm



From on-axis to 60 deg (100 GeV)

$\theta = 0 \rightarrow 15 \rightarrow 30 \rightarrow 45 \rightarrow 60$ degrees



What should be checked at SPS ?

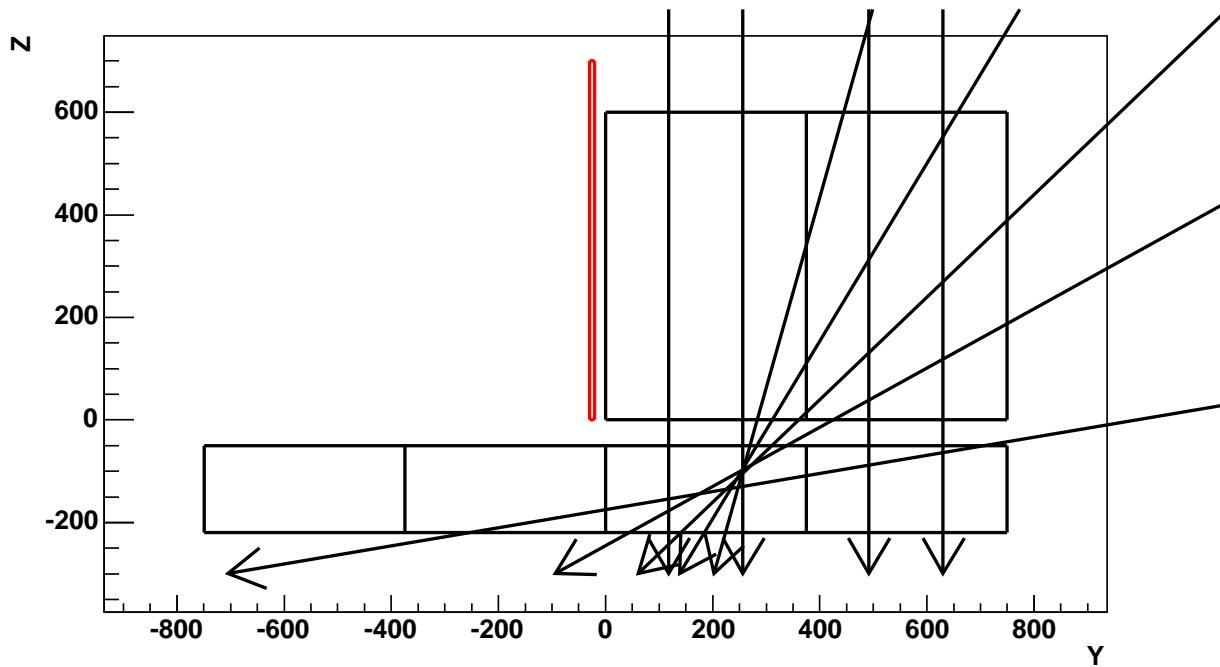
We should check how well the simulation reproduces :

- pure backsplash (intensity, angular distribution, energy dependence)
- the effect of cracks
- the large intensity increase for large incoming angles

Where should we put the ACD tiles ?

