



Retrieving, Filtering and Previewing Data

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Outline



Getting Data from the GSSC Data Server

- Photon Data
- Other files

Looking at the Photon Data

- *d*s9
- *fv*
- gtbin
- Making Cuts on the Data with gtselect
- Good Time Intervals
- Looking at the Exposure History

Getting Data from the GSSC



A quick review from my talk yesterday:

• The main web portal can be found at

http://glast.gsfc.nasa.gov/ssc/dev/databases/DC2

GSSC is providing

- Photon Data
- Pointing and Livetime history Data
- GBM data for the bursts
- Livetime cubes
- Pulsar Ephemeris
- The Source Catalog
- The Galactic Diffuse Emission Model

Getting Photon Data



http://glast.gsfc.nasa.gov/cgi-bin/ssc/LAT/DC2DataQuery.cgi

- Selections can be made on Position, Time and Energy.
- You will get back all event classes
- The valid time ranges are

	Start Time	End Time	
Gregorian Dates	01-01-2008 00:00:01	25-02-2008 00:00:01	
Mission Elapsed Time	220838401	225590401	

- If you want to do object name lookup be sure to select "Object (SIMBAD Resolver)" in the coordinate system box.
- Time only searches will return more data than you ask for at the beginning and end of the query times selected.

Getting Photons (Some Examples)



- Example 1 Galactic anti-center region 20° radius centered on the crab pulsar – full data set (all time and energy)
 - Object Name or Coordinate => crab pulsar
 - Coordinate System => Object (SIMBAD Resolver)
 - Area to Search => 20 Circle
 - Click "Start Search"

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Example 2 – LAT Data for GRB080105885 – 2 hours on either side of burst time.

- Object Name or Coordinate => 322.077, 9.55261 (from Burst page)
- Coordinate System => J2000
- Area to Search => 15 Circle
- Observation Dates => 5-1-08 19:15:39, 5-1-08 23:15:39 Gregorian
- Click "Start Search"



Questions So Far?

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- It's been up for a few days.
- Queries have been made
- Any specific questions?

Other Data at the GSSC



- Pulsar Ephemeris, Source Catalog, FT2 file and Galactic Diffuse Emission Model are each single FITS files that can be downloaded from the main data access page
- Livetime cubes Link on main page takes you to a download page for these files
 - Download the entire set or individual files
 - Each file covers one day (midnight to midnight) of the simulation
 - More on how to use these later in the talk
- Burst Data Link on main page takes you to a page with information about the bursts
 - Name, date and time, duration
 - Position and position error
 - Intensity

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- Link to download GBM Burst data

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FITS File Viewers



ds9

- Third generation SAO Image FITS image viewer
- Download and install from http://hea-www.harvard.edu/RD/ds9
- Used for viewing images

► fv

- FITS file viewer that is part of the NASA HEASARC ftools package
- Download and install from

http://heasarc.gsfc.nasa.gov/docs/software/ftools/fv

- Generic FITS file viewer
 - images, tables
 - Can make plots and histograms

Other FITS file viewers and converters can be found at

http://fits.gsfc.nasa.gov/fits_viewer.html

Quick-view with ds9



So you've got some data and want to look at it.

Try this:

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- ds9 -bin factor 0.5 0.5 -cmap b -scale histequ "<filename>[bin=RA,DEC]"
 - -bin factor 0.5 0.5 sets 0.5 degree bins on RA and Dec
 - -cmap b selects the color map named b (my preference)
 - -scale histequ select the histogram equalized scale for the color map (scale sqrt is also a good one)
- Can eliminate need for [bin=RA,DEC] if you set the environment variable DS9_BINKEY equal to [bin=RA,DEC]

Doing this with the data selection on the anti-center region gives

Anti-center Quick-view Image





Exploring with fv

- fv gives a little more interactive control and allows you to look at the data in different ways.
- Start it up by typing: fv <filename> and you get two windows:
 - At right: main fv menu window
 - Below: Main file option window
 - One line per extension in the file
 - Buttons to look at the extension header, the table data or make plots and histograms

9	v fv: Summary of anticenter_20_full.fits in /local/data/fafnir3/tstephen/DC2/								. 🗆 X	
	File Edit	Tools								Help
	Index	Extension	Туре	Dimension			View			
	0 Primary		lmage	0	Header Image Table		Fable			
	□ 1	EVENTS	Binary	19 cols X 148217 rows	Header	Hist	Plot	All	Select	
	_ 2	GTI	Binary	2 cols X 782 rows	Header	Hist	Plot	All	Select	
Ų										



Making a Map with fv

Click on the "hist" button

- You get the window at top right
- Select RA in the X column
- Select Dec in the Y column
- Set the Min, Max and bin size for each column (image at bottom right)
- You can use a column (such as Energy) as a weight if you desire
- Click on the "Make" button to generate the map.

