



# HARPO: Measurement of polarised gamma rays (1.7 to 72MeV) with the HARPO TPC

Philippe Gros  
for the HARPO collaboration



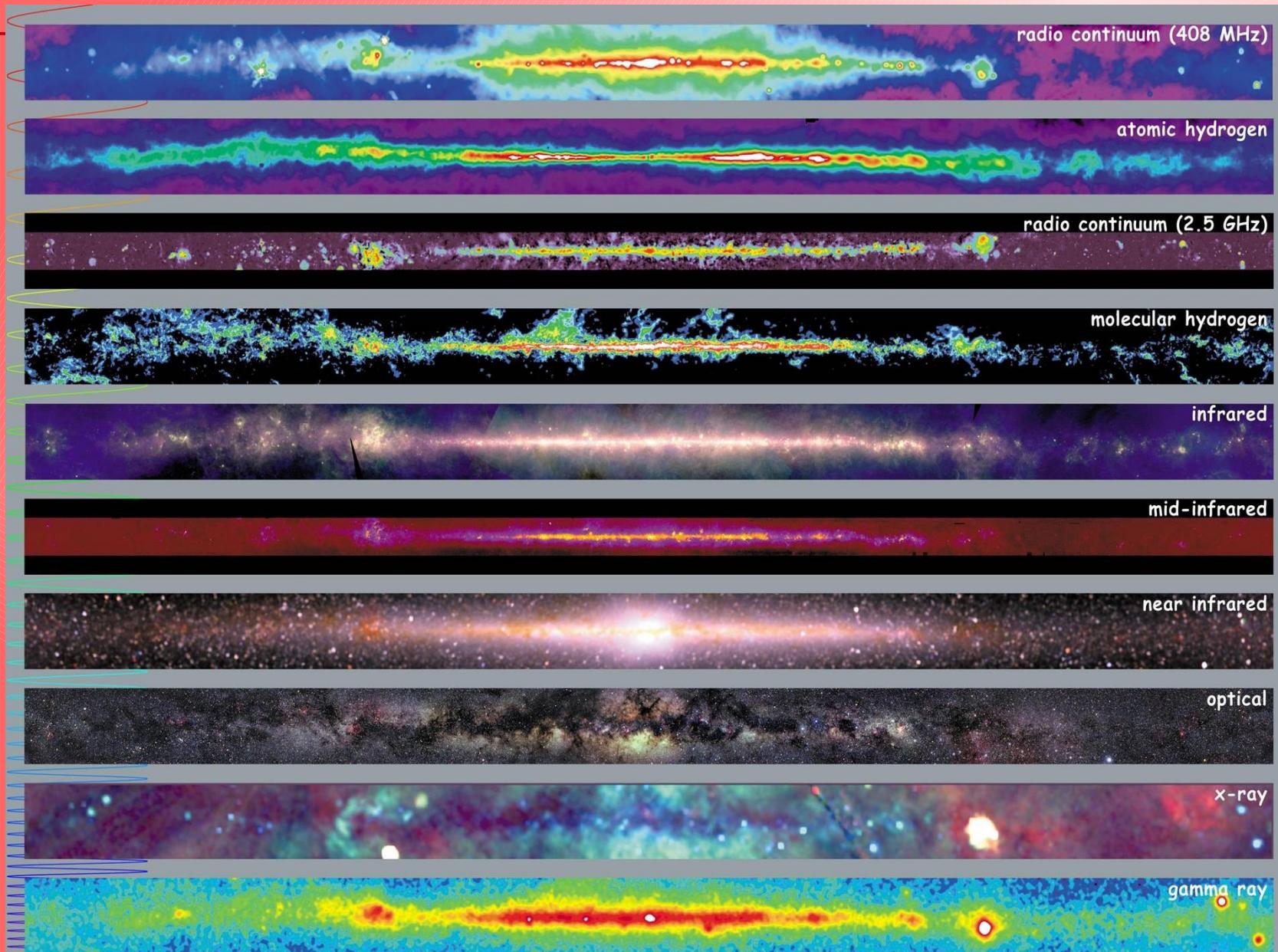
- 
- Introduction to the HARPO project
  - Setup at the NewSUBARU photon beam
  - Gas monitoring
  - Trigger with micromegas signal
  - Other experiences
  - Conclusions



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# Milky Way / 天の河

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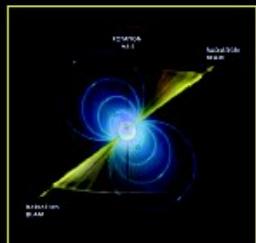


<http://adc.gsfc.nasa.gov/mw>

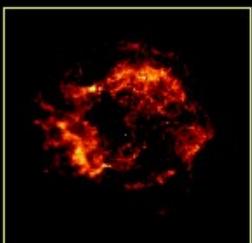
# Gamma Astrophysics non thermal phenomena

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- Galactic targets



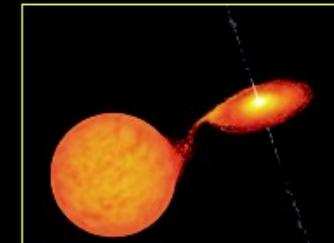
Pulsar



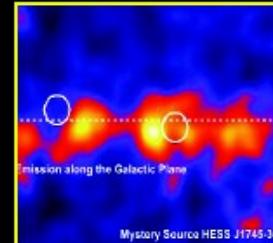
Supernova Remnants



Pulsar wind nebulae



Micro-quasars

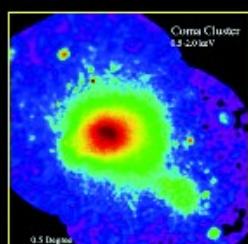


Galactic center

- Extragalactic targets



Active Galactic Nuclei



Galaxy Cluster



Starburst galaxies

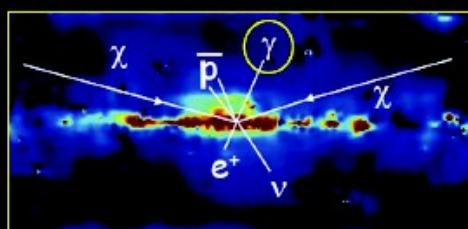


Merging Galaxies

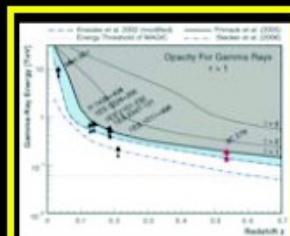


Gamma-ray Bursts

- Fundamental physics



Dark Matter annihilation



Universe transparency

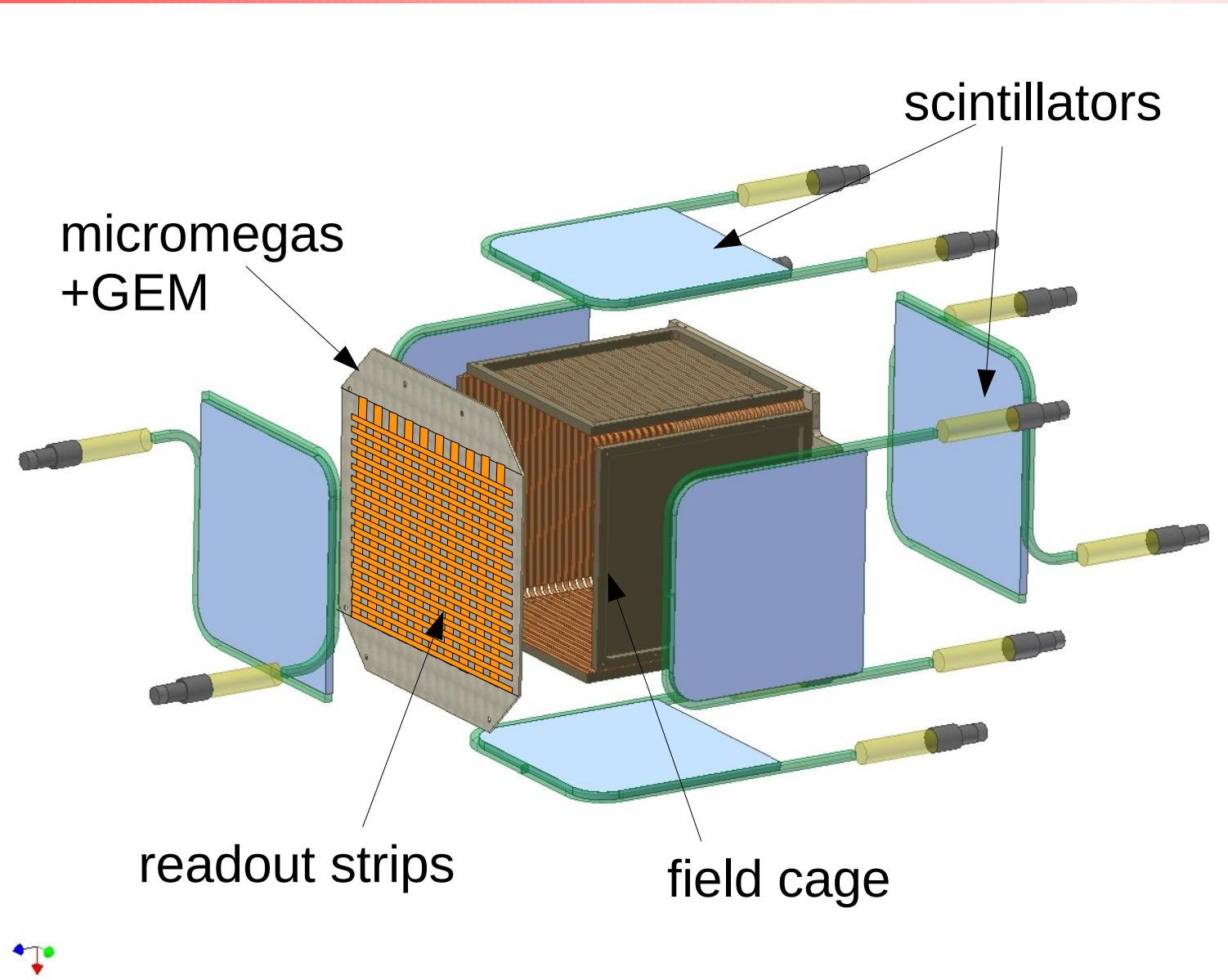
- CR physics
- Lorentz invariance
- Quantum gravity
- Axion-photons obsc



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# HARPO Demonstrator

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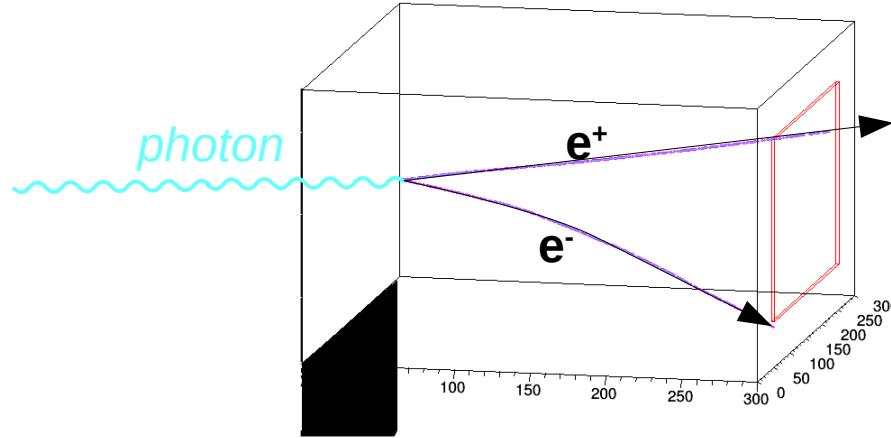




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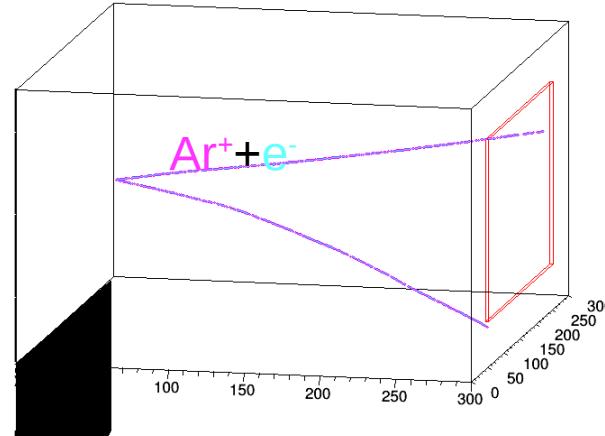
# TPC: photon conversion

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The incoming photon interacts with the gas and decays into an electron-positron pair

# TPC: Gas ionisation



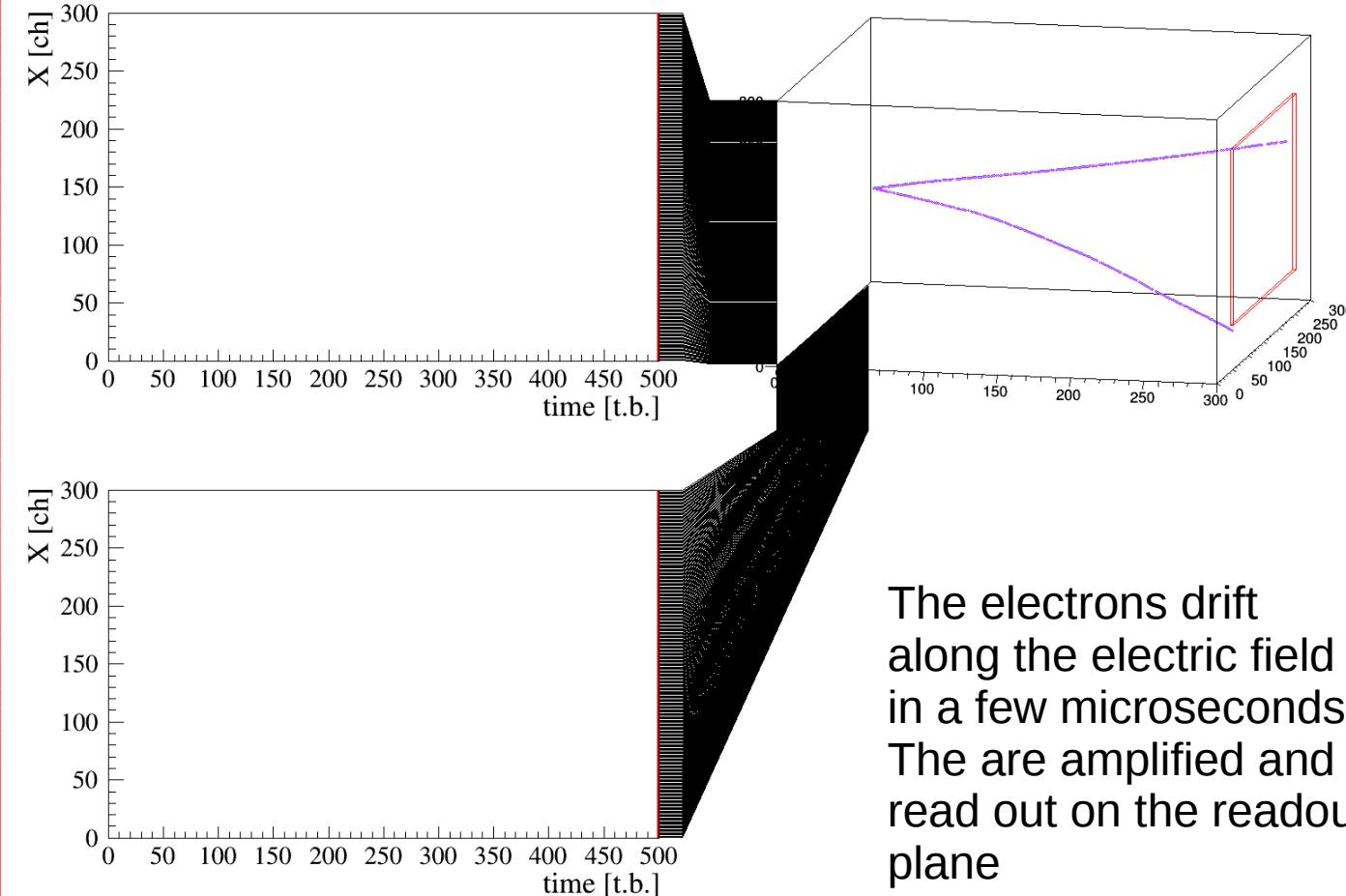
The electron and positron travel through the gas (mostly Argon) and ionises it, freeing many electrons and positive ions  
This takes a few nanoseconds