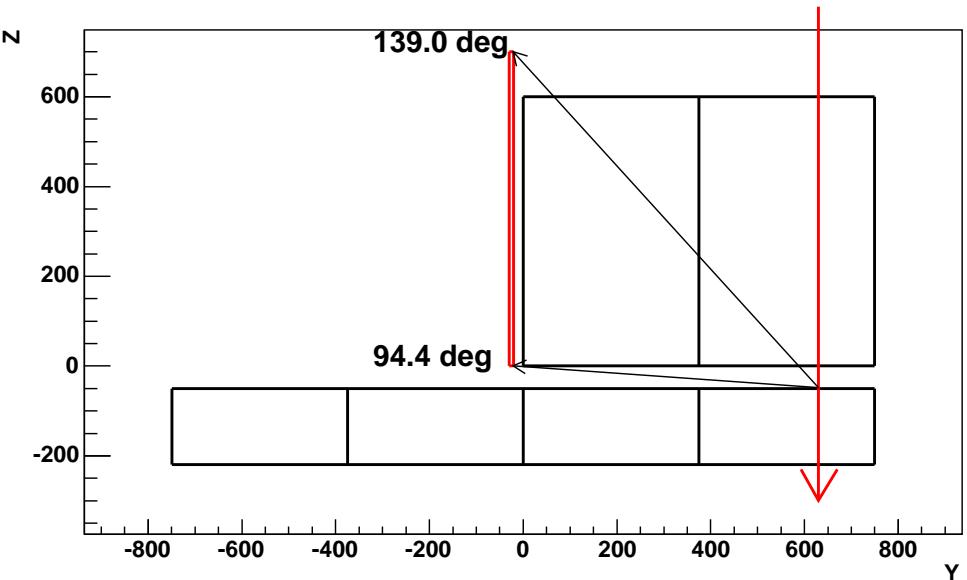
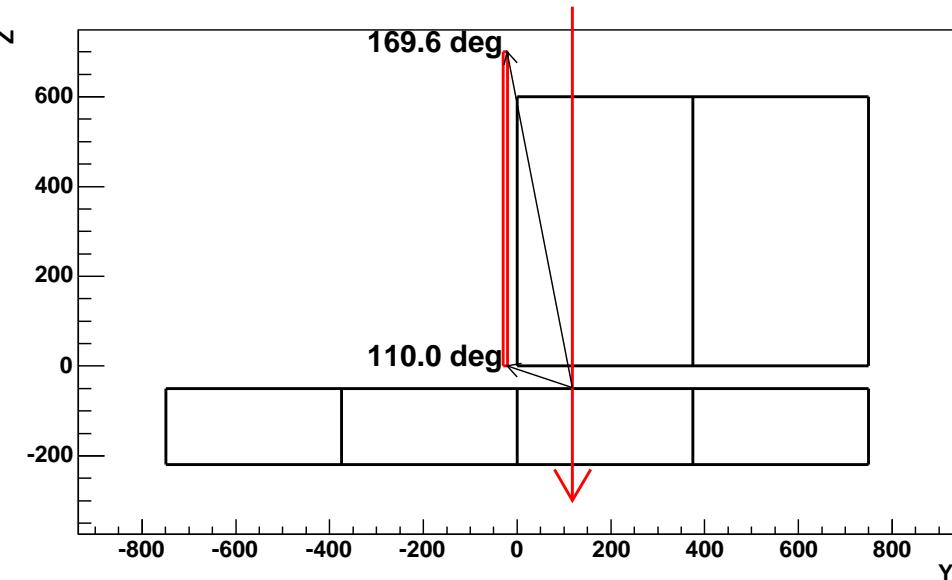
- not in the beam trajectory
 - ⇒ not on the top of the two complete towers
 - ⇒ not on the right hand side of CU