TV: Histogram			
Mak	e a 1D or 2D histogra	am by binning 1 or 2 tabl	e columns
	Х	Y	Weight
Column Name	•		
TLMin			
TLMax			
Data Min			
Data Max			
Min			
Max			
Bin Size			
Row Range			
		e selected rows	
Make/Close	Make	Cancel	Heln

💙 fv: Histogram			_ O X				
Make a 1D or 2D histogram by binning 1 or 2 table columns							
	×	Y	Weight				
Column Name	RA 💌	DEC					
TLMin	0.0	-90.0					
TLMax	360.0	90.0					
Data Min	61.9665336609	2.0107450485					
Data Max	105.2365798950	41.9806327820					
Min	61	2					
Мах	106	42					
Bin Size	0.1	0.1					
Row Range							
Use selected rows							
Make/Close Make Cancel Help							





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fv Generated Map





Making a Histogram with FV

You already know how as we used the interface to make a map.

- This time just fill in the X column and leave the Y column blank.
- If we use the Energy column for the anti-center region we get the image at right
- Not very interesting on a linear scale so select Edit->Axes Transform->Log-Log to get





Making a Histogram with FV

You already know how as we used the interface to make a map.

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- This time just fill in the X column and leave the Y column blank.
- If we use the Energy column for the anti-center region we get the image at right
- Not very interesting on a linear scale so select Edit->Axes Transform->Log-Log to get this image
- Note: scale on bottom is wrong. I don't know why.







Binning on Time

- Let's look at a different file
- Binning the data we extracted on the burst by time (parameters below) gives the image on the right

•	fv: H	listogram						
Make a 1D or 2D histogram by binning 1 or 2 table columns								
'	X	Υ	Weight					
Column Name	TIME	_						
TLMin	0.0							
TLMax								
Data Min	221259273.3918867111							
Data Max	221266596.9589495659							
Min	221259273							
Мах	221266597							
Bin Size	30							
Row Range								
Use selected rows								
Make/C	lose Make	Cancel	Help					



Looking at the Raw Data

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If you just want to look at the data, click on the "All" button for the extension you want to look at. This will give you a table like the one below

×	✓ fv: Binary Table of burst.fits[1] in /home/tstephen/DC2/							
File Edi	t Tools							Help
		BA	DEC	L	B	THETA	_ PHI	
Select	E	E	E	E	E	E	E	
🗌 Ali	MeV	deg	deg	deg	deg	deg	deg	
Invert								
1	8.488889E+01	3.196125 E+ 02	3.661224E-01	5.210165E+01	-3.197961E+01	3.298398E+01	7.285530E+01	ΠAI
2	6.169908E+01	3.227498E+02	1.073976E+01	6.394409E+01	-2.841150E+01	3.206598E+01	9.433624E+01	
3	1.968599E+02	3.255253E+02	3.484162E-01	5.608533E+01	-3.689734E+01	3.312002E+01	8.426746E+01	
4	1.214815E+02	3.237915E+02	1.148063E+01	6.533534E+01	-2.872891E+01	2.864745E+01	1.049374E+02	
5	9.264123E+01	3.217905E+02	8.900364E+00	6.163348E+01	-2.881450E+01	2.497474E+01	1.029084E+02	
6	8.429630E+01	3.222105E+02	3.864159E+00	5.723811E+01	-3.216133E+01	6.111385E+00	1.023245E+02	
7	1.329176E+02	3.210327E+02	6.610307E+00	5.902215E+01	-2.959751E+01	4.006523E+00	1.270756E+02	
8	3.750802E+02	3.209325 E+ 02	8.264104E+00	6.046975E+01	-2.852794E+01	3.996705E+00	1.510800E+02	
9	1.621546E+02	3.196081E+02	5.888457E+00	5.739981E+01	-2.887485E+01	2.804486E+00	1.081583E+02	
10	1.033333E+02	3.205126E+02	1.372985E+00	5.367768E+01	-3.218105E+01	5.541411E+00	3.385061E+01	
11	2.455396E+02	3.201808E+02	1.104513E-01	5.221192E+01	-3.259584E+01	6.494734E+01	3.148200E+02	
12	6.681481E+01	3.327098E+02	4.945316E+00	6.635645E+01	-3.967527E+01	3.704285E+01	9.803329E+01	
Go to:	Edit	cell:						