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# TPC: Drift and Readout

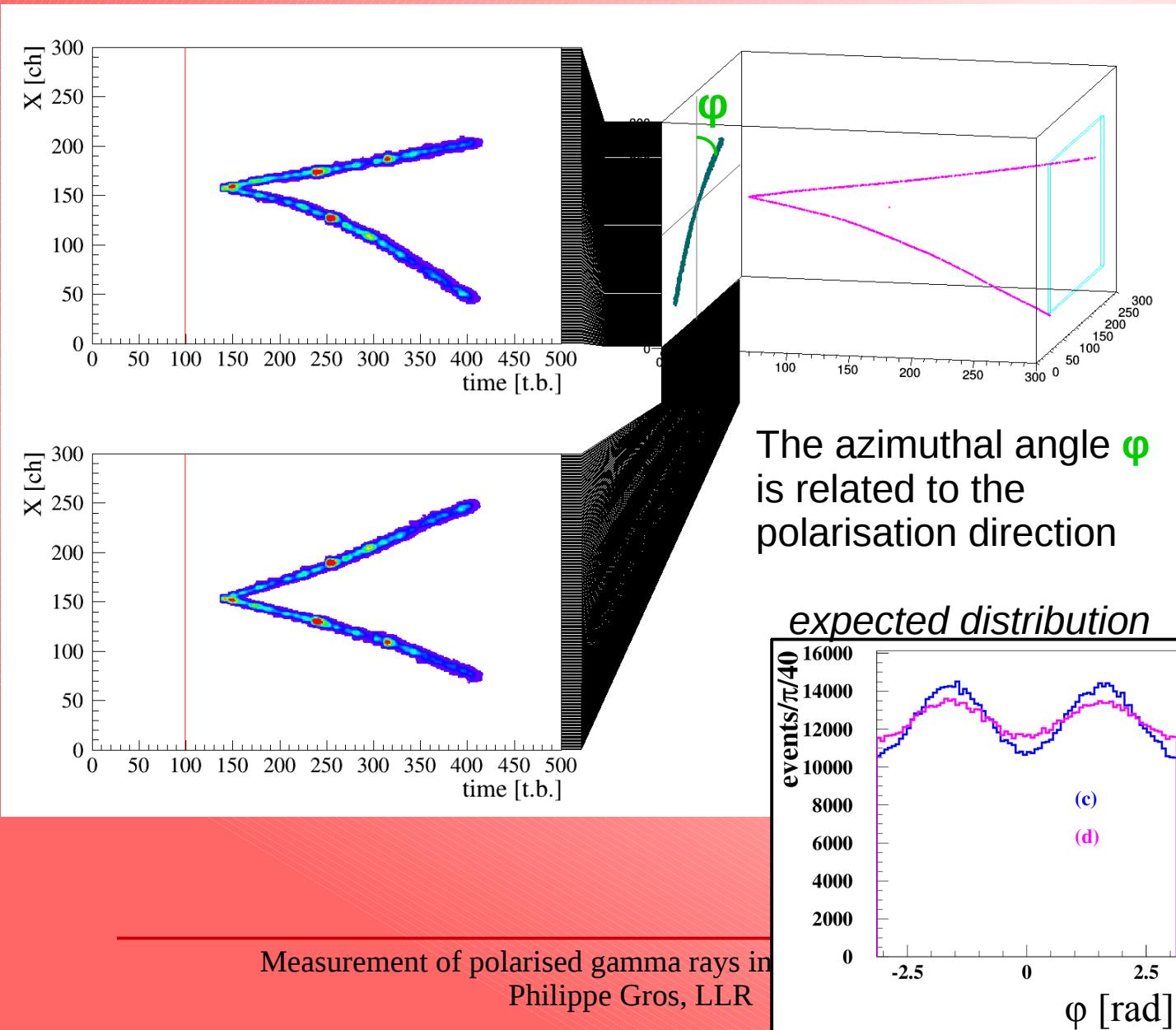
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The electrons drift along the electric field in a few microseconds. They are amplified and read out on the readout plane



# Polarisation measurement



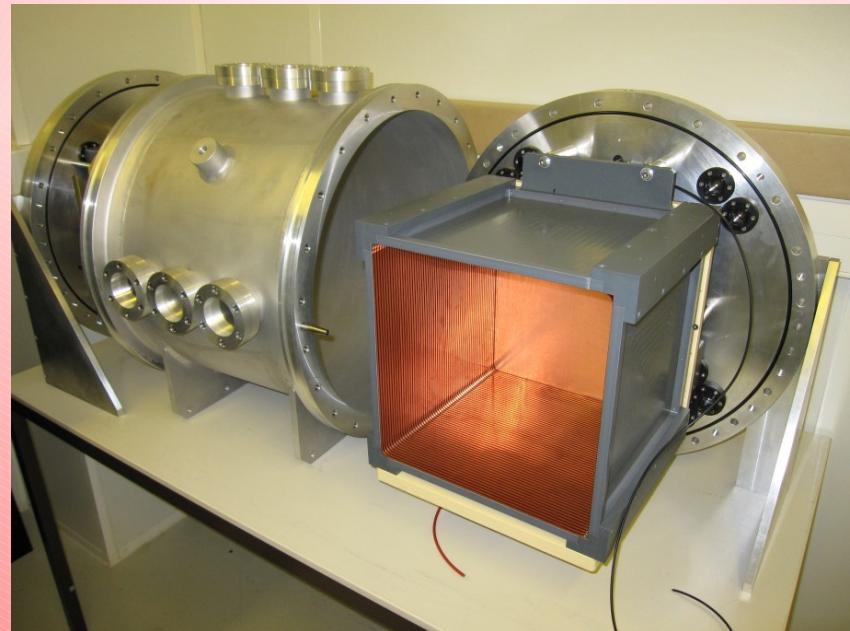


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# HARPO Demonstrator



- Purpose
  - Assess challenges
  - Demonstrate performance in test beam
- Realisation
  - 30cm cubic TPC
  - Ar/iC<sub>4</sub>H<sub>10</sub> 95/5 up to 5bar
  - micromegas+2GEM amplification
  - 2x288 strips readout (x&y), 1mm pitch
  - AFTER readout electronics,  
511 time bins, up to 50MHz (33 used)
  - trigger: 6 scintillators

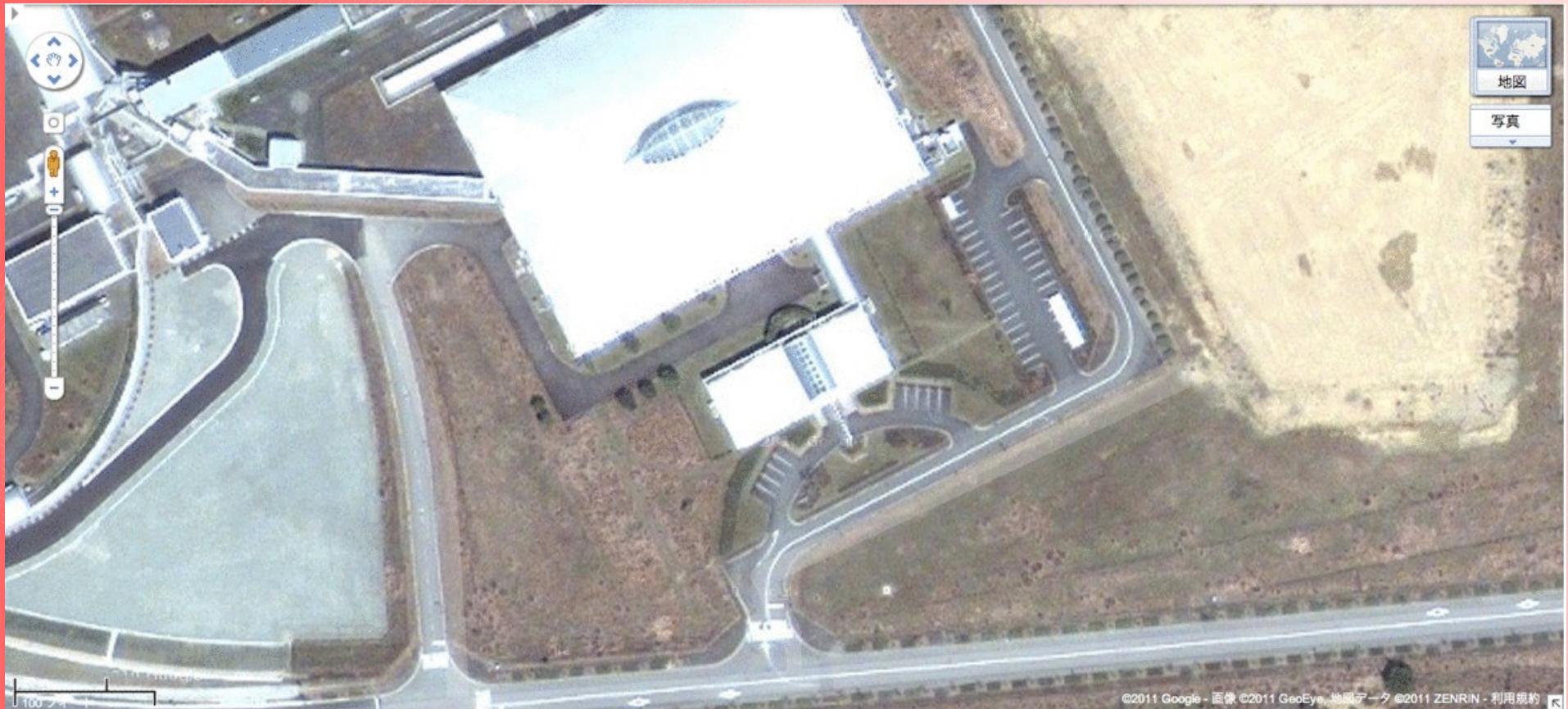




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# Going to Japan

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# Shipping equipment

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Departure from LLR



Arrival in NewSUBARU



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# Shipping equipment

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Departure from LLR



Arrival in NewSUBARU



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# Intermède animalier 1

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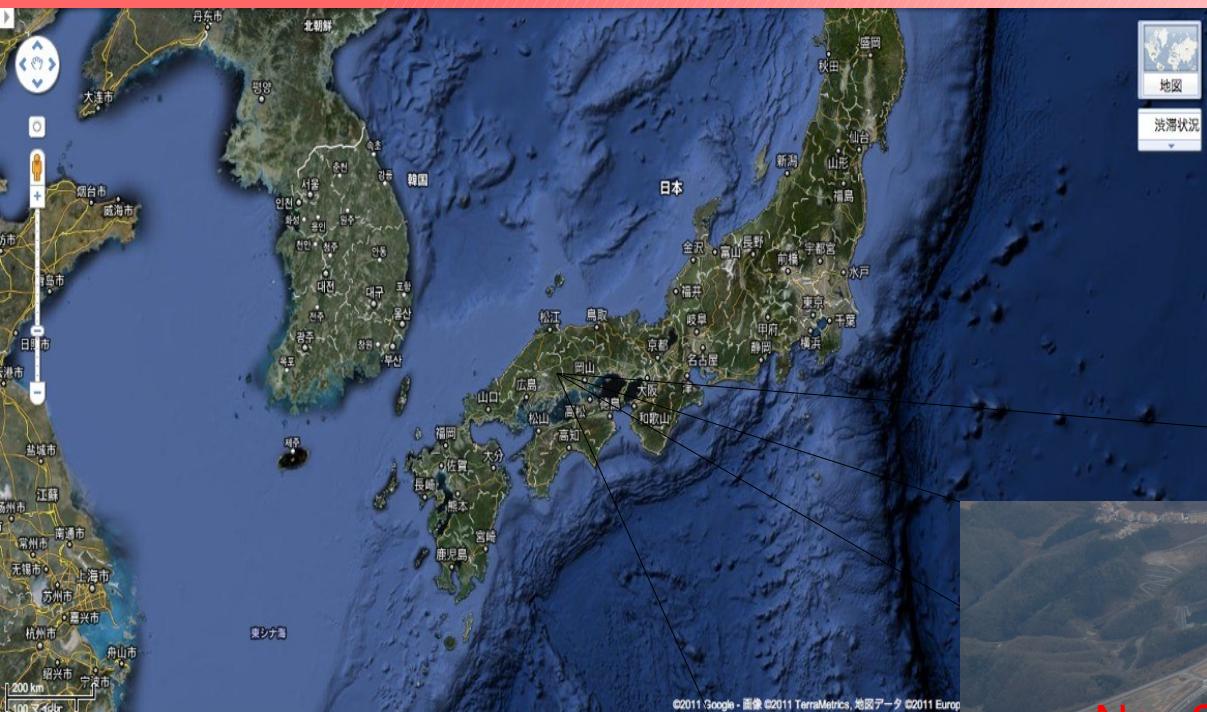




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# NewSUBARU

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2014-12-11

Measurement of pol.

Philippe Gros, LLR

# NewSUBARU photon beam



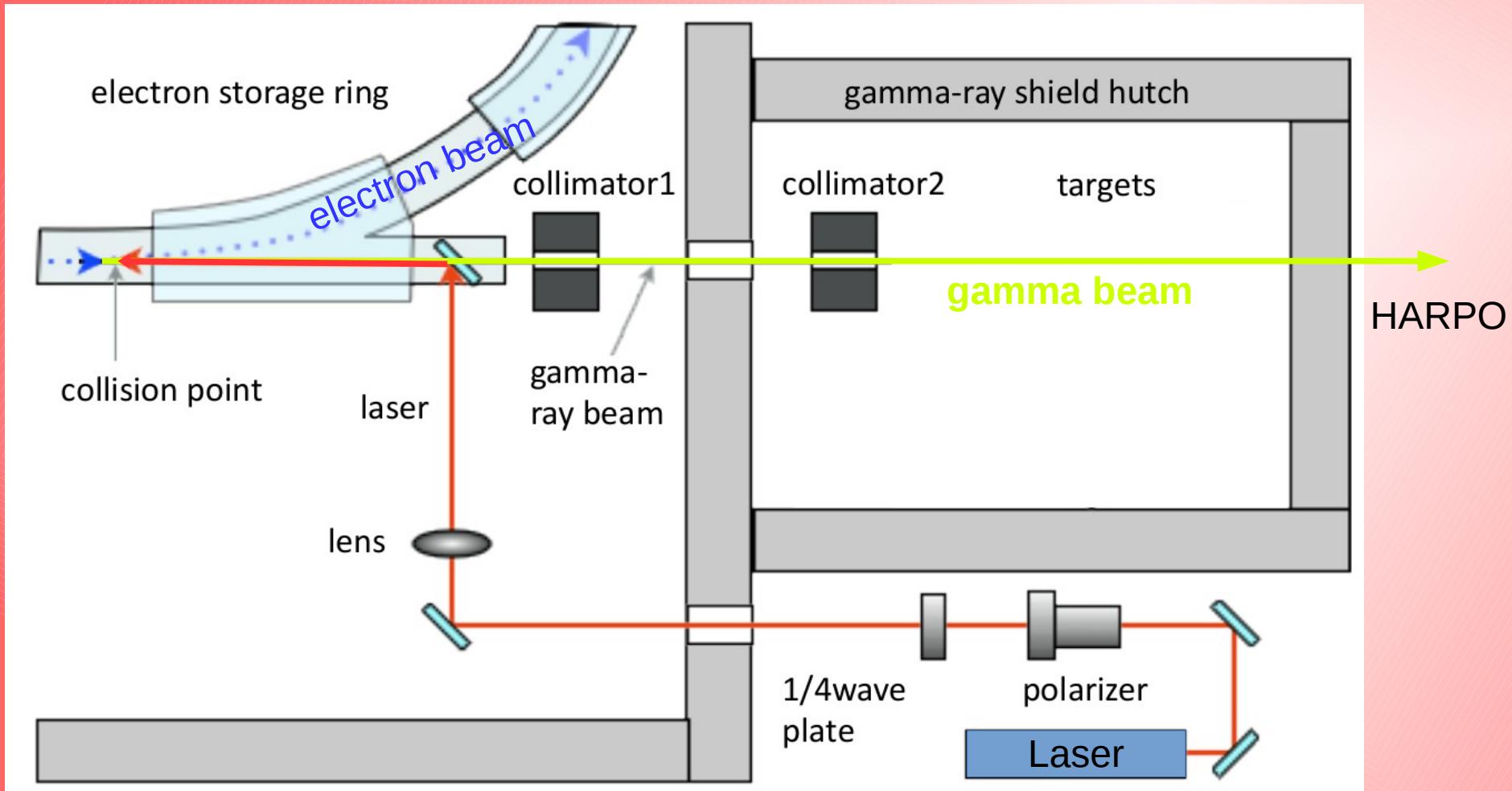
- Polarised gamma ray beam
  - Inverse Compton
  - electron beam 0.6, 1., 1.2 or 1.5 GeV
  - laser Nd (1 $\omega$  or 2 $\omega$ ), Er or CO2
    - => polarised photons 1.71 to 72.3MeV
- Pulsed mode
  - Nd: 20kHz, Er:200kHz, CO2: not