put ACD tiles on the left hand side of the first complete tower



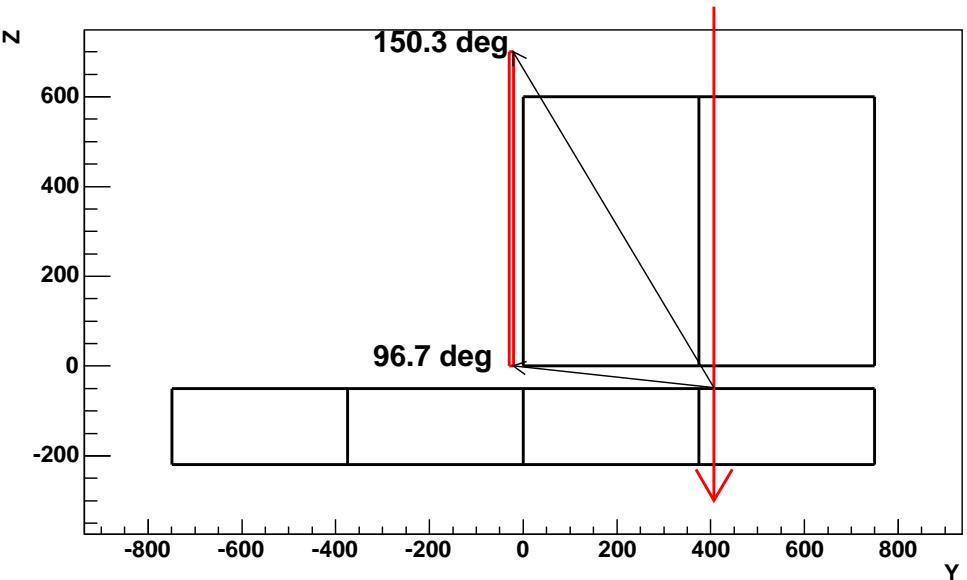
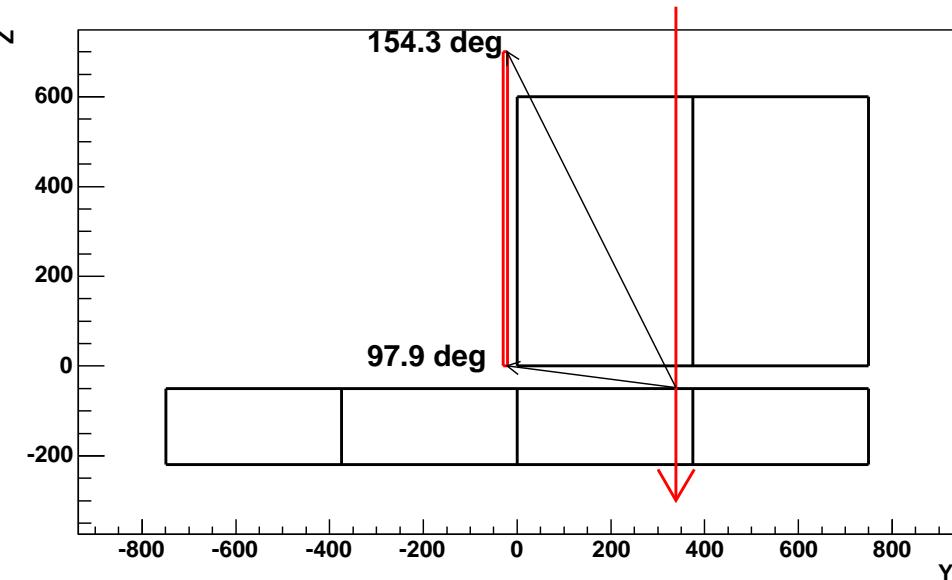
Pure backsplash