Looking at the Headers



- Clicking on the "Header" button displays the FITS header keywords and values for the selected extension
- There is a variety of useful information here
 - Start and stop times of the files in both MET and UTC
 - Selection keywords

- Descriptions of the table data
- Total exposure time

gtbin

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- The Science Tools provide a tool to bin up the photon data into different representations
 - Images (maps)
 - Light curves
 - Energy spectra (PHA files)
- Let's take a quick look at each of these
- Full details can be found in the User Workbook at

http://glast-ground.slac.stanford.edu/workbook/sciencetools/sciTools_Home.htm

- Click on Data Selection in the blue navigation bar
- Click on Extracting Data in the gray navigation bar

Making a Map with gtbin

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- The previous two methods don't really incorporate the coordinate systems properly.
- gtbin can make standard FITS images from the event file according to parameters specified by the user.
- We want to use the CMAP option (counts map) for this
- A sample run on the anti-center data we extracted looks like:

```
[tstephen@80 DC2]$ gtbin
This is gtbin version v0r16p3
Type of output file <CMAP|LC|PHA1|PHA2> [CMAP] : CMAP
Event data file name [anticenter_20_full.fits] :
Output file name [anticenter_image.fits] :
Spacecraft data file name [NONE] :
Size of the X axis in pixels [1024] : 500
Size of the Y axis in pixels [1024] : 410
Image scale (in degrees/pixel) [0.1] :
Coordinate system (CEL - celestial, GAL -galactic) <CEL|GAL> [CEL] :
First coordinate of image center in degrees (RA or galactic 1) [86.3] : 83.6
Second coordinate of image center in degrees (DEC or galactic b) [22] : 22
Rotation angle of image axis, in degrees [0] :
[tstephen@80 DC2]$ []
```

gtbin Generated Map



 This time the image is in the proper orientation as the WCS keywords have been added by gtbin.



gtbin Generated Map

NASA

 This time the image is in the proper orientation as the WCS keywords have been added by gtbin.

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 The green crosses are location of sources listed in the DC2 source catalog



Using gtbin to make a lightcurve



- Let's remake that light curve from the burst
- This time we use the LC option to create the light curve
 - For start and start time use the TSTART and TSTOP values from the FITS header
 - Let's make linearly spaced bins with 5 second intervals

• Here's the input:

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[tstephen@80 DC2]\$ gtbin This is gtbin version v0r16p3 Type of output file <CMAP|LC|PHA1|PHA2> [LC] : LC Event data file name [burst.fits] : burst.fits Output file name [burst_lc.fits] : burst_lc.fits Spacecraft data file name [NONE] : Algorithm for defining time bins <FILE|LIN|SNR> [LIN] : Start value for first time bin [221253339] : 221253339 Stop value for last time bin [221267739] : 221267739 Width of linearly uniform time bins [10] : 5 [tstephen@80 DC2]\$ []

gtbin Lightcurve

- We'll look at it with fv
- Use the "Plot" button on the second extension
- Chose the Time and Counts columns
- Get the light curve at right

Se Se	elect Plot Co	olumns = • *			
RowNumber ElementNumber	Click on a column name then select the corresponding plot axis or error bar				
TIME	Axis	Column			
TIMEDEL COUNTS	×	TIME			
	Y	COUNTS			
	X Error				
	Y Error				
	Rows:				
		Use selected rows			
	A	dd My curve to current graph			
<u> </u>	Go	Clear Cancel Help			







Creating an Energy Spectrum with gtbin



- Now let's remake the energy spectrum we make earlier using gtbin instead of fv
- This time we use the PHA1 option to create the light curve
 - For the energy range lets use 20 MeV to 200 GeV
 - We have a lot of events (148217) so let's make 100 equally spaced logarithmic bins

• Here's the input:

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[tstephen@80 DC2]\$ gtbin This is gtbin version v0r16p3 Type of output file <CMAP|LC|PHA1|PHA2> [PHA1] : PHA1 Event data file name [anticenter_20_full.fits] : anticenter_20_full.fits Output file name [energyspec.fits] : energyspec.fits Spacecraft data file name [NONE] : Algorithm for defining energy bins <FILE|LIN|LOG> [LOG] : Start value for first energy bin [20] : 20 Stop value for last energy bin [200000] : 200000 Number of logarithmically uniform energy bins [100] : 100

gtbin energy spectrum



- Again, use the "Plot" button in fv to look at it
- This time we want the CHANNEL and COUNTS columns in the SPECTRUM extension