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# Gamma beam

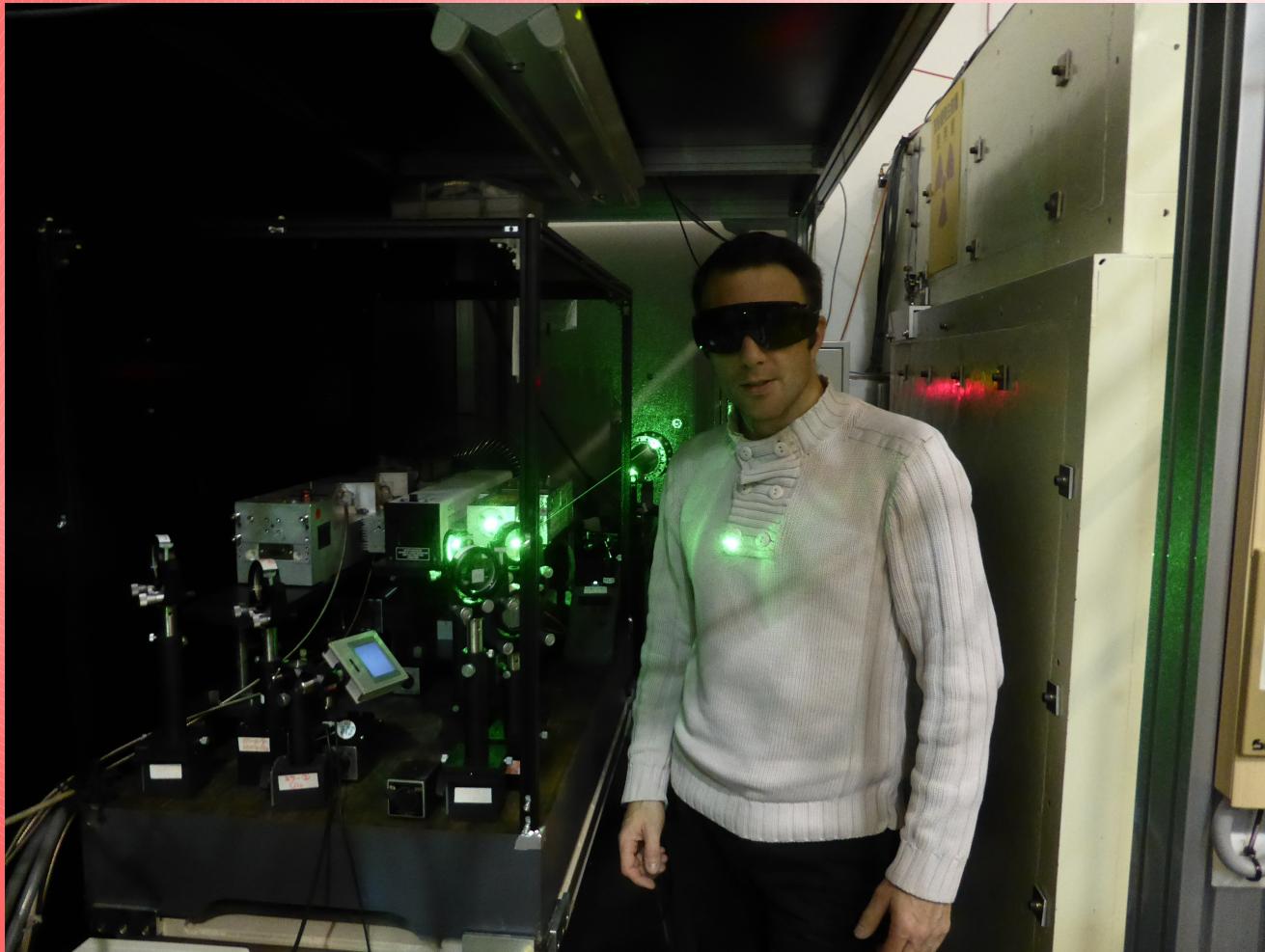




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# Laser beam

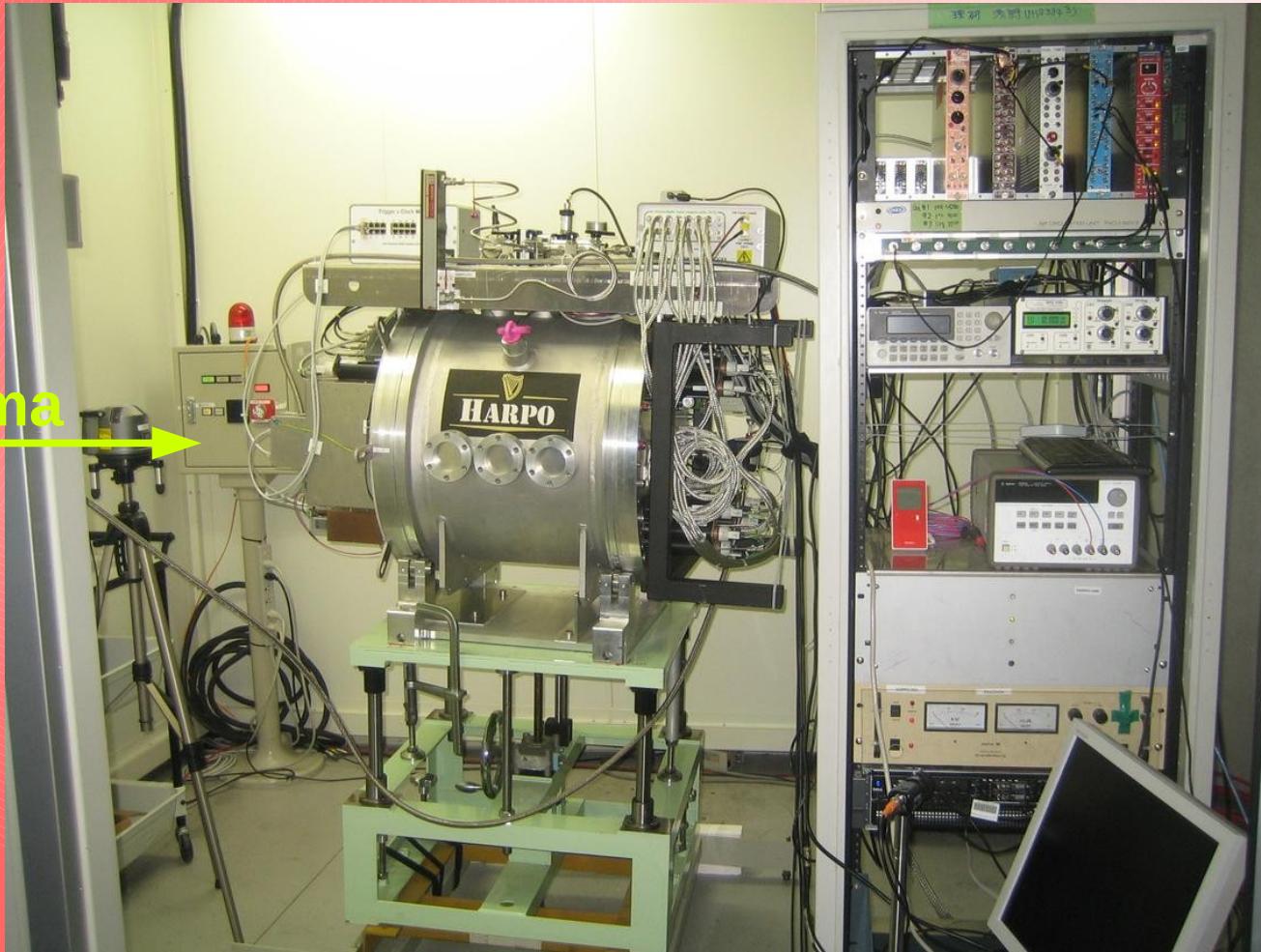




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# HARPO in beam

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# Rotation

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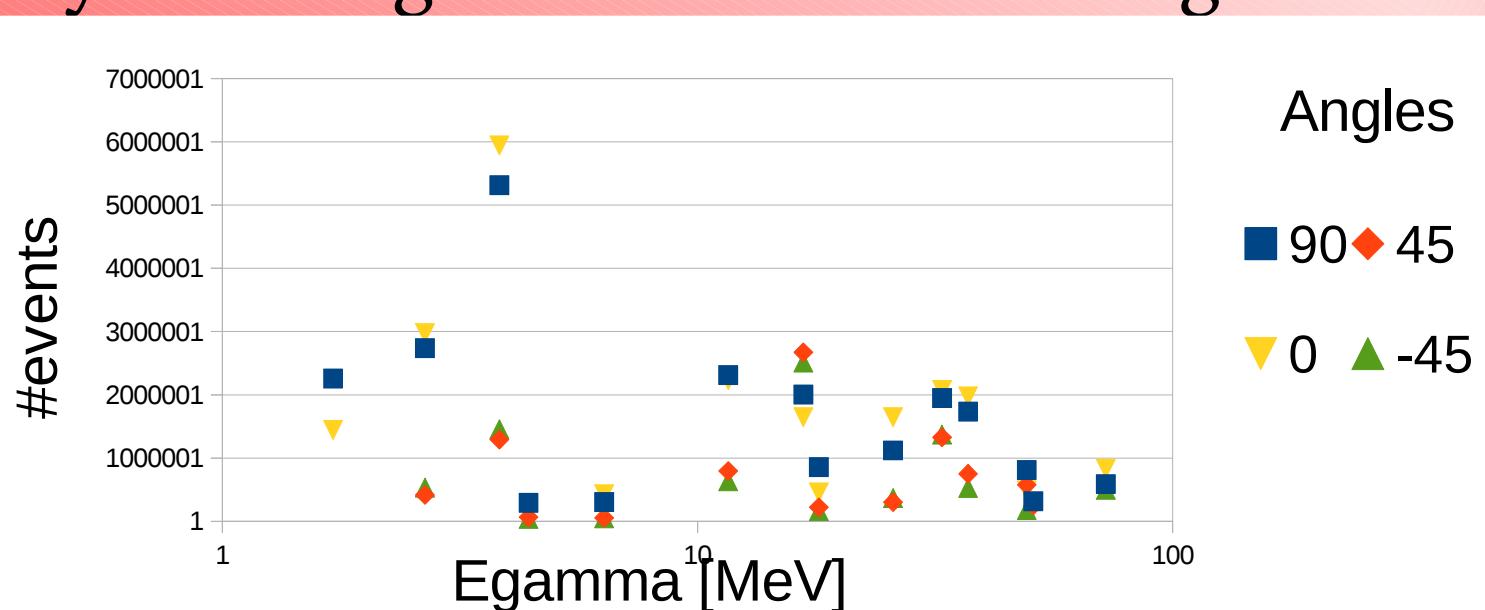


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# Beam campaign



- ~20 days of data taking
- 13 gamma energies, beam polarised or not
- 4 TPC orientation for angle systematics
- >60Mevents, >1TB of data
- probably >20% gamma converted in gas

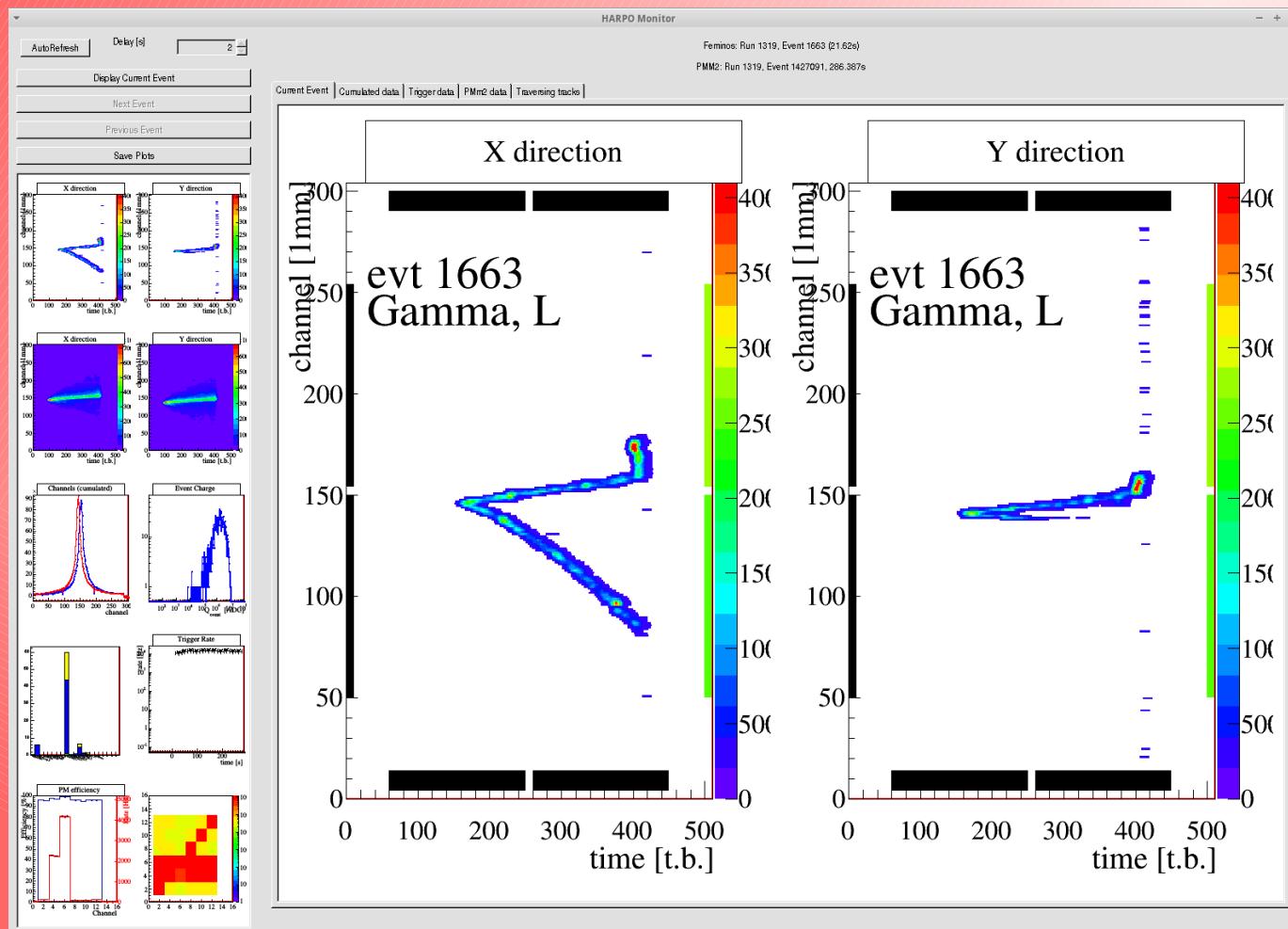




# HARPO

# Monitoring

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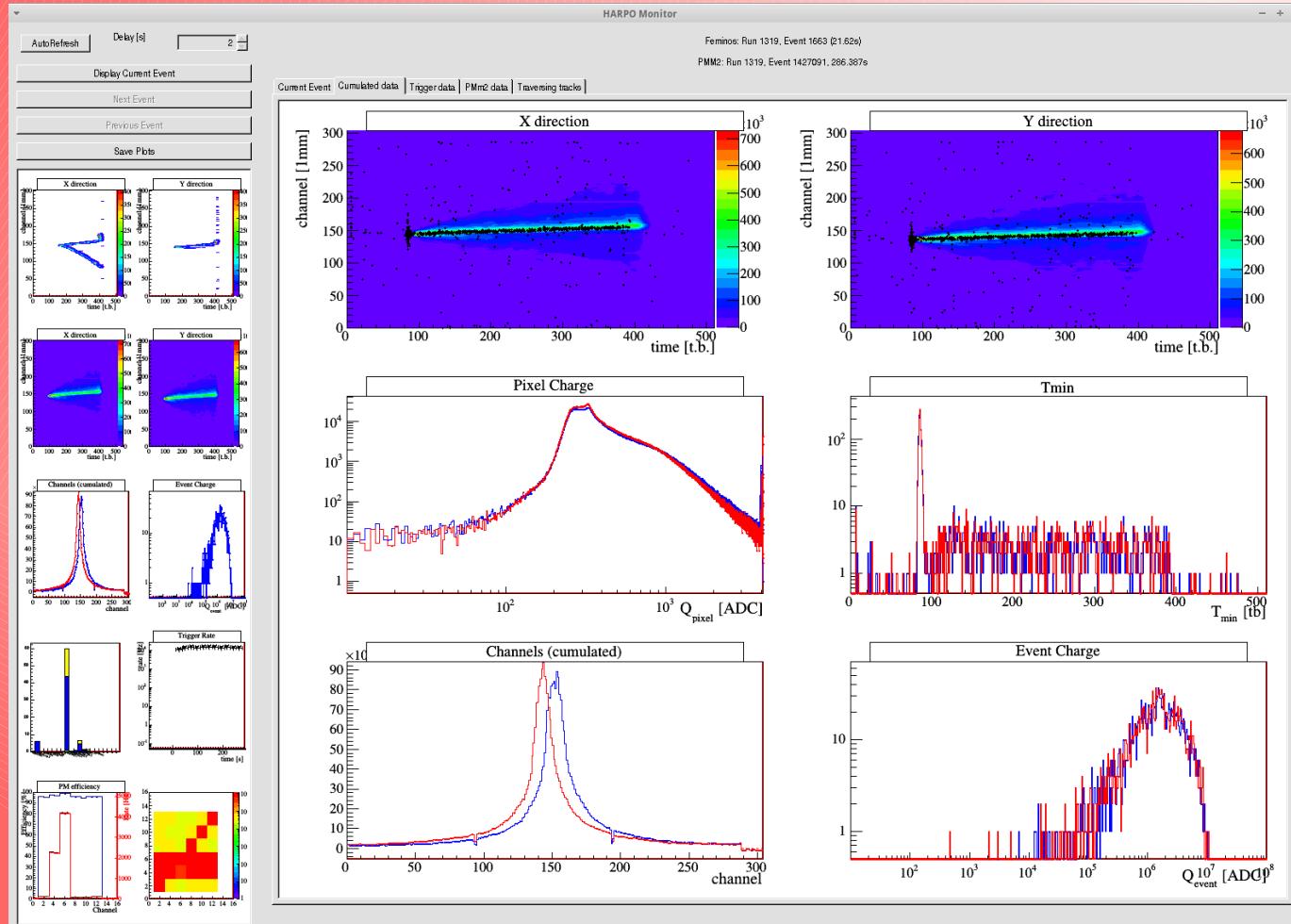