try to cover up to 180 deg. while being far enough from cracks



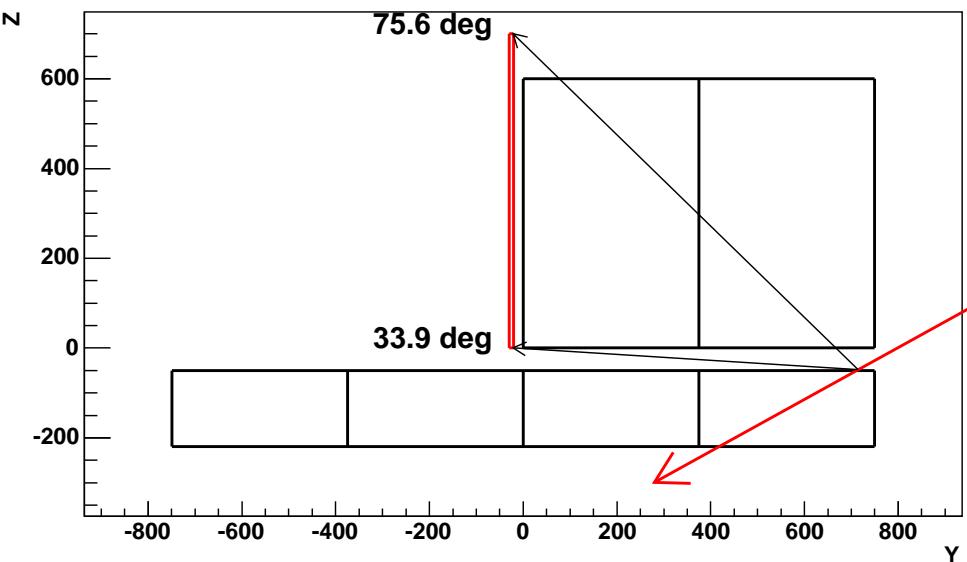
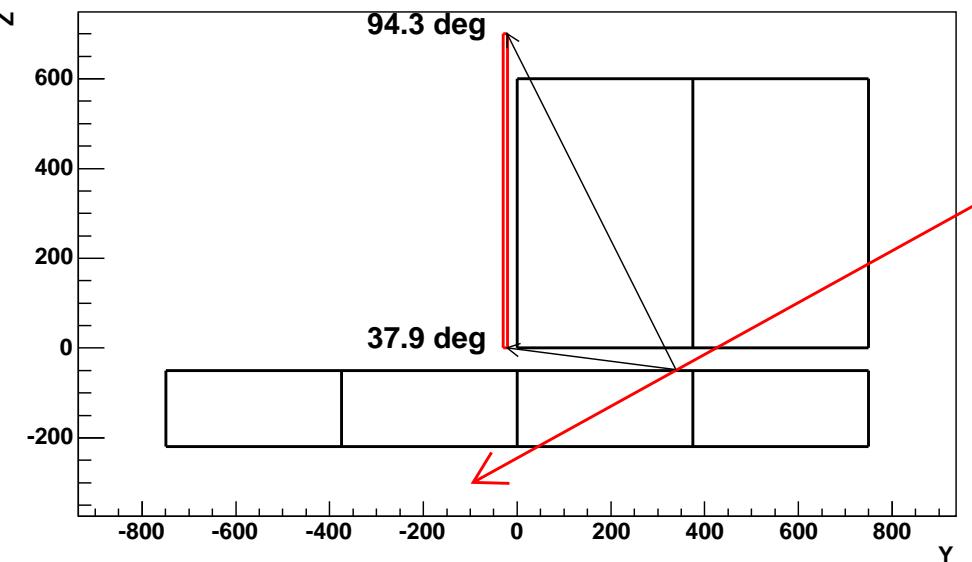
Crack effect

the crack effect is significant above 140 deg.

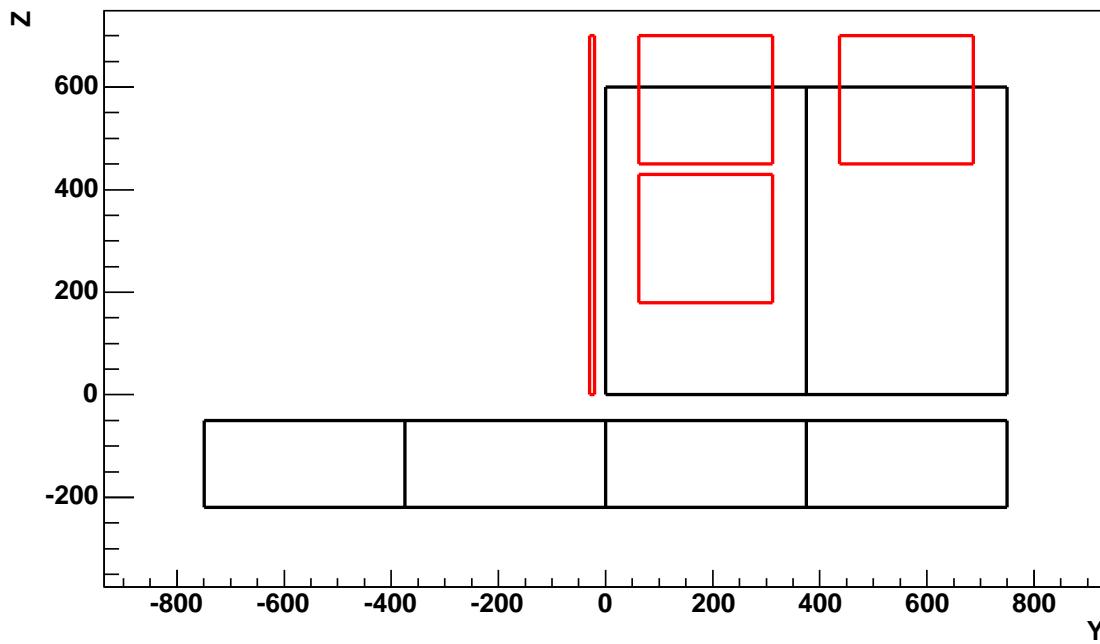


Large incoming angles

incoming angle = 60 deg. \Rightarrow the maximum lies at ~ 80 deg.