Se Se	Select Plot Columns						
RowNumber	Click on a column name then select the corresponding plot axis or error bar						
CHANNEL	Axis	Column					
COUNTS	×	CHANNEL					
	Y	COUNTS					
	X Error						
	Y Error						
	Rows:						
		Use selected rows					
		dd My curve to current graph					
	Go	Clear Cancel Help					



Making Further Data Cuts



- Eventually you'll want to make additional cuts on the data
- If they are only cuts on position, time and/or energy you could use the data server
- If you want to cut on these and/or other parameters use the Science Tool gtselect
- gtselect allows cuts on

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- Basic parameters
 - Position, Time and Energy
 - Event Class
- Advanced Parameters
 - Instrument coordinates (Theta and Phi)
 - Zenith Angle
 - Latitude and Longitude

Cutting Out the Crab



- Let's zero in on the Crab pulsar
- We now want a data set with the following characteristics
 - 4 degree radius circle centered on the crab pulsar
 - Only class A events
 - Events above 100 MeV

• Here's the input:

```
[tstephen@sigurd DC2]$ gtselect
Input FT1 file [test_events_0000.fits] : anticenter_20_full.fits
Output FT1 file [filtered_events_0000.fits] : crab_100MeV.fits
RA for new search center (degrees) <0 - 360> [86.4] : 83.6
Dec for new search center (degrees) <-90 - 90> [28.9] : 22
radius of new search region (degrees) <0 - 180> [20] : 4
start time (MET in s) [0] : 220838401.126248
end time (MET in s) [0] : 225590400.922648
lower energy limit (MeV) [30] : 100
upper energy limit (MeV) [200000] :
Event classes (-1=all, 0=FrontA, 1=BackA, 2=FrontB, 3=BackB, 4=class A) <-1 - 4>
[-1] : 4
Done.
[tstephen@sigurd DC2]$ []
```

The Crab Cutout

- Displayed with ds9 and 0.05 degree bins.
- There are now 6077 events from the original 148217.



What is a GTI anyway?



- If you've been paying attention, you've probably noticed that all of these files have a GTI extension in them.
- So what is a GTI?

- The Good Time Interval is a time range when the data can be considered valid.
- The GTI extension contains a list of these GTI's for the file
- How do we interpret these for GLAST
 - From the Data Server
 - The GTI's are the list of times that the LAT was collecting data over the time range you selected
 - Your object will most likely not be in the field of view during the entire time.
 - Additional data cuts made with gtmaketime will update the GTI's based on the specified cuts
 - The Science Tools use the GTI when calculating exposure

Using gtmaketime

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- gtmaketime uses the FT2 file to create a new set of GTI intervals based on the selected parameters
 - Takes an FT1 and FT2 file as input
 - Creates the new GTIs
 - cuts out all the events that don't fall into those GTI's
 - creates a new FT1 file
- You can make cuts on any field in the FT2 file
 - The default is to select time not in the SAA (IN_SAA!=T)
 - Cuts are made using C-style?? relational syntax
 - ! -> not, &&-> and, || -> or, ==, !=, >, <, >=, <=
 - Things like ABS(), COS(), SIN() seem to work as well
- Let's make a cut on the anti-center region file to exclude the SAA and any events collected when the Geomagnetic Latitude was greater than 10°.

Input to gtmaketime

GTL

2

Binary

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Here's the input to make the cut:

[tstephen@dhcpvisitor218227 DC2]\$ gtmaketime Spacecraft data file [test_scData_0000.fits] : DC2_FT2.fits Filter expression [IN_SAA!=T] : IN_SAA!=T&&LAT_GEO<10 Event data file [test_events_0000.fits] : anticenter_20_full.fits Output event file name [filtered.fits] :

It did something, here are the file summaries from fy:

 Image: A start of the start of	fv: Summary of anticenter_20_full.fits in /home/tstephen/DC2/						
File Edit	Tools			Help			
Index	Extension	Туре	Dimension		View	,	
0	Primary	Image	0	Header Image Table		Table	
□ 1	EVENTS	Binary	19 cols X 148217 rows	Header	Hist Plot	All Select	
_ 2	GTI	Binary	2 cols X 782 rows	Header	Hist Plot	All Select	
	fv:	Summary	of filtered fits in /home/ts	stenhen/D	C2/		
File Edit	File Edit Tools						
Index	Extension	Туре	Dimension		View		
	Primary	Image	0	Header Image		Table	
1	EVENTS	Binary	19 cols X 72813 rows	reheell	Hist Plot	All Select	

2 cols X 1690 rows

Header

Header

Hist

Hist

Plot

Plot

All

All

Select

Select

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The New Data Lightcurve





Looking at the Exposure



- Two steps to generating an exposure map
 - Making an exposure cube from the FT2 file
 - Making an exposure map from the exposure cube
- To help speed up the process pre-generated exposure cubes are available for each day of the simulation
- Let's look at how to use these to generate an exposure map
 - With a single exposure cube
 - With multiple exposure cubes

Using a Single Exposure Cube



- In this case we just run the exposure_map tool
- This allows us to control the map generation parameters
 - Map center, size and scale
 - Projection type

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- Many choices (Aitoff, Cartesian, Mercator, Tangential, etc)
- But it's a hidden parameter and doesn't prompt for an entry
- Default is Aitoff
- Energy range and number of energy bins

Note: The default parameter file wants to use the Ext1 name for the exposure extension.

- The exposure cubes use EXPOSURE as the extension name
- You need to give the command line argument table=EXPOSURE or in the GUI select the advanced options box and enter this in the table field

Running exposure_map

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Let's make an all sky Aitoff map of the first day's exposure - 8 bins from 20 MeV to 200 GeV and 0.5 degree spatial bins

• Here is the input and output for this run:

[tstephen@dhcpvisitor218227 DC2]\$ time exposure_map table=EXPOSURE This is exposure_map version N/A Exposure cube input file name [expCubeO.fits] : Count map input file name (NONE for manual input of map geometry) [NONE] : Exposure map output file name [day1exp.fits] : Response function to use, e.g. DC1F/DC1B, G25F/G25B, TestF/TestB [DC2] : Size of the X axis in pixels [720] : Size of the Y axis in pixels [360] : Image scale (in degrees/pixel) [0.5] : Coordinate system (CEL - celestial, GAL -galactic) <CEL|GAL> [CEL] : First coordinate of image center in degrees (RA or galactic 1) [180] : Second coordinate of image center in degrees (DEC or galactic b) [0] : Rotation angle of image axis, in degrees [0] : Start value for first energy bin [20] : Stop value for last energy bin [200000] : Number of logarithmically uniform energy bins [6] : 8 Creating an Exposure object from file expCubeO.fits total elapsed time: O Using Aeff(s) DC2::FrontA, DC2::BackA, DC2::FrontB, DC2::BackB, Creating an Image, will write to file day1exp.fits Generating layer 0 at energy 20 MeV Aeff(0): 6.86067 cm^2 Generating layer 1 at energy 63.2456 MeV Aeff(0): 1140.65 cm^2 Generating layer 2 at energy 200 MeV Aeff(0): 4407.49 cm^2 Generating layer 3 at energy 632.456 MeV Aeff(0): 8239.31 cm^2 Generating layer 4 at energy 2000 MeV Aeff(0): 8996.34 cm^2 Generating layer 5 at energy 6324.56 MeV Aeff(0): 8955.51 cm^2 Generating layer 6 at energy 20000 MeV Aeff(0): 9041.17 cm^2 Generating layer 7 at energy 200000 MeV Aeff(0): 7683.49 cm^2



Exposure Map – Day 1





200-632 MeV







Using Multiple Exposure Cubes



- In order to use multiple exposure cubes we must first combine them into a single cube using gtaddlivetime
 - It only adds two cubes
 - You can't append to an existing cube and keep the same name
 - But it's very quick

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Here's a sample input:

[tstephen@dhcpvisitor218227 DC2]\$ ds9 day1exp.fits [tstephen@dhcpvisitor218227 DC2]\$ gtaddlivetime Livetime cube 1 [expCube_00.fits] : expCube0.fits Livetime cube 2 [expCube_01.fits] : expCube1.fits Output file [expCube.fits] : expCube.fits

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200-632 MeV



2-6.32 GeV

20-63.2 GeV



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Exposure Map – Full Simulation



20-63.2 MeV

200-632 MeV



2-6.32 GeV



20-63.2 GeV



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