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# Monitoring

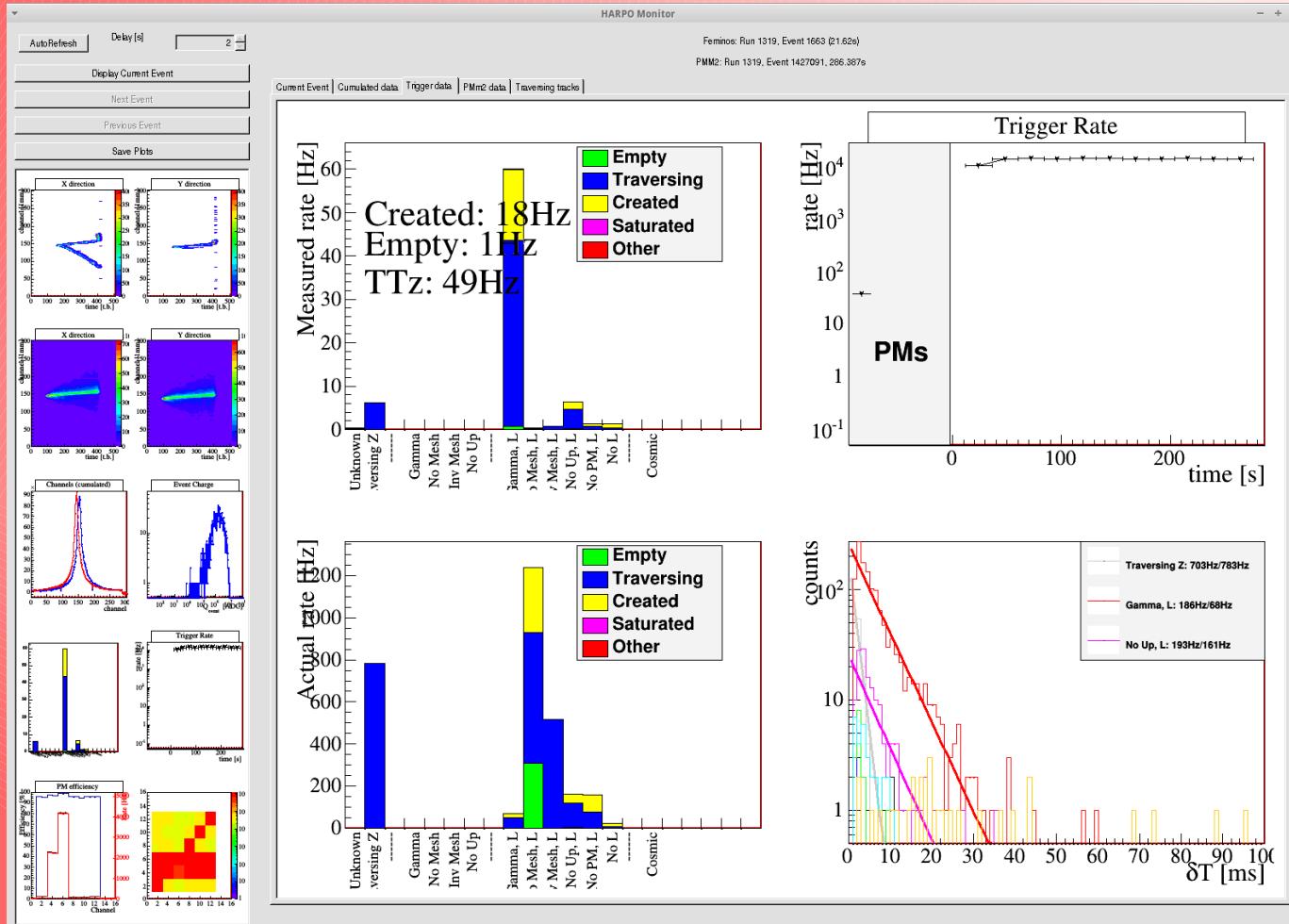




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# Monitoring

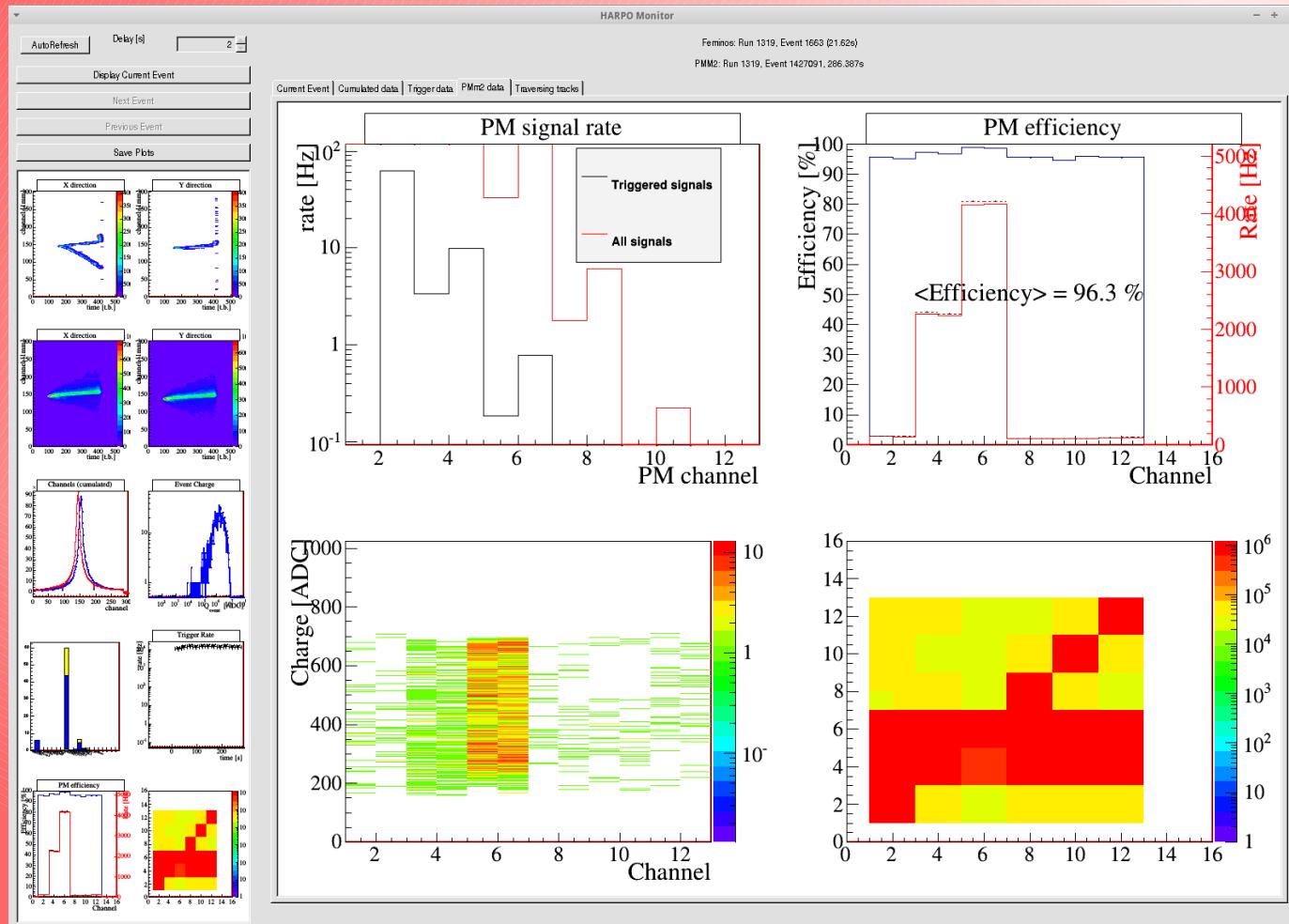




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# Monitoring

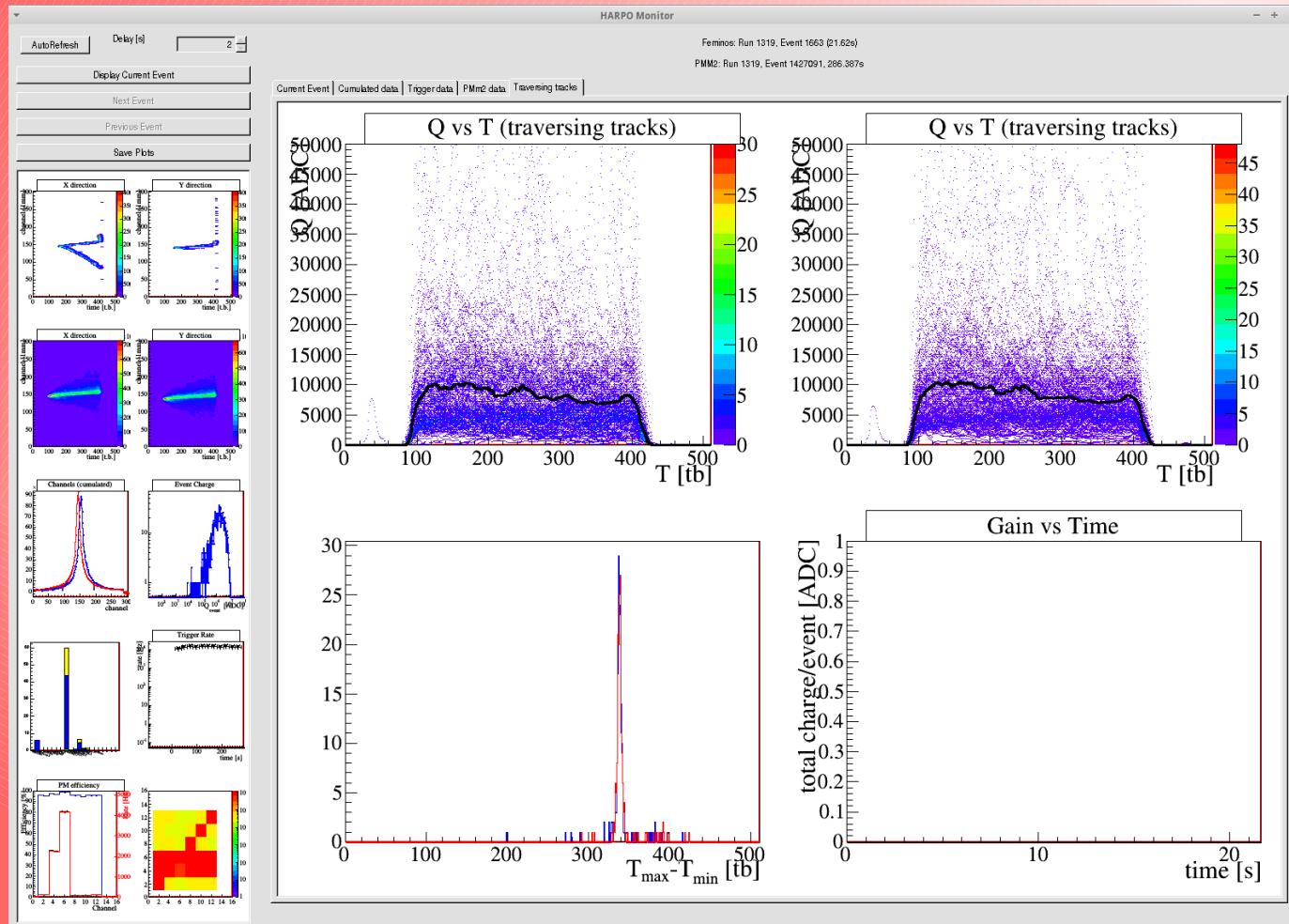




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# Monitoring





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# Intermède animalier 2

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2014



d gamma  
Philippe Gros, ECR



# HARPO gas system



- Sealed vessel
- Leaks minimised in vacuum with He system
- Gas filling procedure
  - vacuum pumping ( $<10^{-5}$ bar)
  - “rinsing” with gas mixture  $\sim 100$ mbar
  - vacuum pumping ( $\sim 10^{-7}$ bar)
  - 2 bar fill



# Gas monitoring



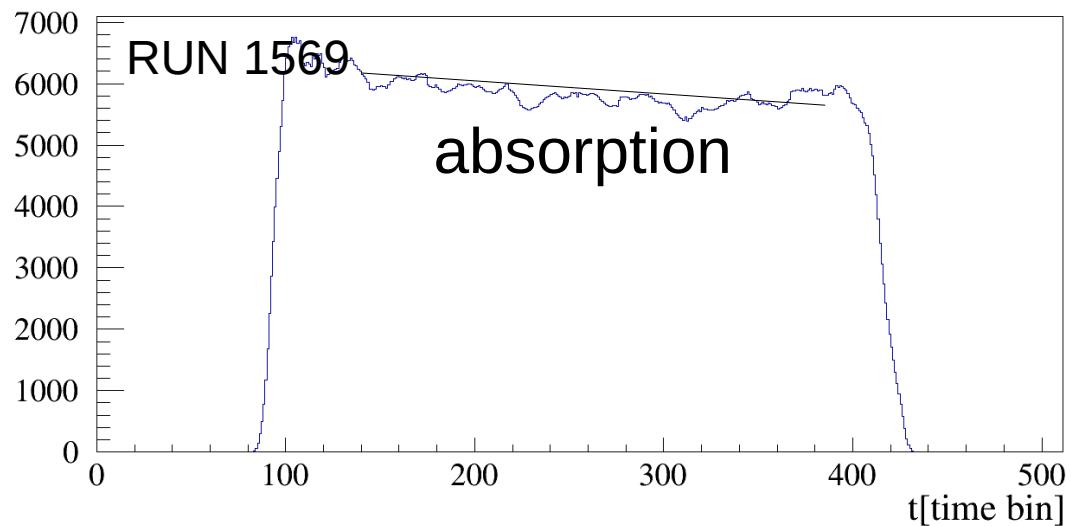
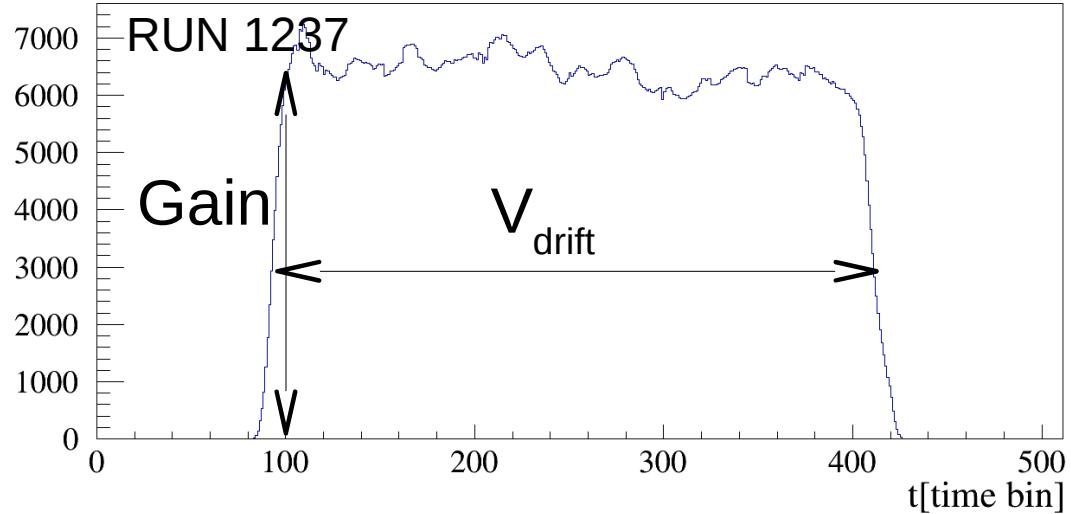
- No direct measurement of the gas
- Monitoring with track data
  - High momentum traversing tracks in Z
  - $\sim$ uniform energy deposition along Z (if corrected for angle)
  - Simple access to gain, drift velocity and absorption
  - Dedicated trigger line =>  $\sim$ 2% of the events



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# Cosmic rays

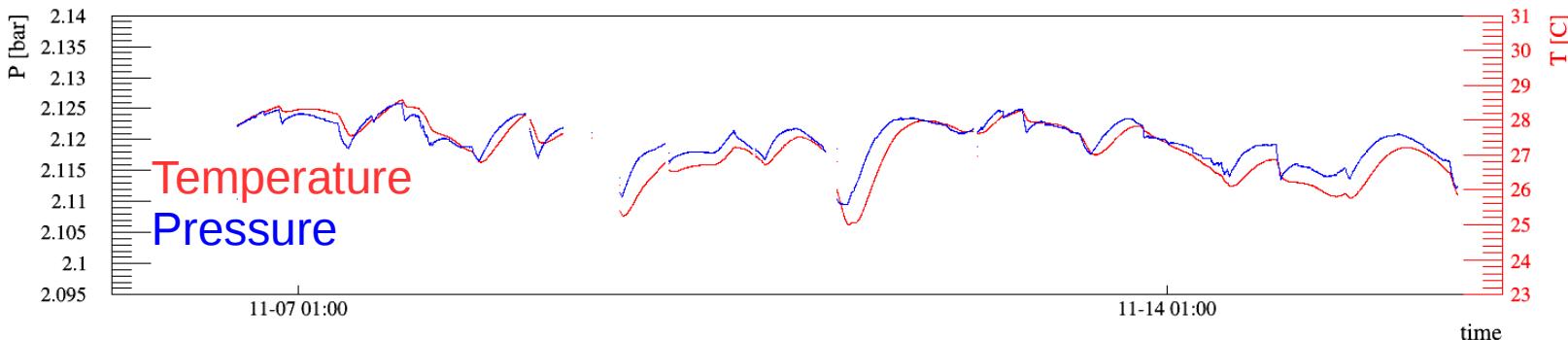
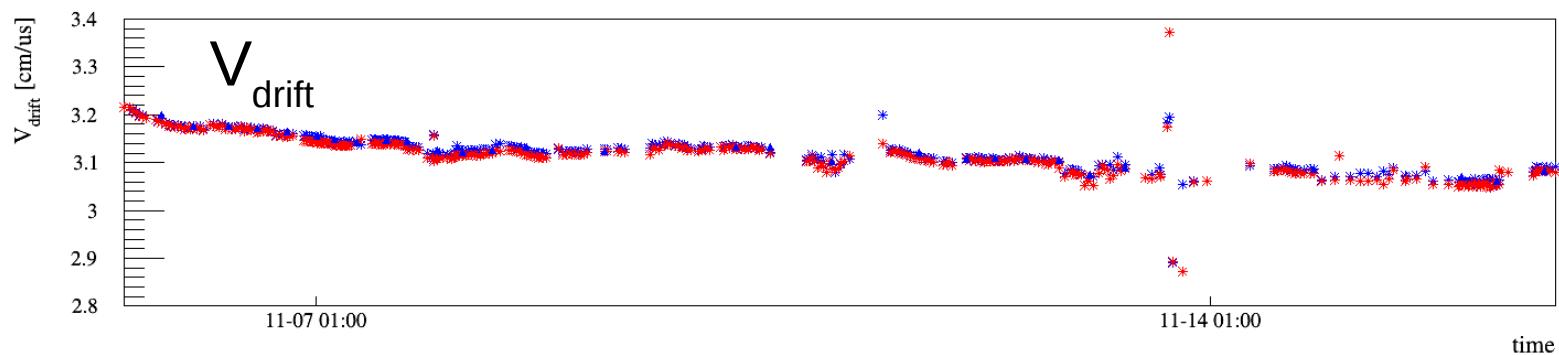
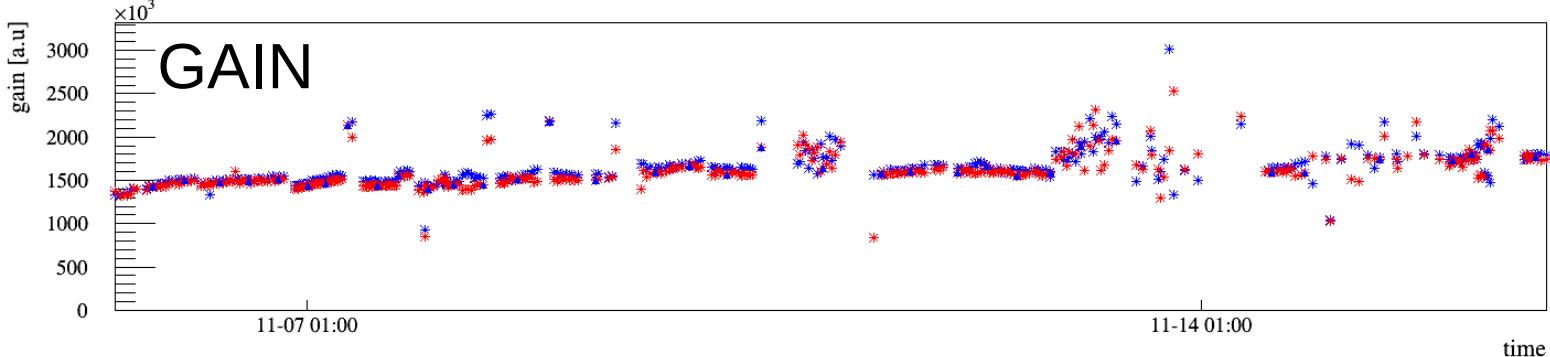