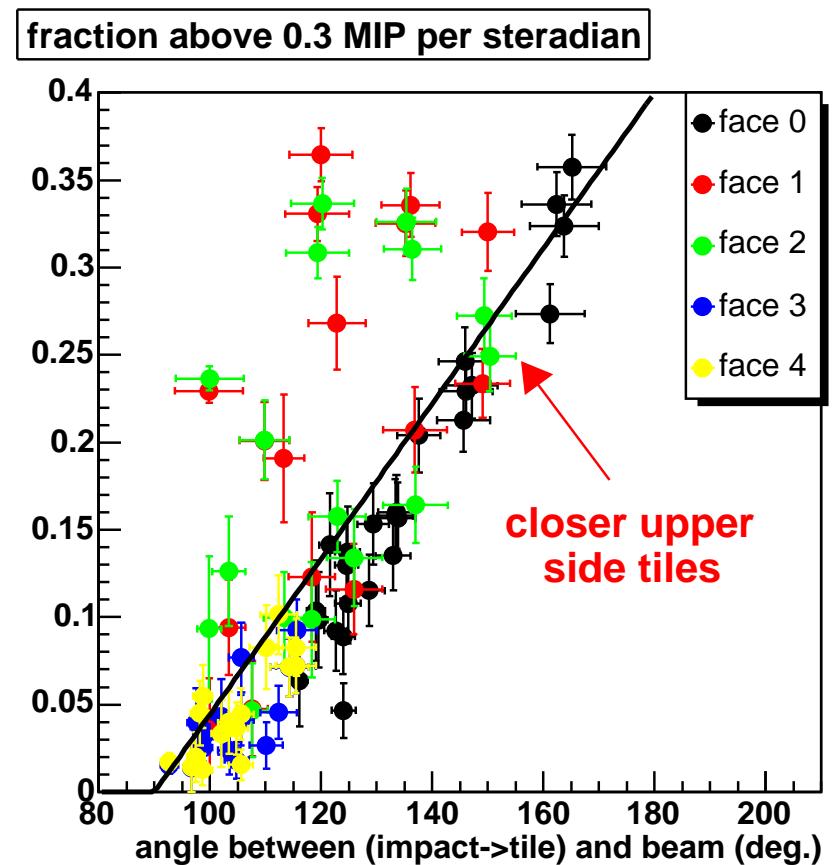
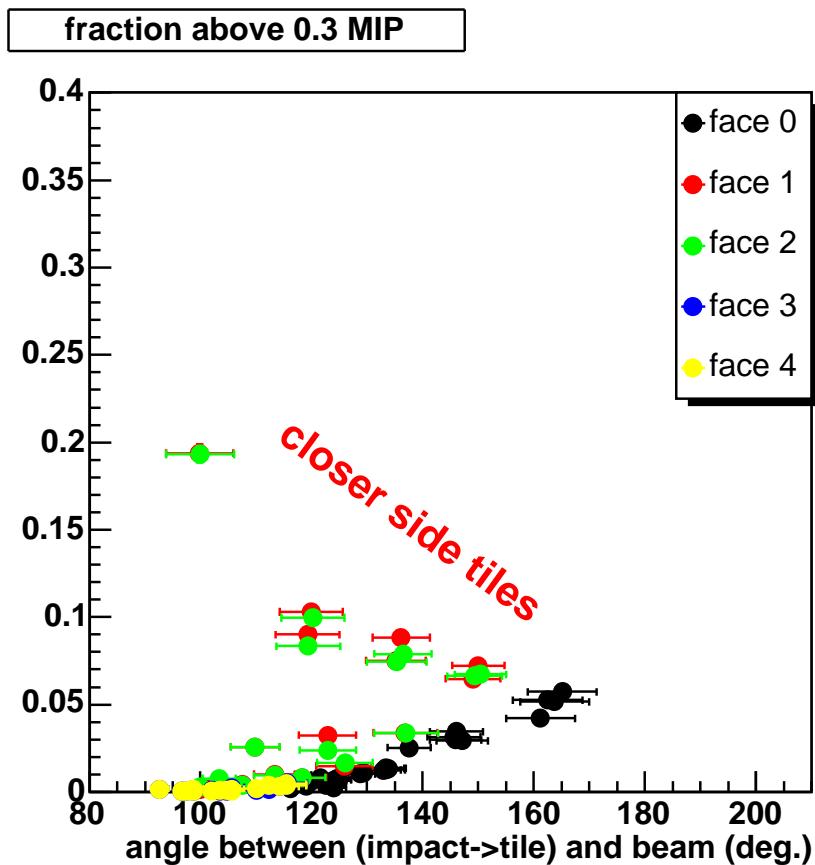