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# Time evolution

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# Gas stability

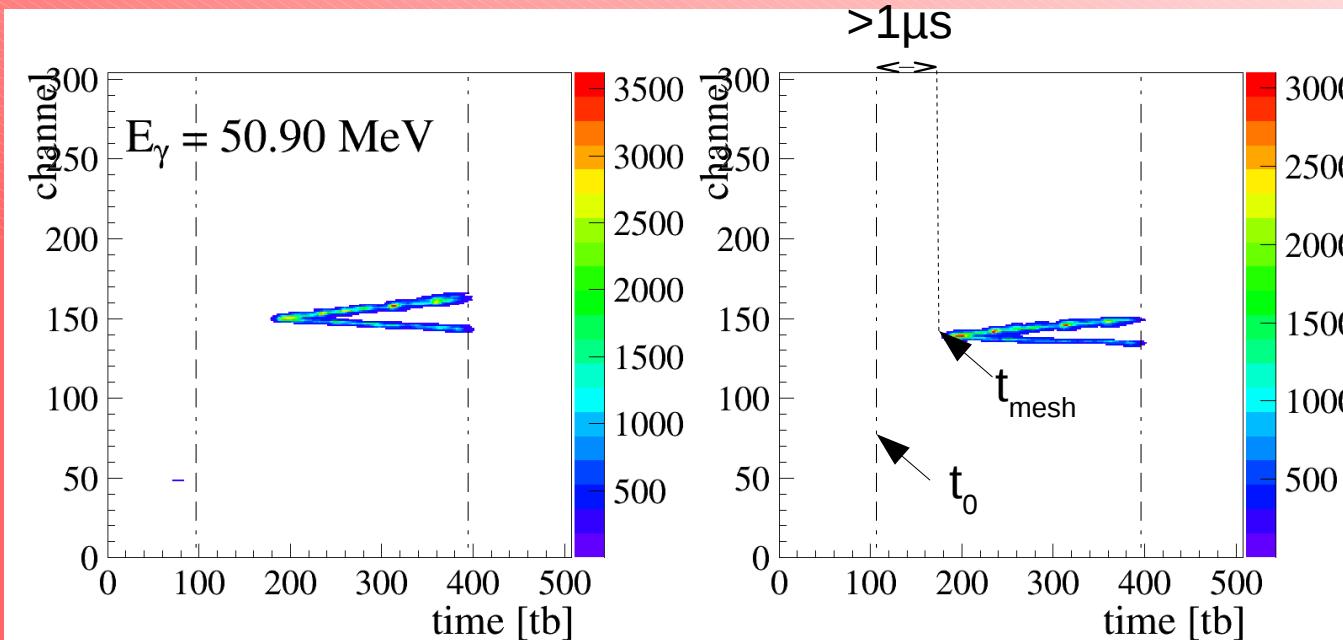
- 23 days in beam with same gas
- Almost no leak (no pressure loss)
- Slight decrease of Vdrift
- Stable gain
- Not much contamination (absorption). Longer analysis (with more statistics) necessary.



# Gamma trigger

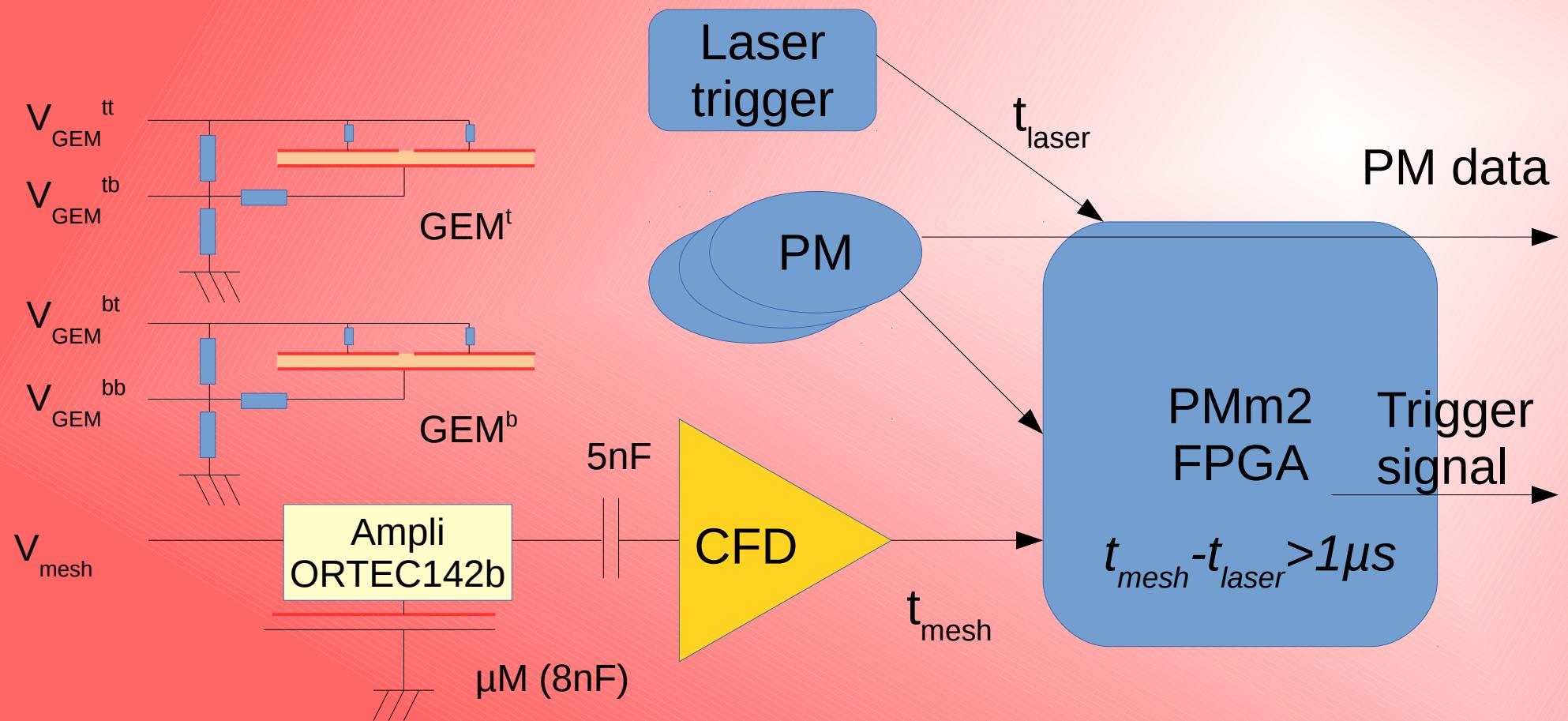
- Trigger on conversion in the gas volume

- at least one scintillator ( $\Rightarrow$  timing + trigger information)
- if available, laser trigger signal
- mesh signal  $>1\mu\text{s}$  ( $\rightarrow 3\text{cm}$ ) later



# Micromegas trigger signal

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# Trigger performance (?)

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#events

10<sup>3</sup>

10<sup>2</sup>

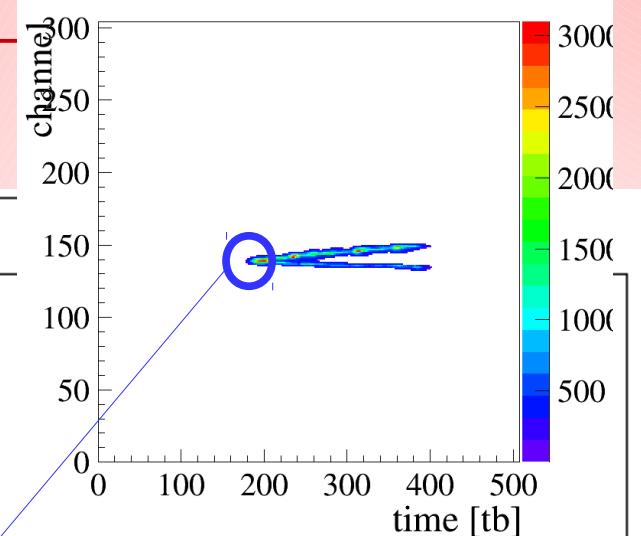
10<sup>1</sup>

0

fTmin

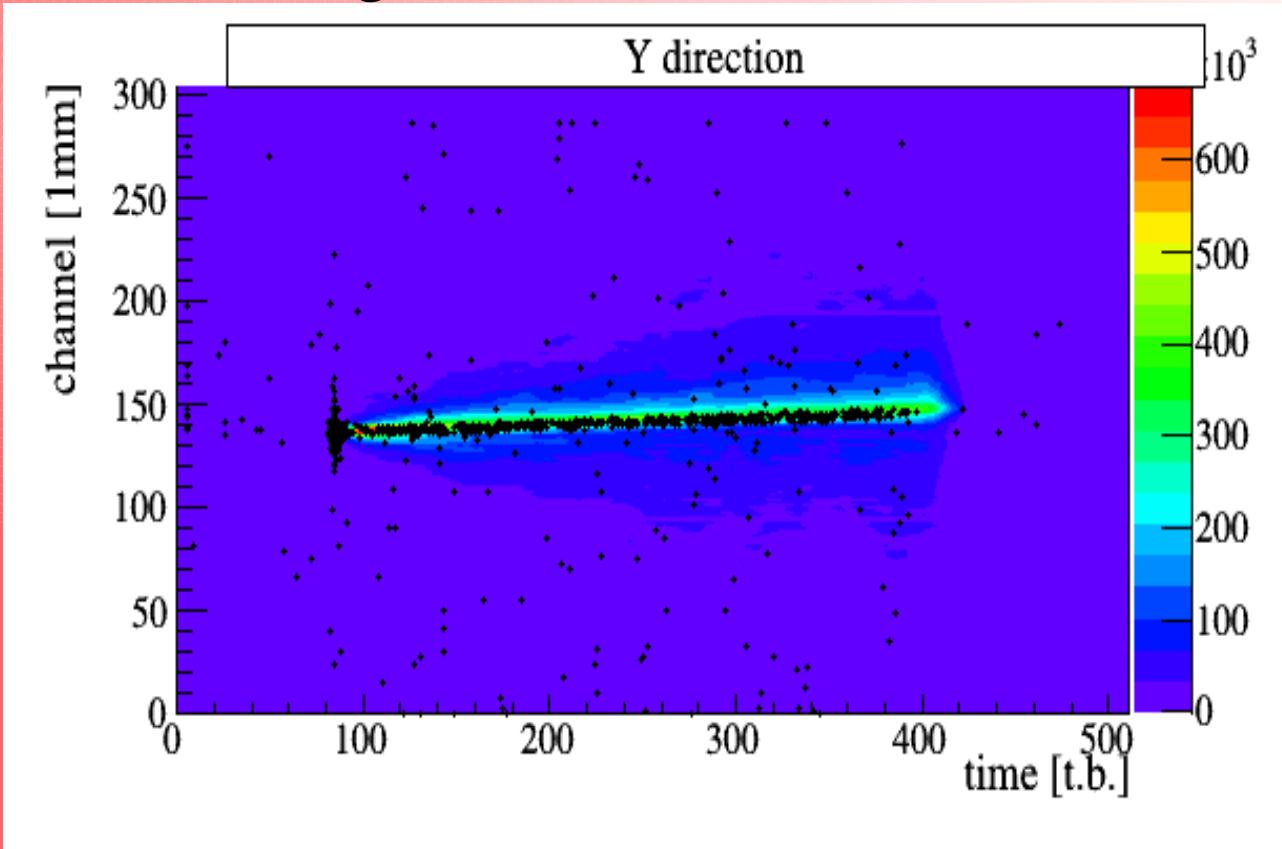
Conversion outside  
of the gas volume

T<sub>min</sub> [30ns bins]





- Further analysis of course needed
- “By eye” it worked well
  - Most events originate on the beam

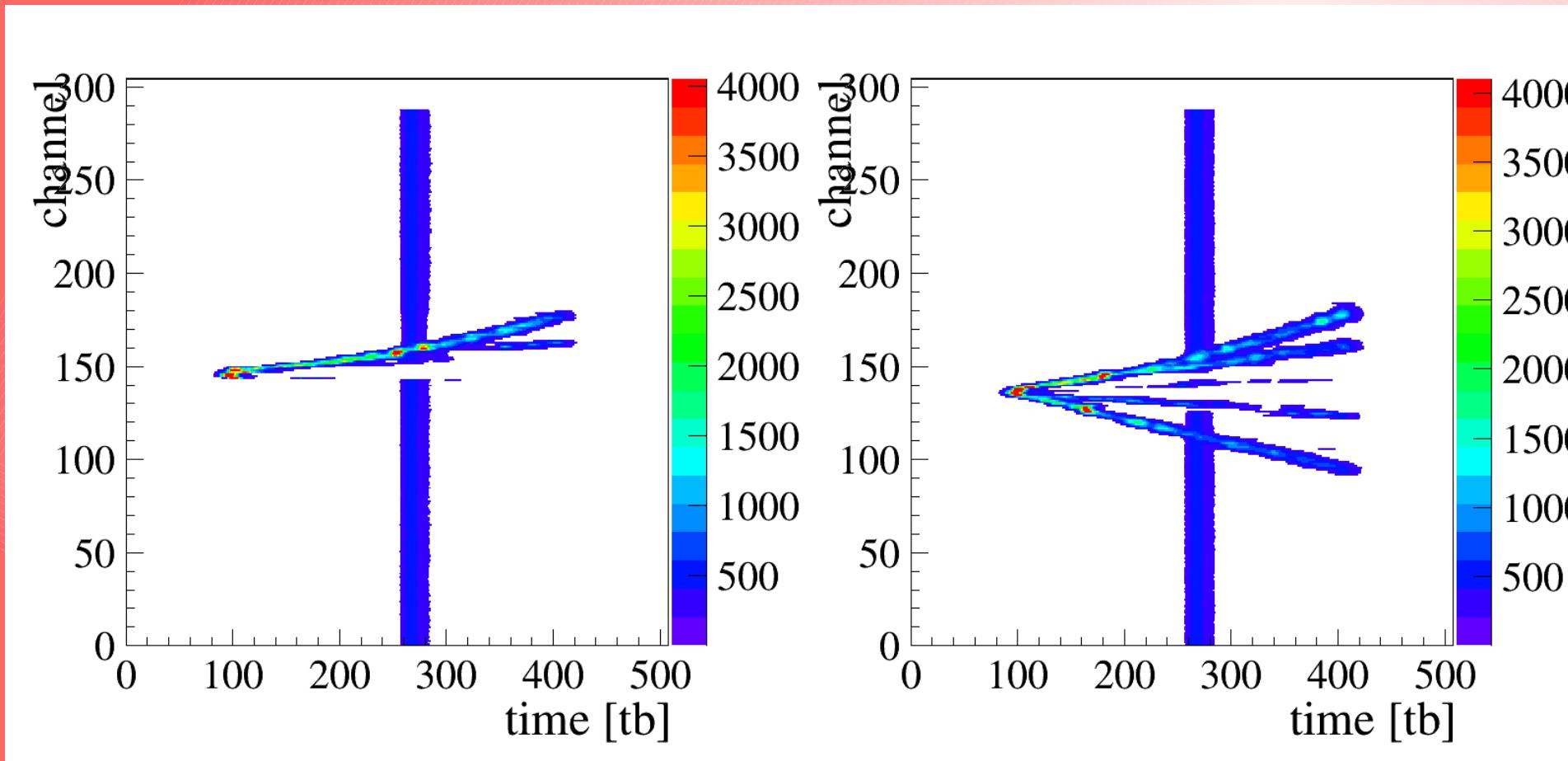




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# A puzzling effect

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# Space charge? Ballistic deficit?



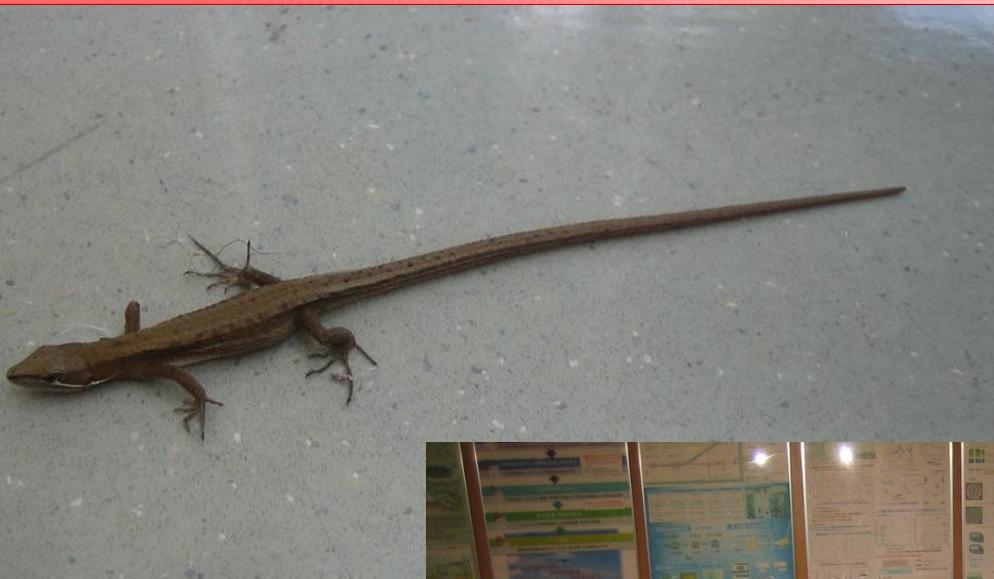
- Observed signal loss at the center of the detector (beam axis)
  - Only at high rate
  - space charge? ballistic deficit? other?
- “Ballistic deficit”
  - shaping 116ns
  - tracks in Z direction => several  $\mu s$  signals
- Improved by misaligning the detector



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# Intermède animalier 3

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# Conclusion and outlook



- Successful beam campaign with gamma rays
- Good gas stability over more than 20 days
- Good trigger performance
- Most issues related to high rate in beam
- Lots of work necessary for quantitative results...

# Merci!

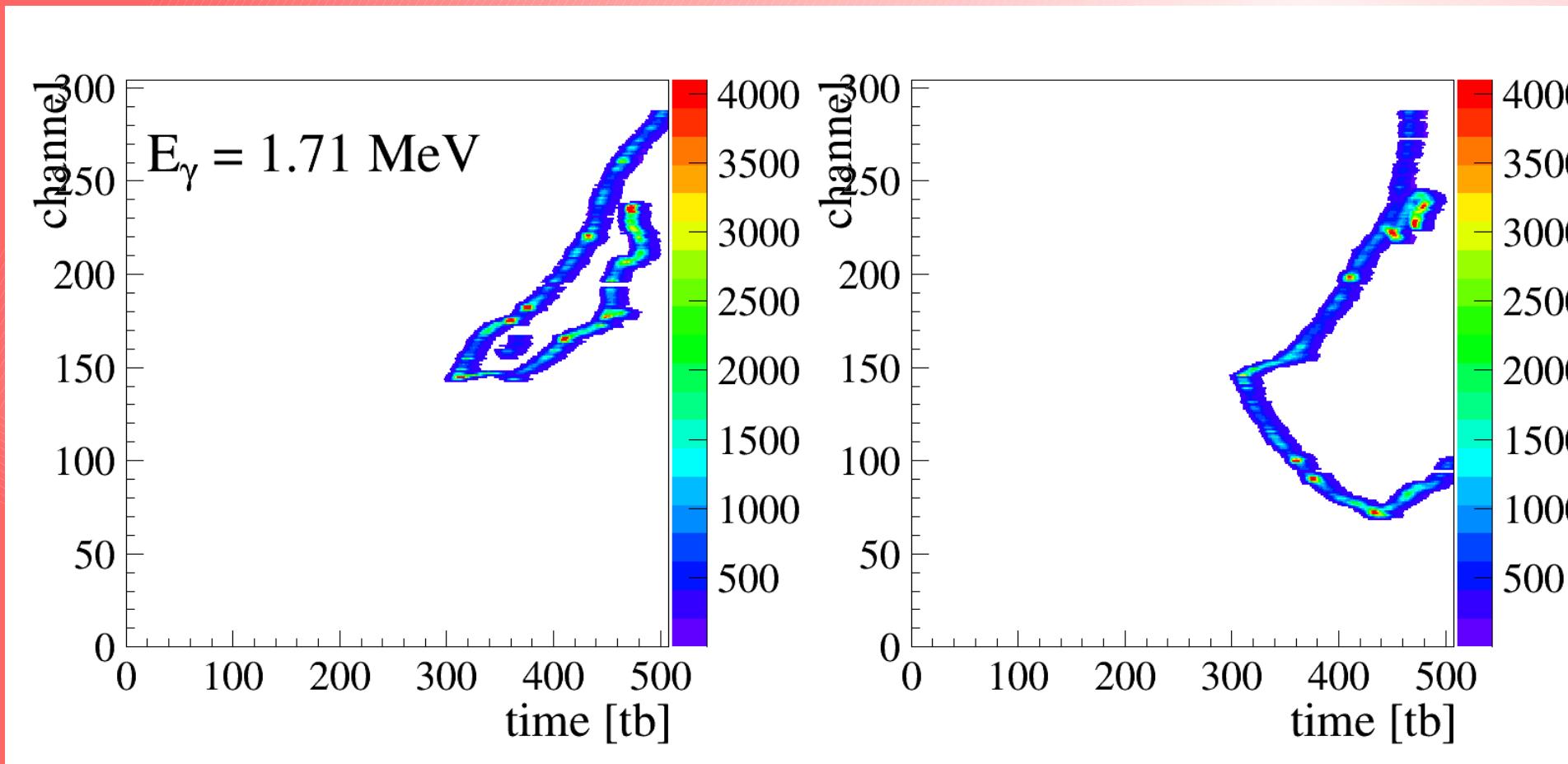


- Service de mécanique
  - conception, fabrication
- Service électronique
  -
- Service informatique
  -
- Service administratif
  - mission Japon, MoU, ...



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# Backup

# Pressure scan

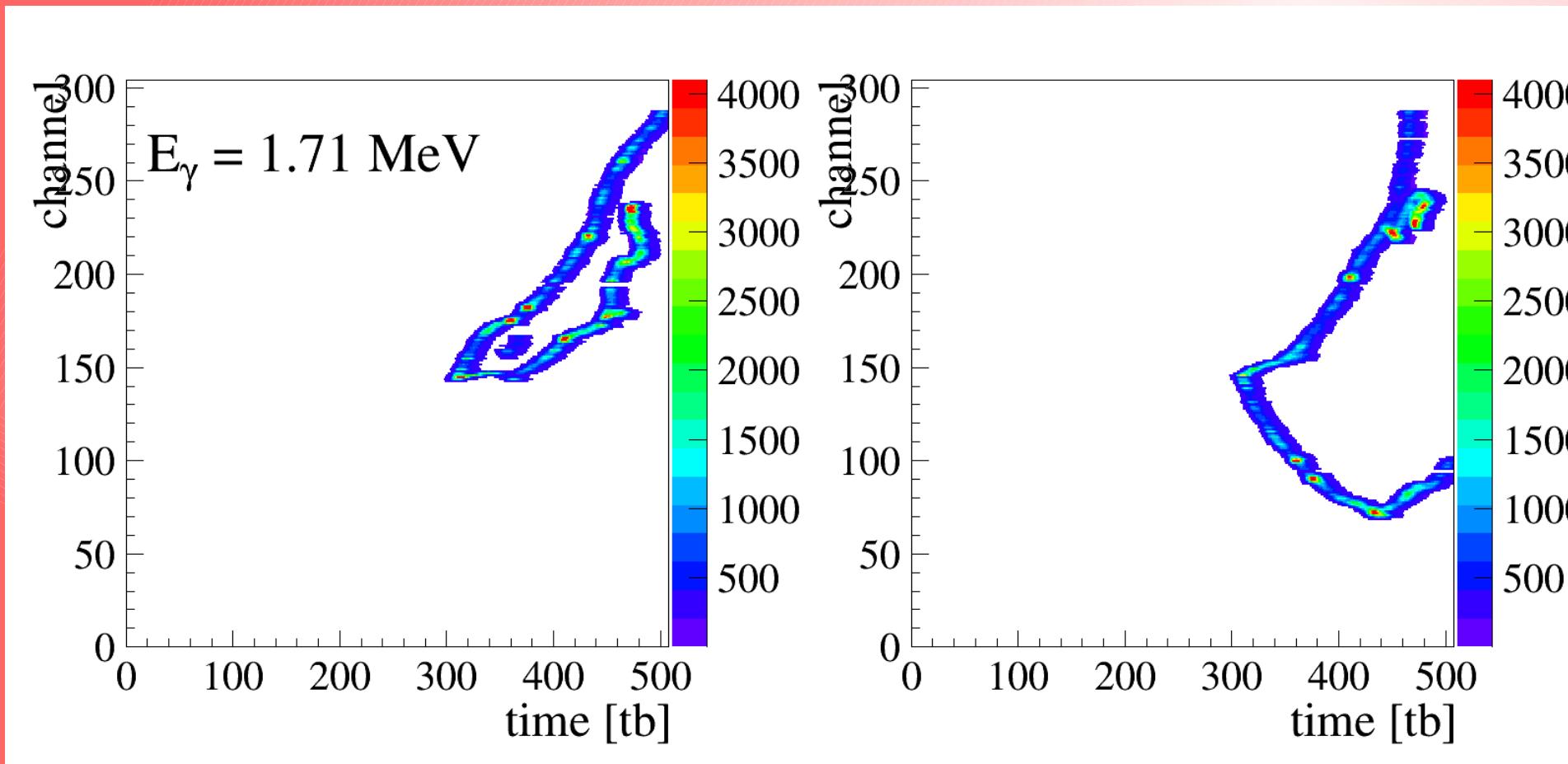


- One day dedicated to pressure scan
  - “Clean” gas at 1, 1.5, 2, 3 and 4 bar
  - Signal amplitude ( $dE/dx \times \text{gain}$ )  $\sim$ fixed  
(Adjusted on the fly)
  - At high pressure, cathode voltage was limited
- Good running condition at all pressures
- Increasing micromegas current spikes at high pressure
  - matching GEM over current => physics?