If more tiles are available...

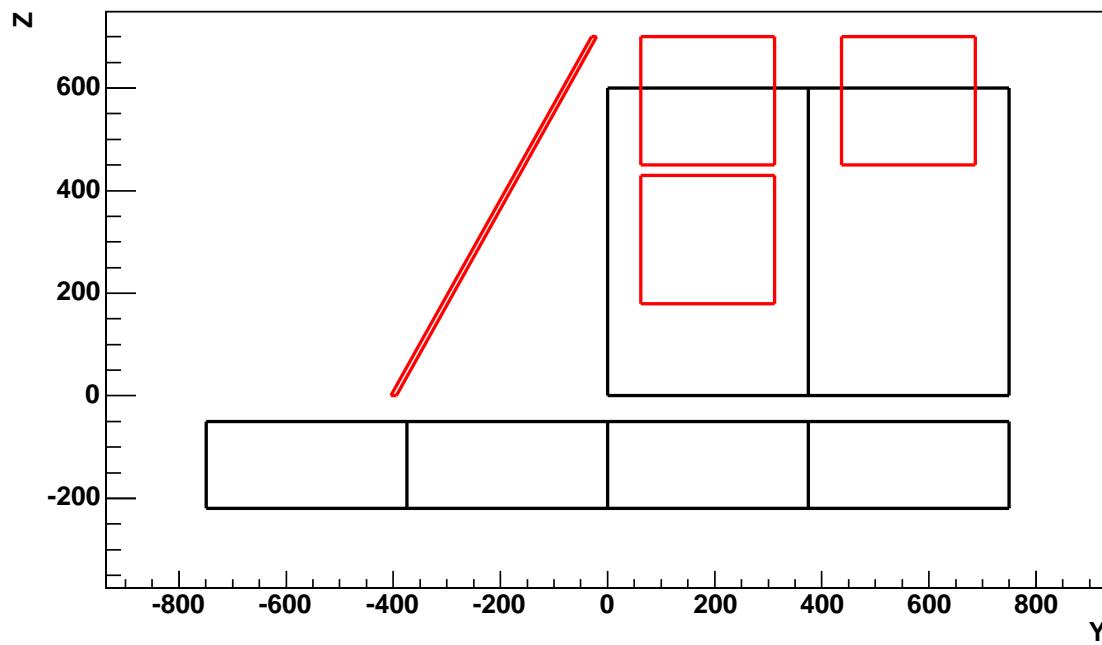


BUT : in the center of a corner tower

- the correction does not work so well...
- should the distance be taken into account ?
- direct “contamination” from the shower ?



A solution ?



Conclusions

- towards a backsplash beam test plan :
 - backsplash, cracksplash and sidesplash
 - preferred position of tiles
 - statistics and configurations
- some remarks :
 - should be confirmed with realistic simulation of CU + ACD tiles with electrons
 - should look at what the tracker measures
 - sensitivity to the beam background