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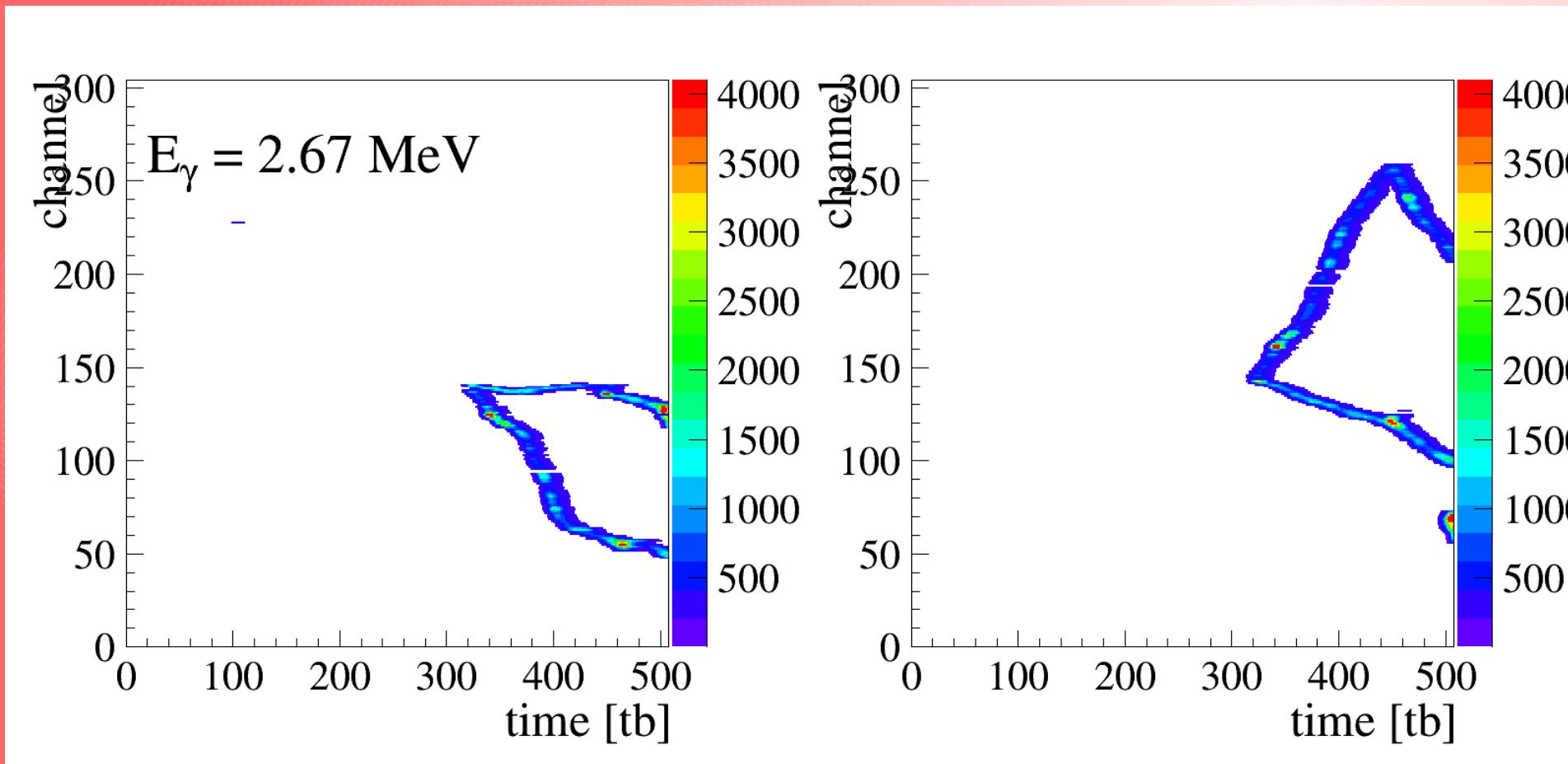
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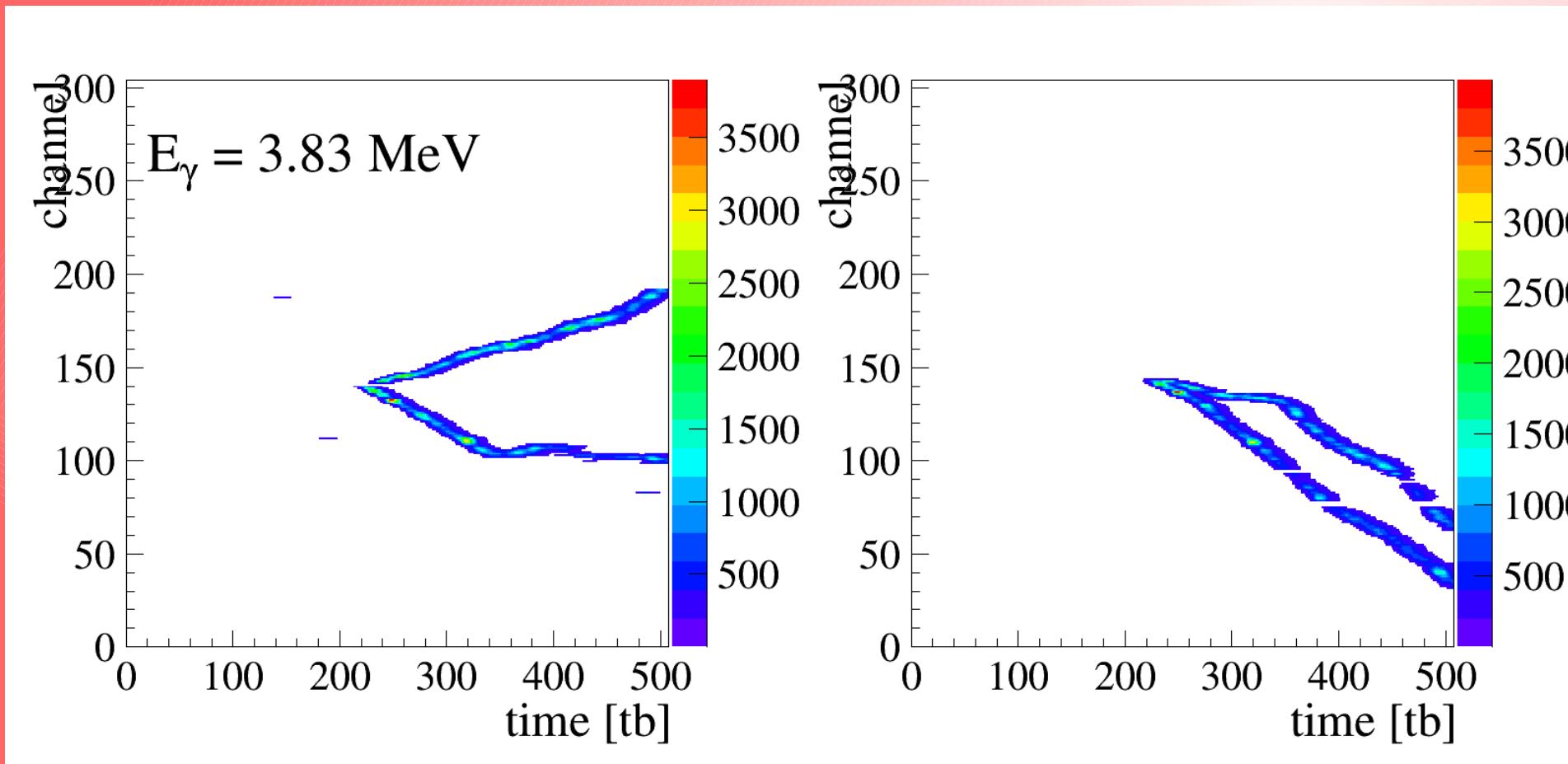
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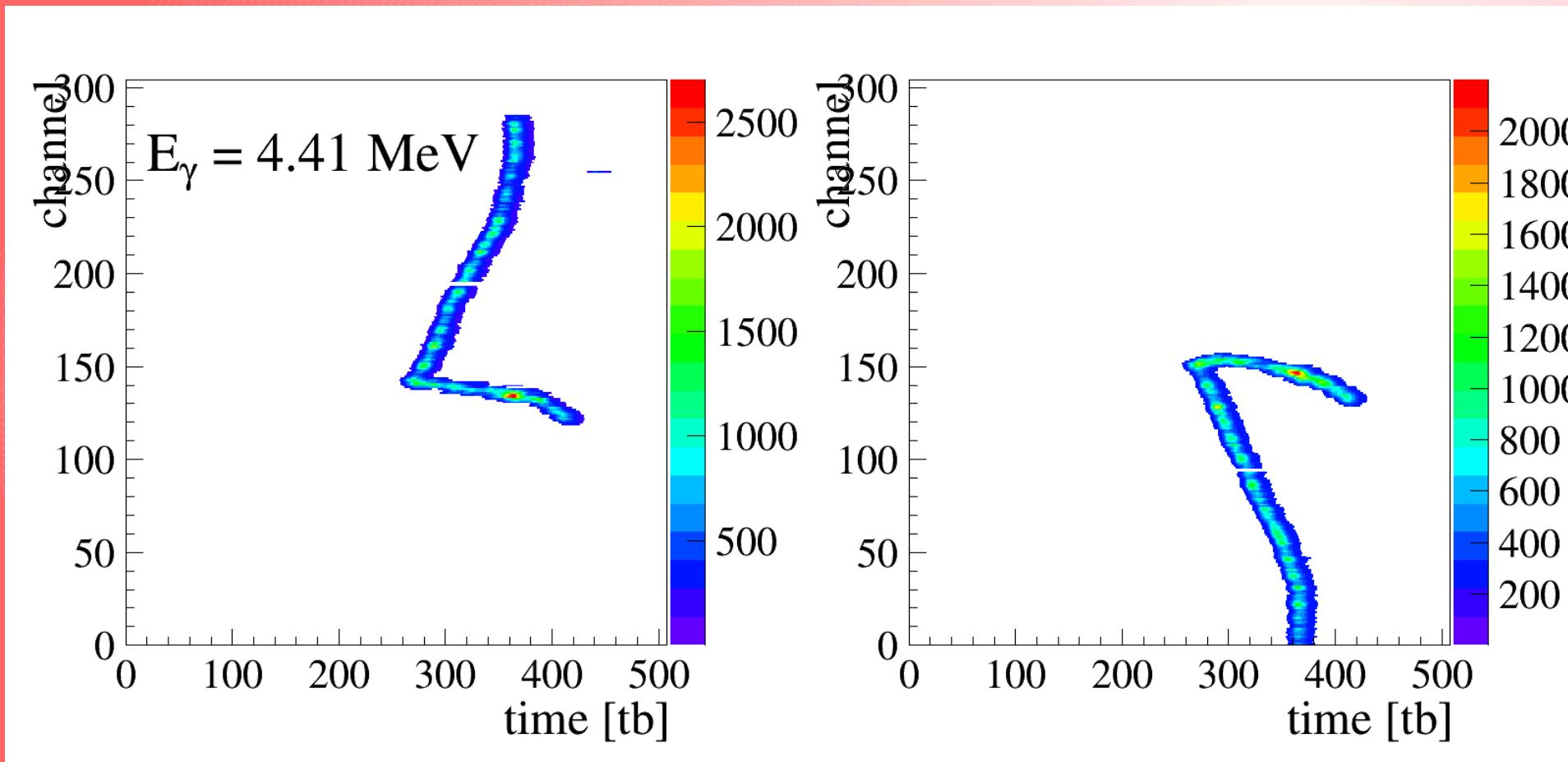
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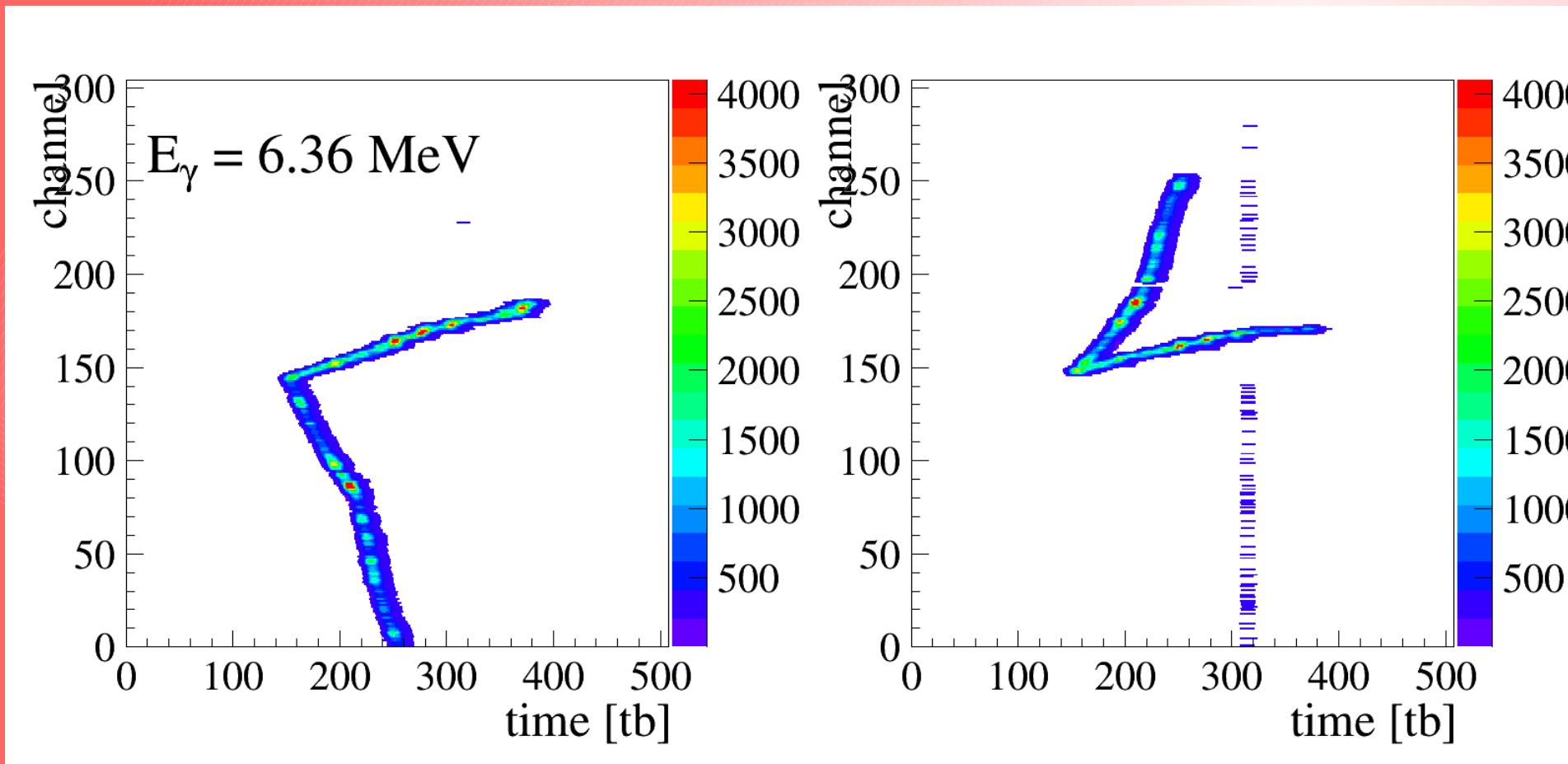
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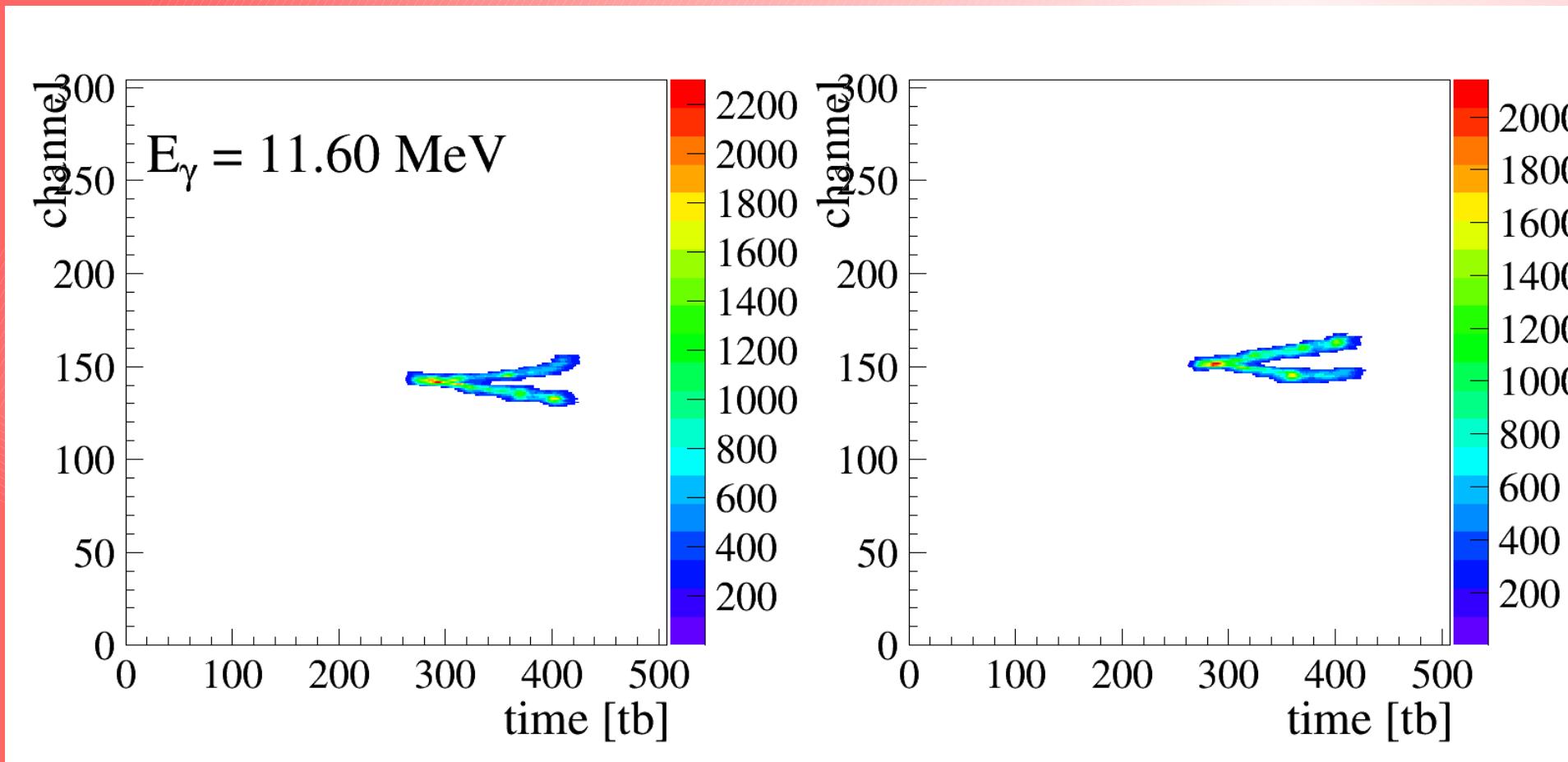
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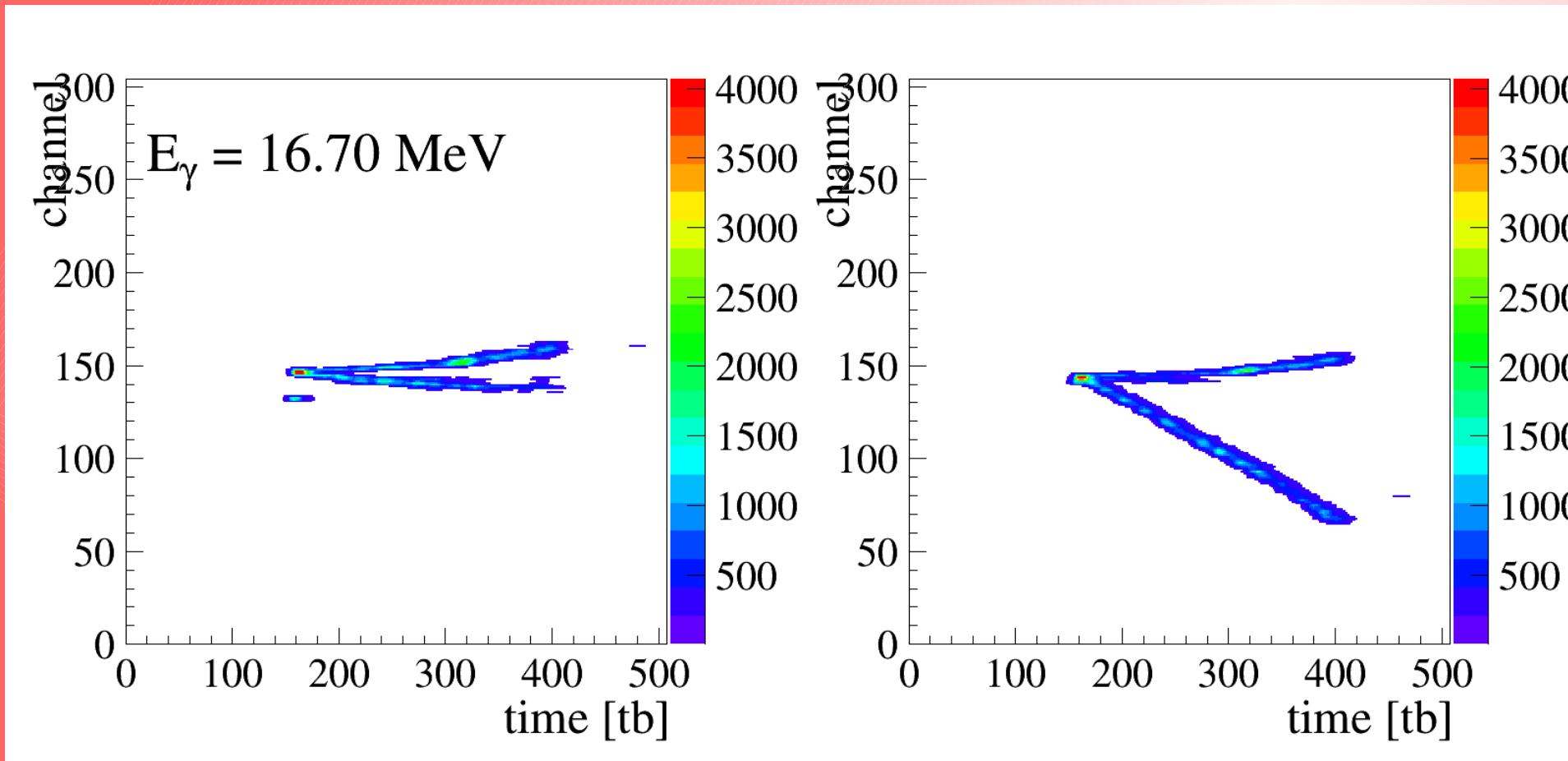
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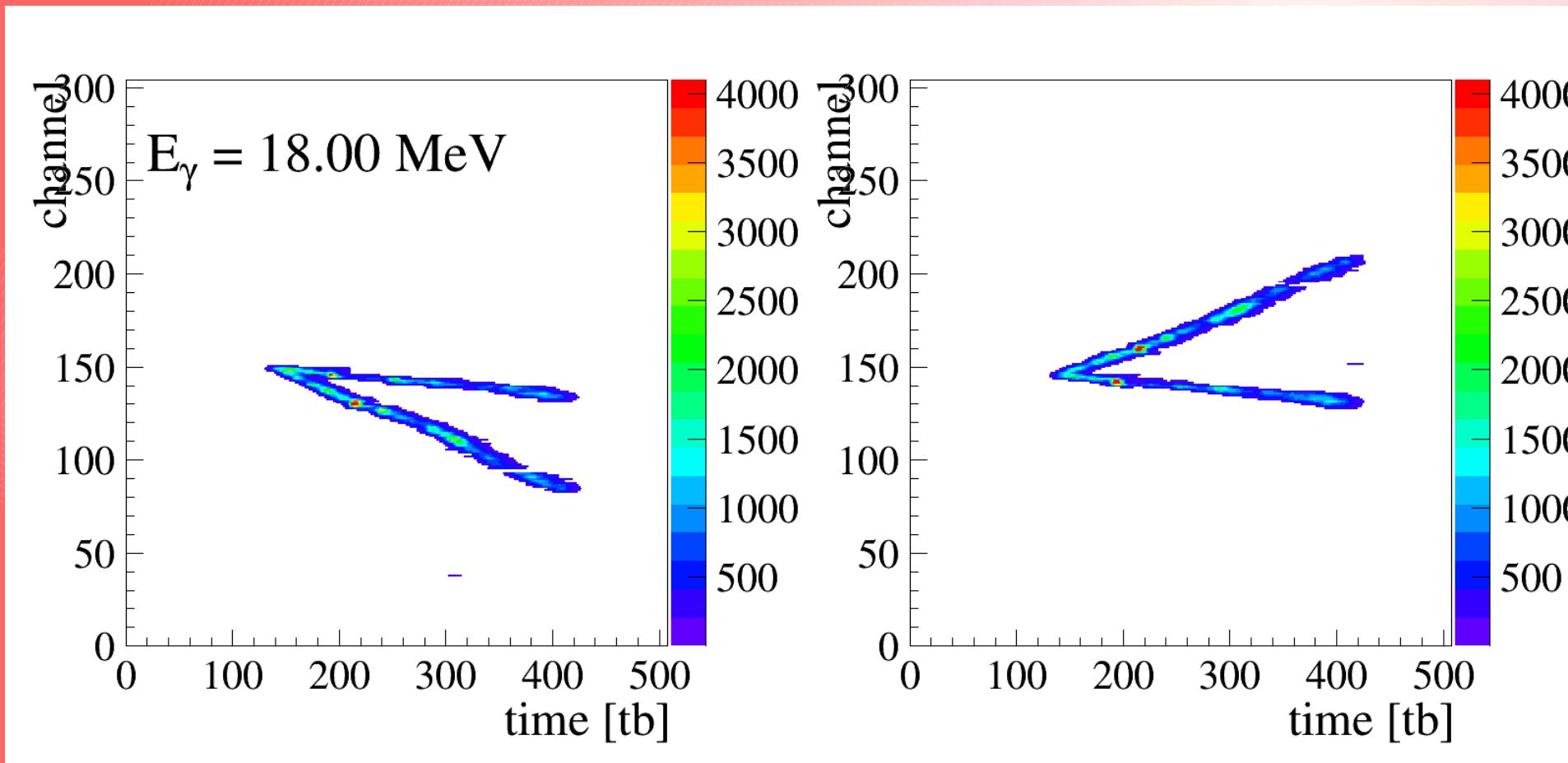
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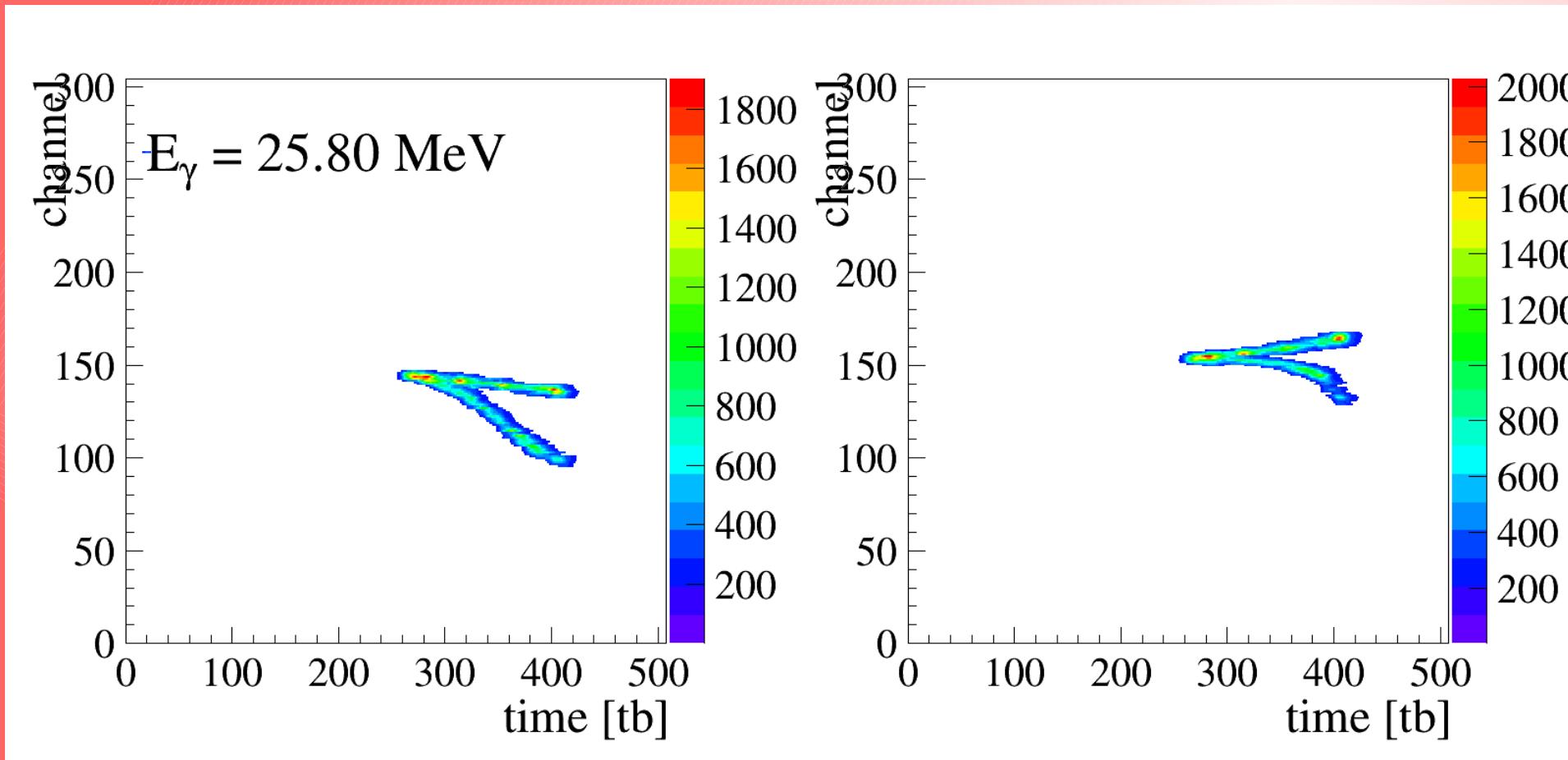
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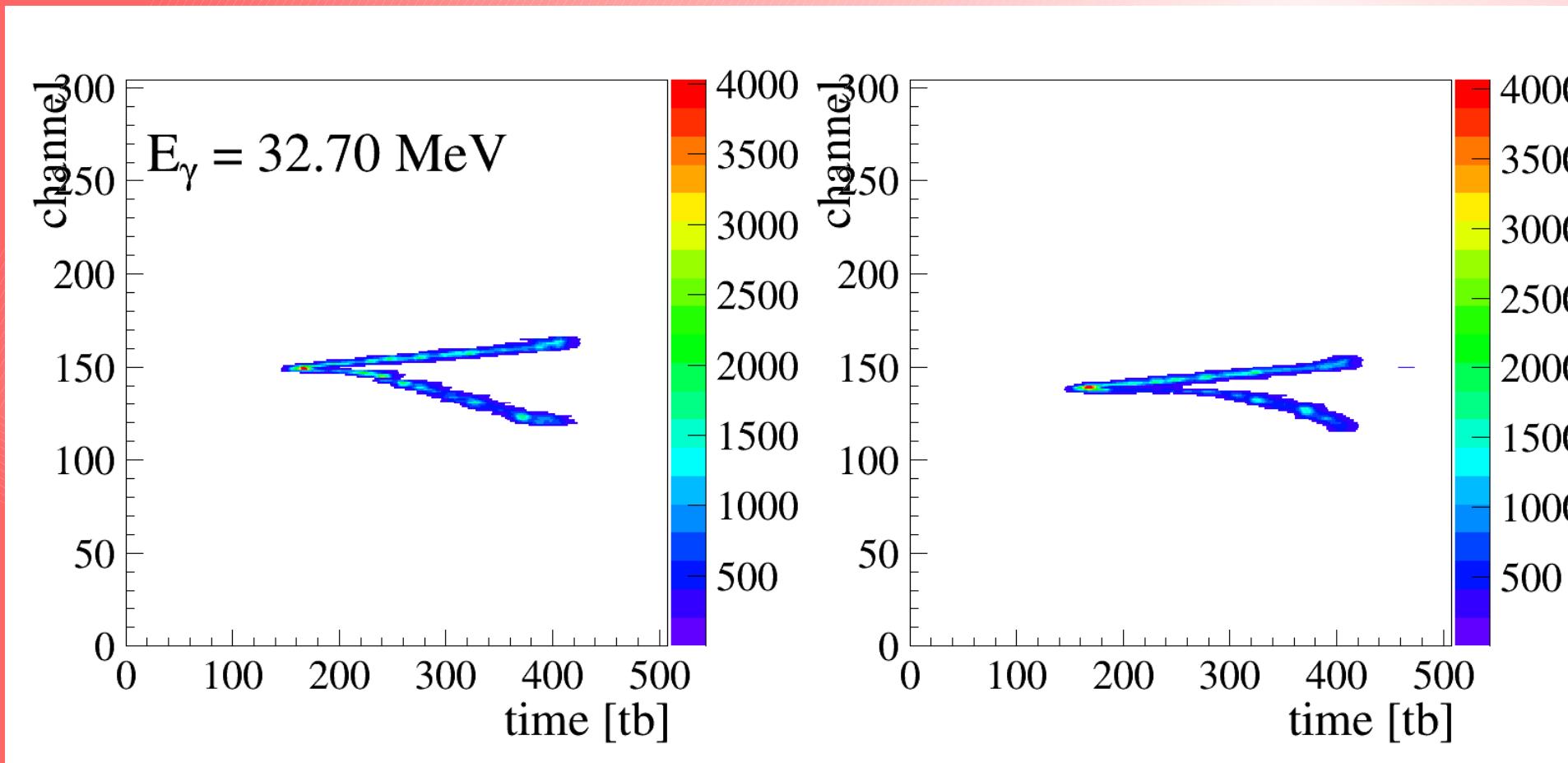
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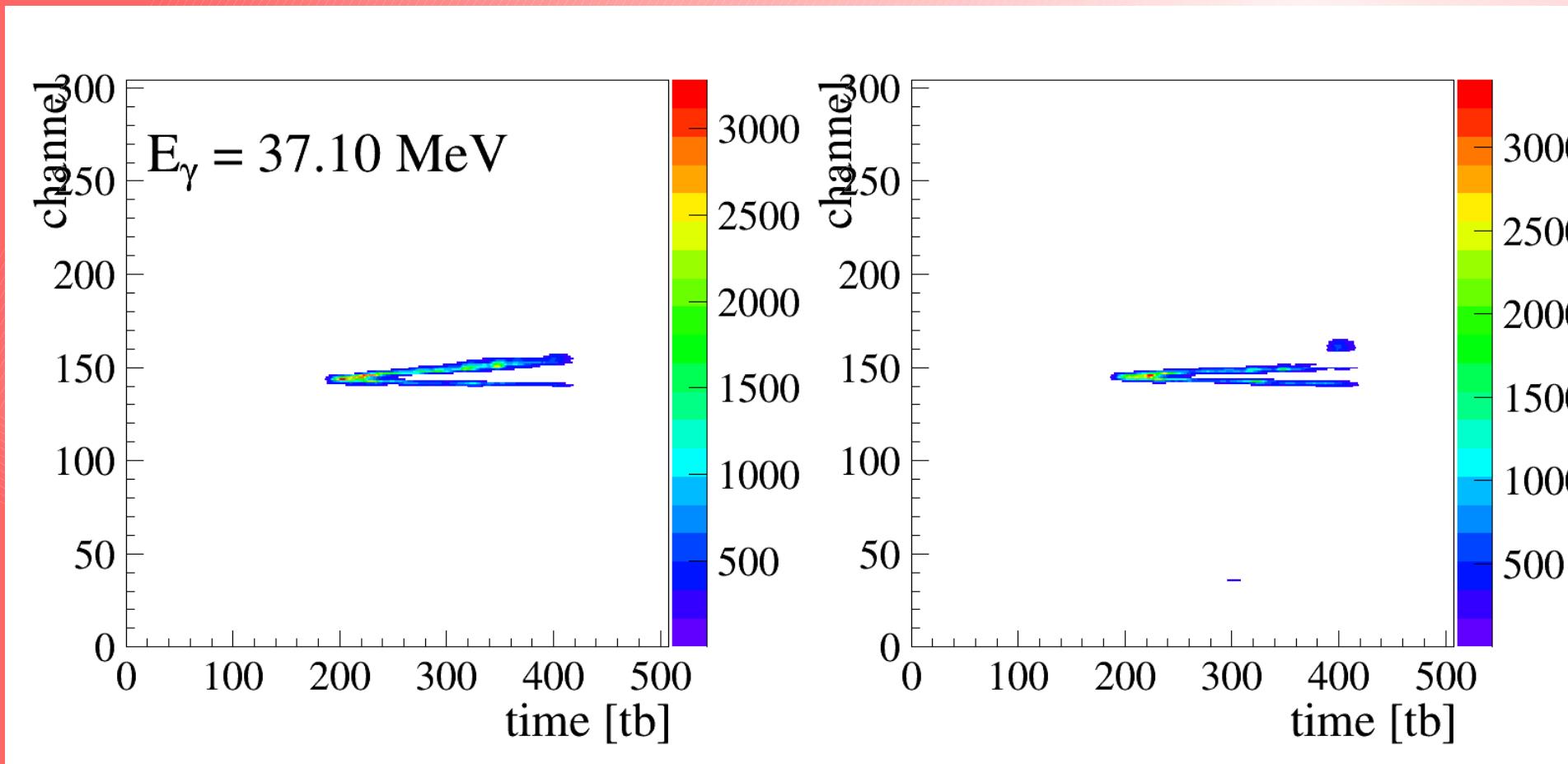
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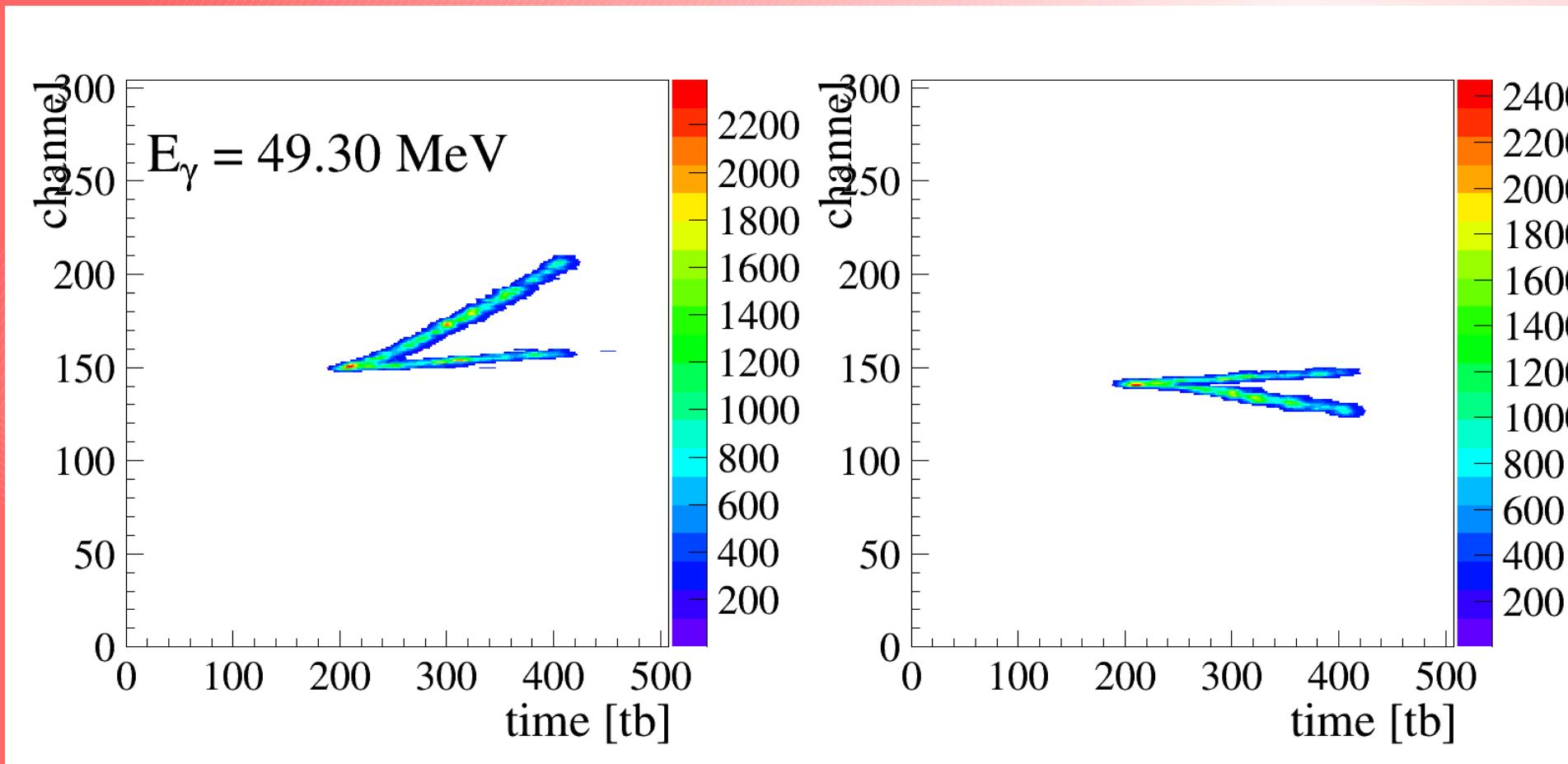
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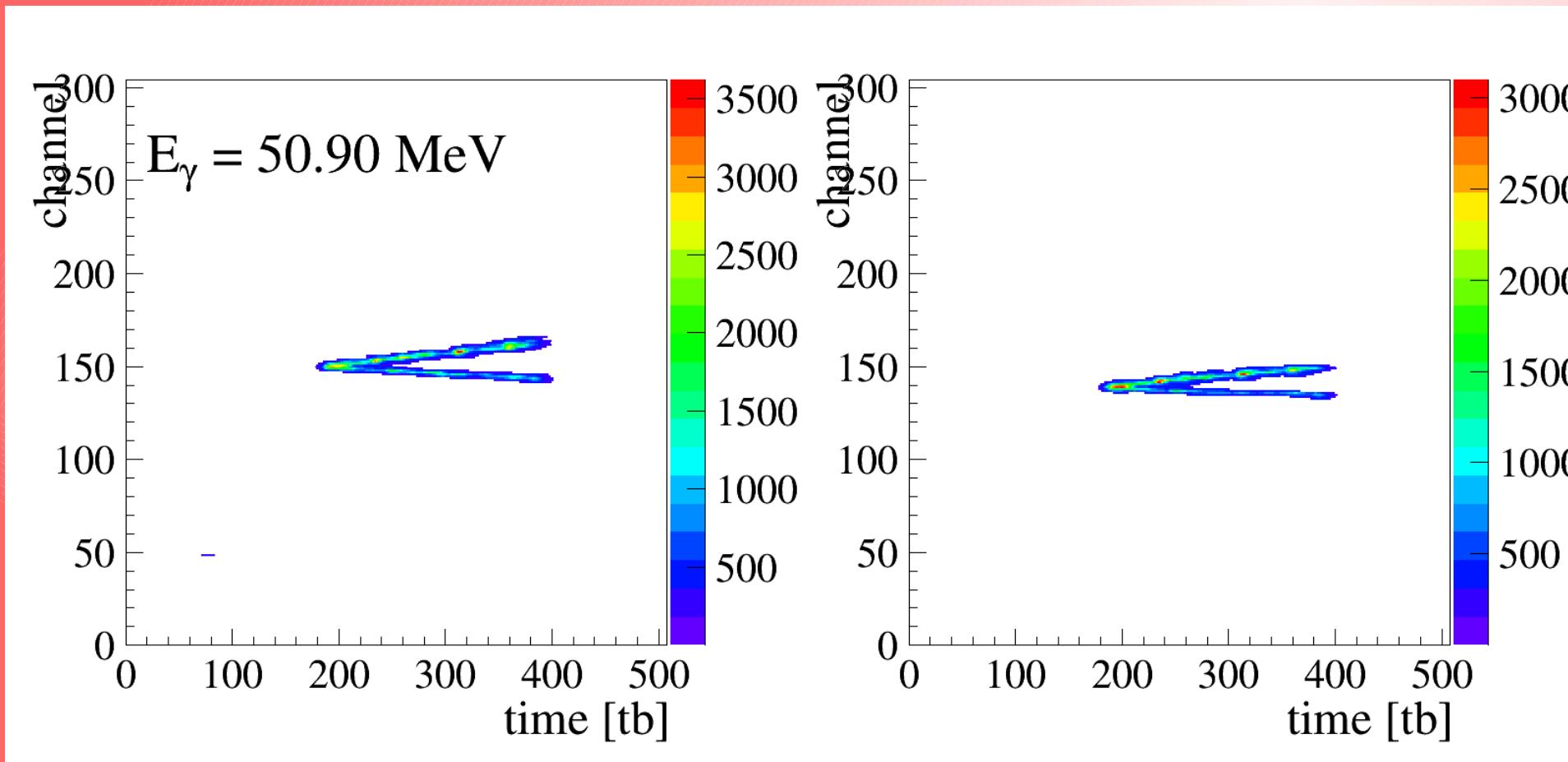
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